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Statistics of Some Atmospheric Turbulence Records Relevant to Aircraft Response Calculations

William D. Mark and Raymond W. Fischer

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Statistics of Some Atmospheric Turbulence Records Relevant to Aircraft Response Calculations

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Prepared for
Langley Research Center
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and Space Administration

**Scientific and Technical
Information Branch**

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SUMMARY

This report illustrates the results of application to turbulence velocity records of several new methods for characterizing atmospheric turbulence that are described in Ref. 1. The methods illustrated include maximum likelihood estimation of the integral scale and intensity of records obeying the von Karman transverse power spectral form, constrained least-squares estimation of the parameters of a parametric representation of autocorrelation functions, estimation of the power spectral density of the instantaneous variance of a record with temporally fluctuating variance, and estimation of the probability density functions of various turbulence components. The report also contains descriptions of the computer programs used in the computations, and a full listing of these programs. The computational methods illustrated herein were developed by the first named author. The computer programs and their explanation contained in the Appendices were written and exercised by the second named author.

TURBULENCE MODEL

In the work described in this report, we shall assume that the turbulence velocity records under consideration can be modeled as

$$\begin{aligned}w(t) &= w_s(t) + w_f(t) \\ &= w_s(t) + \sigma_f(t)z(t),\end{aligned}\tag{1.1}$$

where

$$w_f(t) = \sigma_f(t)z(t)\tag{1.2}$$

with

$$\sigma_f(t) \geq 0$$

and

$$E\{z(t)\} = 0, \quad E\{z^2(t)\} = 1.\tag{1.3}$$

The three processes $\{w_s(t)\}$, $\{\sigma_f(t)\}$, and $\{z(t)\}$ are assumed to be stationary and mutually statistically independent. Furthermore, we shall assume that $\{z(t)\}$ is a Gaussian process. The "slow" turbulence component $w_s(t)$ is assumed to contain predominately very low frequencies (or large wave-numbers) relative to the "fast" component $w_f(t)$ which may be regarded as ordinary turbulence with a slowly varying standard deviation $\sigma_f(t)$. This model is more completely described in Section 1 of the companion report [1] or Section 2 of its predecessor [5].

MAXIMUM LIKELIHOOD ESTIMATION OF THE INTEGRAL SCALE AND
INTENSITY OF THE VERTICAL RECORD FROM FLIGHT 8 RUN 2
(CONVECTIVE CONDITIONS)

The vertical record shown in Fig. 1 illustrates a turbulence velocity history with negligible low-frequency component $w_s(t)$. The power spectral density of the vertical record in Fig. 1 is shown in Fig. 2. The method used to compute the power spectral density of Fig. 2 is described in Appendix B of Ref. 2, where the value used for M was 6590.5 m which corresponds to 1024 temporal sample points. Before computing the power spectral density of the record, its mean value was computed and removed.

Also shown in Fig. 2 is the von Karman transverse power spectral density

$$\Phi_{KT}(k) = \sigma^2 L \frac{1+188.75L^2k^2}{[1+70.78L^2k^2]^{1\frac{1}{6}}} \quad (2.1)$$

The values of L and σ^2 in Eq. (2.1) — as plotted in Fig. 2 — are

$$L = 309.4 \text{ m}, \quad \sigma^2 = 1.326 \text{ (m/sec)}^2. \quad (2.2)$$

These values were computed using the maximum likelihood method derived in Sec. 3 of Ref. 1. The specific equation used to compute the value of L in Eq. (2.2) was Eq. (3.26) of Ref. 1 with the aid of Eqs. (3.34) and (3.35) of Ref. 1. Details of this computation are described in Appendix F of Ref. 1. Using the value of L obtained by Eq. (3.26) of Ref. 1, Eq. (3.25) of Ref. 1 was then used to compute the values of σ^2 given in Eq. (2.2) above.

The von Karman transverse spectrum shown in Fig. 2 provides an excellent fit to the spectrum computed from the turbulence record. In particular, note that asymptotic (high wavenumber) slopes of the empirical and von Karman spectra agree very well. We also have computed the value of σ^2 directly — by squaring and averaging the time history sample points. The value of σ^2 obtained in this manner was

$$\sigma^2 = 1.331 \text{ (m/sec)}^2. \quad (2.3)$$

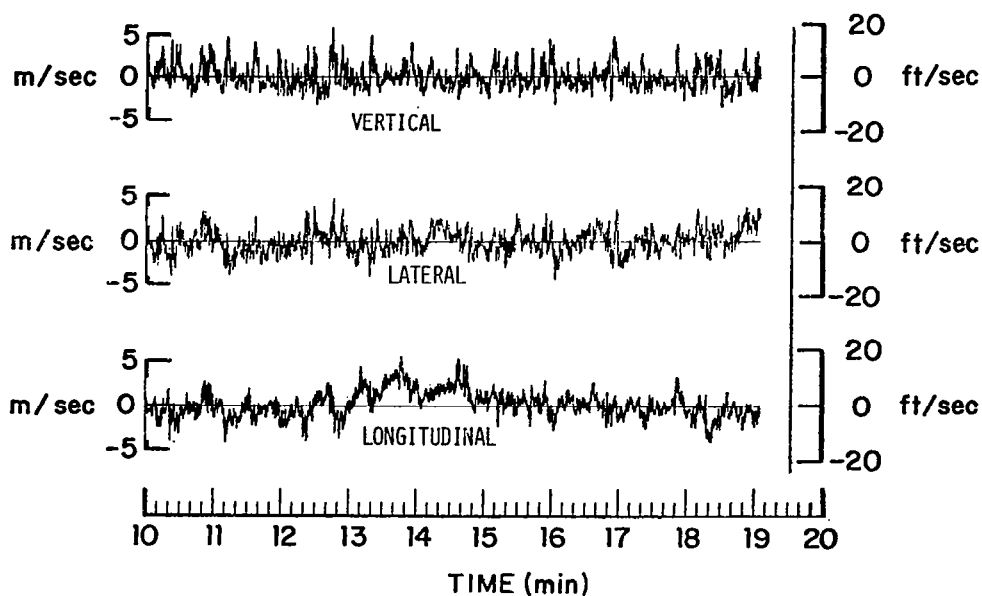
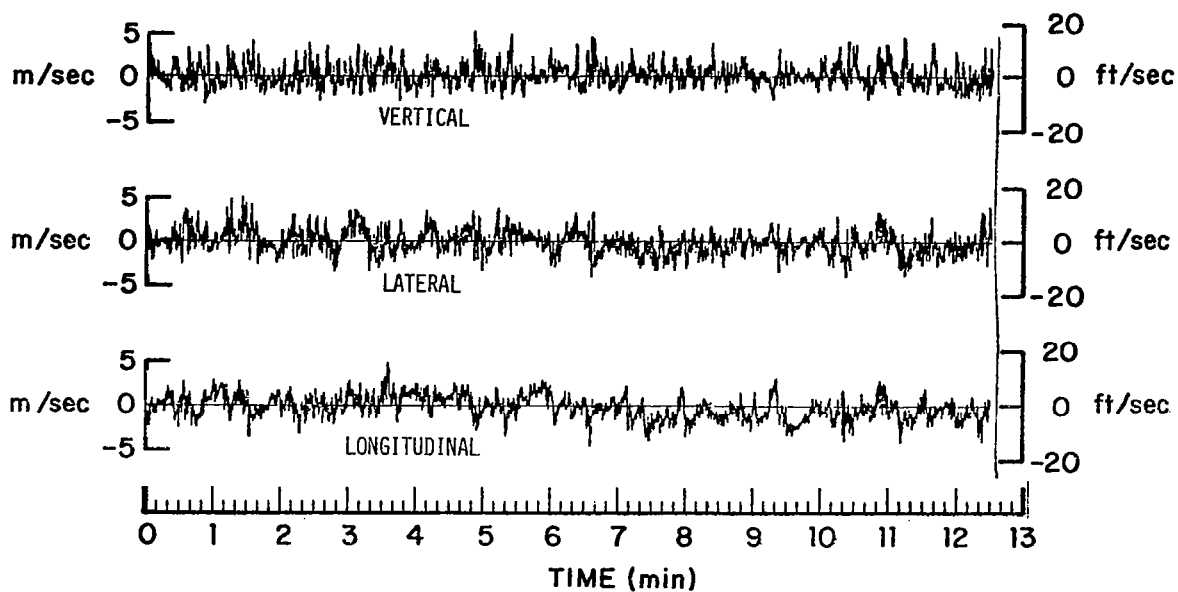


FIG. 1. LOW-ALTITUDE TURBULENCE RECORDS. [CONVECTIVE CONDITIONS. AIRCRAFT SPEED 129 m/sec (422 ft/sec).] (Ref. 3, Fig. 4, p. 282).

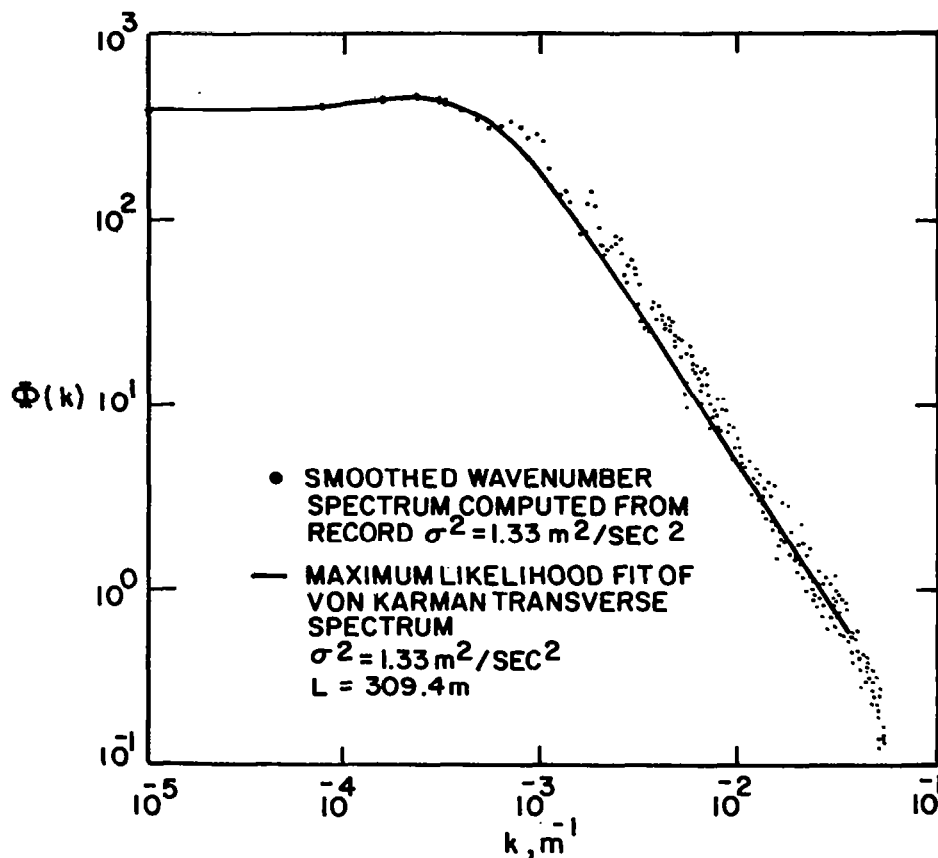


FIG. 2. COMPARISON OF SMOOTHED WAVENUMBER SPECTRUM COMPUTED FROM VERTICAL RECORD SHOWN IN FIG. 1 AND MAXIMUM LIKELIHOOD FIT OF von KARMAN TRANSVERSE SPECTRUM.

The maximum likelihood method used to compute the value of σ^2 given by Eq. (2.2) is *not* the same as the squaring and averaging procedure used to compute the value of Eq. (2.3). However, quite remarkably, the two values agree to the first three significant figures. Since the assumption of a von Karman transverse spectrum was used in the computation of σ^2 given by Eq. (2.2), the close agreement of the values of σ^2 given by Eqs. (2.2) and (2.3) provides verification of the excellent representation of the empirical spectrum that is provided by the von Karman transverse spectrum of Eq. (2.1).

The autocorrelation function of the vertical record shown in Fig. 1 is compared in Fig. 3 with the von Karman transverse autocorrelation function:

$$\begin{aligned} \phi_{KT}(\xi) \triangleq \sigma^2 \frac{2^{2/3}}{\Gamma(1/3)} (\beta\xi/L)^{1/3} [K_{1/3}(\beta\xi/L) \\ - \frac{\beta\xi}{2L} K_{-2/3}(\beta\xi/L)], \end{aligned} \quad (2.4)$$

where

$$\beta \triangleq \frac{2\sqrt{\pi}}{5} \frac{\Gamma(11/6)}{\Gamma(4/3)}, \quad (2.5)$$

where the $K_n(\cdot)$ in Eq. (2.4) are modified Bessel functions of the second kind of order n and $\Gamma(\cdot)$ is the gamma function. Values of the Bessel functions in Eq. (2.4) were obtained from the tabulation of p. 228 of Ref. 4, where we note that $K_{-2/3}(x) = K_{2/3}(x)$. The empirical autocorrelation function in

Fig. 3 was computed from the vertical record of Fig. 1 by the method described in Appendix B of Ref. 2. Both autocorrelation functions shown in Fig. 3 are normalized to unity at the origin. The value of integral scale L used in the von Karman form of Eq. (2.4) is that given by Eq. (2.2).

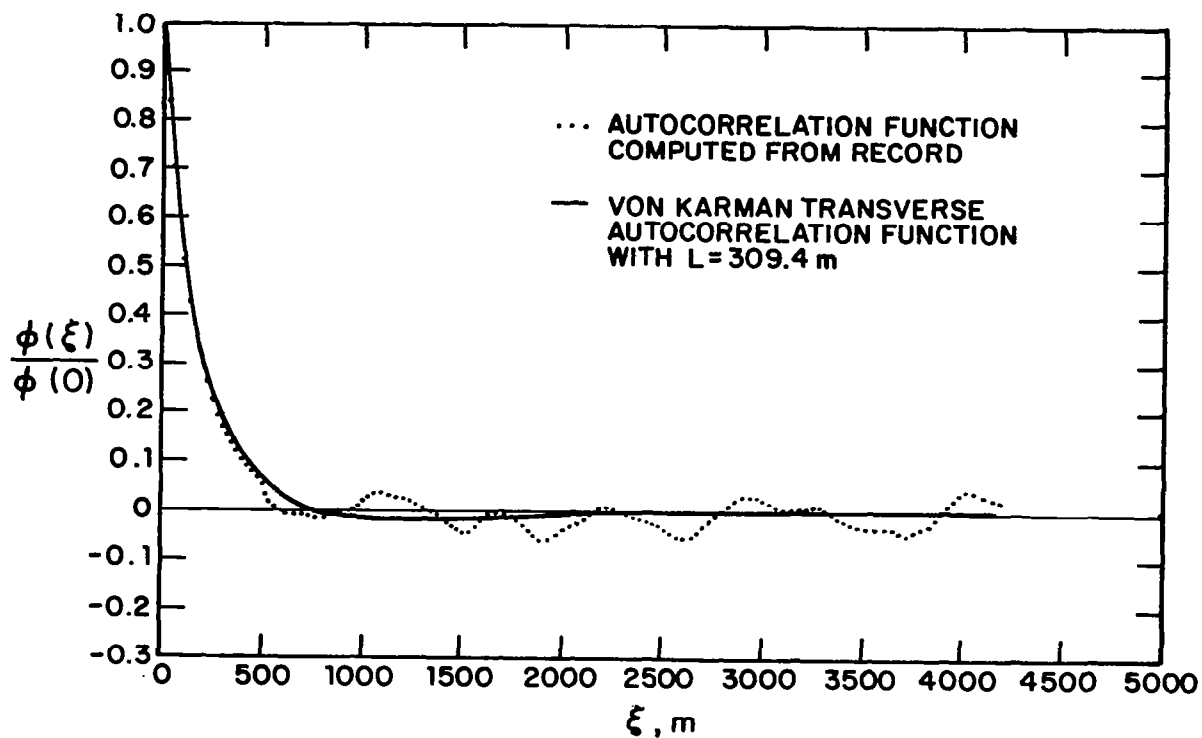


FIG. 3. COMPARISON OF AUTOCORRELATION FUNCTION COMPUTED FROM VERTICAL RECORD SHOWN IN FIG. 1 AND von KARMAN TRANSVERSE AUTOCORRELATION FUNCTION.

CONSTRAINED LEAST-SQUARES ESTIMATION OF AUTOCORRELATION FUNCTION PARAMETERS OF LATERAL RECORD FROM FLIGHT 32 RUN 4 (WIND-SHEAR CONDITIONS)

The lateral record shown in Fig. 4 illustrates a turbulence velocity history with a relatively strong low-frequency component $w_s(t)$. The power spectral density of the lateral record in Fig. 4 is shown in Fig. 5 (solid dots), which was computed by the method described in Appendix B of Ref. 2 — the value of M used in the computation was 9613.3 m which corresponds to 1024 temporal sample points. Before computing the power spectral density of the record, its mean value was computed and removed.

Also plotted in Fig. 5 is the von Karman transverse spectrum of Eq. (2.1) evaluated from the parameters

$$L = 265.5 \text{ m}, \quad \sigma^2 = 5.315 \text{ m}^2/\text{sec}^2. \quad (3.1)$$

These parameter values were arrived at using the constrained least-squares estimation method described in Sec. 4 of Ref. 1. This method postulates that within an interval $0 \leq \xi \leq \xi_H$, the autocorrelation function of a record is of the form

$$\phi(\xi) \triangleq \sigma_f^2 \phi_K(\xi; L) + \sum_{i=0}^m a_i \xi^i, \quad 0 \leq \xi \leq \xi_H, \quad (3.2)$$

where $\sigma_f^2 \phi_K(\xi; L)$ is the appropriate (transverse or longitudinal) von Karman autocorrelation function, and the m th degree polynomial in Eq. (3.2) represents the autocorrelation function of the "slow" turbulence component $w_s(t)$ within the interval $0 \leq \xi \leq \xi_H$. The least-squares estimation procedure constrains the relationship between σ_f^2 and L using the portion of the wavenumber spectrum of the record (in the "high-frequency" region) between two wavenumbers k_ℓ and k_u as described in Sec. 4 of Ref. 1, where in the present case, we used $k_\ell = 10^{-3} \text{ m}^{-1}$ and $k_u = 4 \times 10^{-2} \text{ m}^{-1}$. Equation (4.4) of Ref. 1 is the equation of constraint. The resulting relationship between σ_f^2 and L for the present example is plotted in Fig. 6.

Figure 7 displays the autocorrelation function of the lateral record shown in Fig. 4. To determine the general behavior of the constrained least-squares estimation method

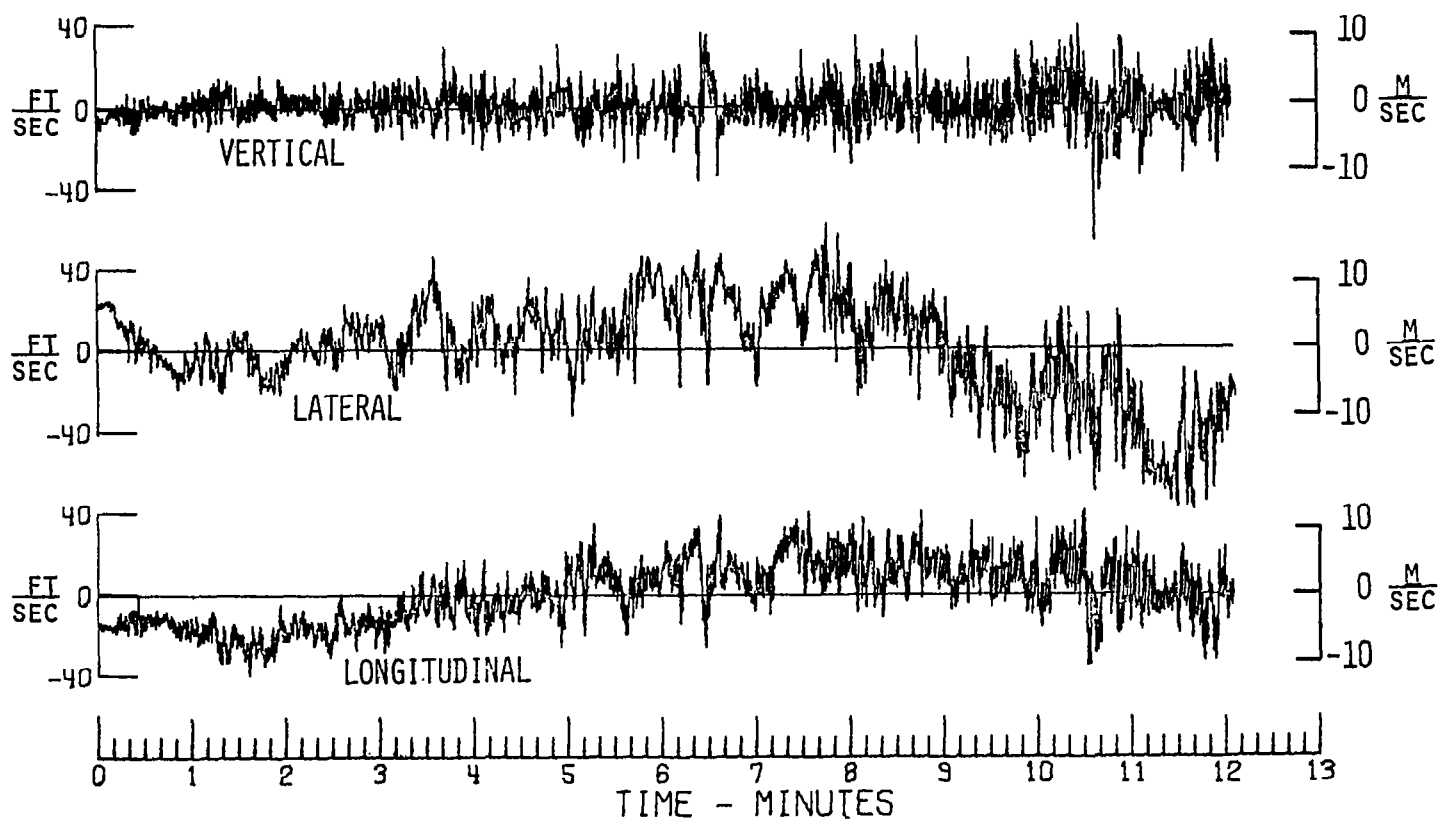


FIG. 4. TURBULENCE RECORDS CONTAINING STRONG "SLOW" COMPONENTS $w_s(t)$. [WIND SHEAR CONDITIONS. AIRCRAFT SPEED 188 m/sec (616 ft/sec).] Ref. 3, Fig. 6, p. 283.)

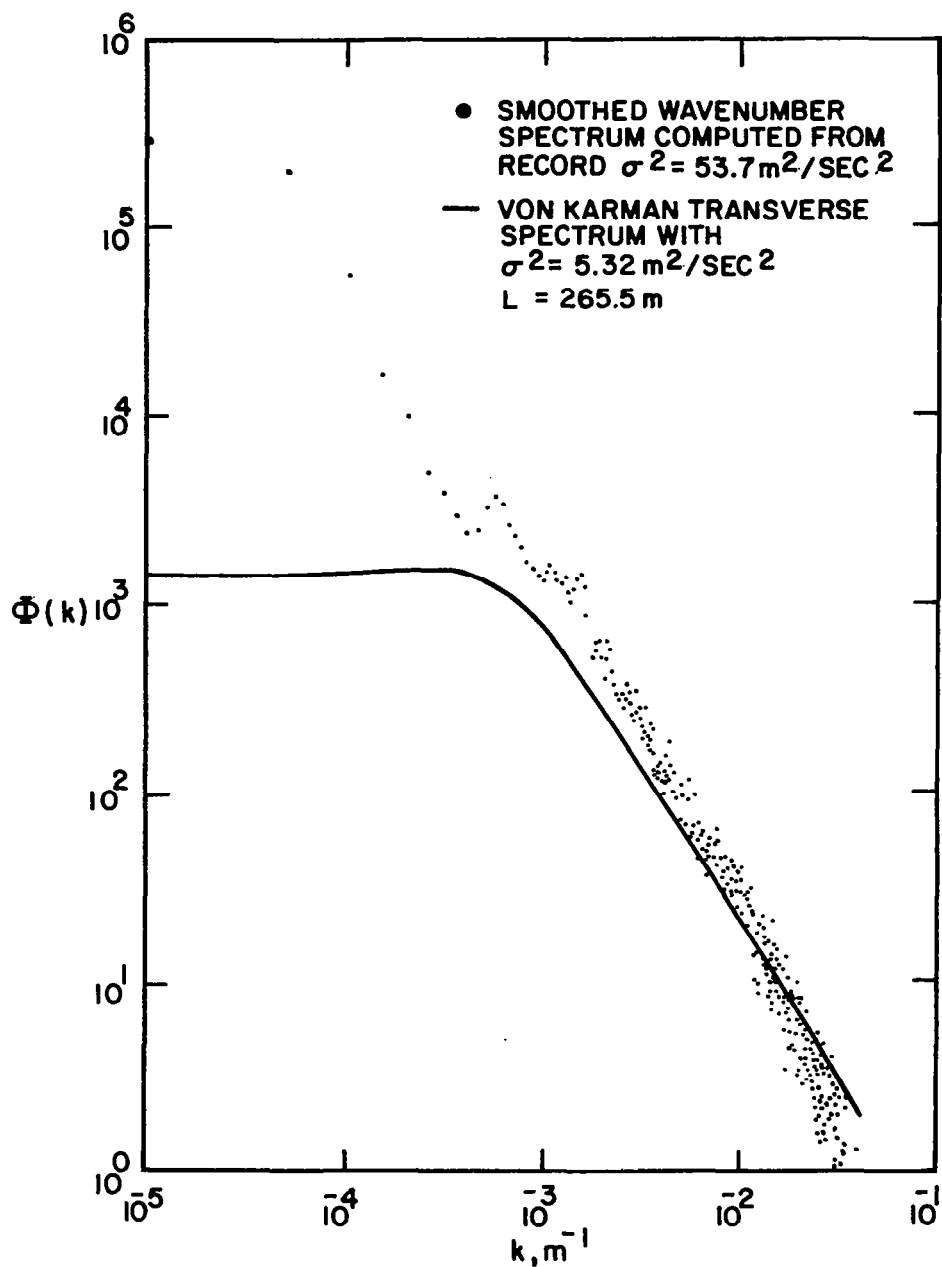


FIG. 5. COMPARISON OF SMOOTHED WAVENUMBER SPECTRUM COMPUTED FROM LATERAL RECORD SHOWN IN FIG. 4 AND von KARMAN TRANSVERSE SPECTRUM OBTAINED BY CONSTRAINED LEAST-SQUARES FIT TO THE (EMPIRICAL) AUTOCORRELATION FUNCTION. Von KARMAN SPECTRUM CHARACTERIZES "FAST" TURBULENCE COMPONENT ONLY.

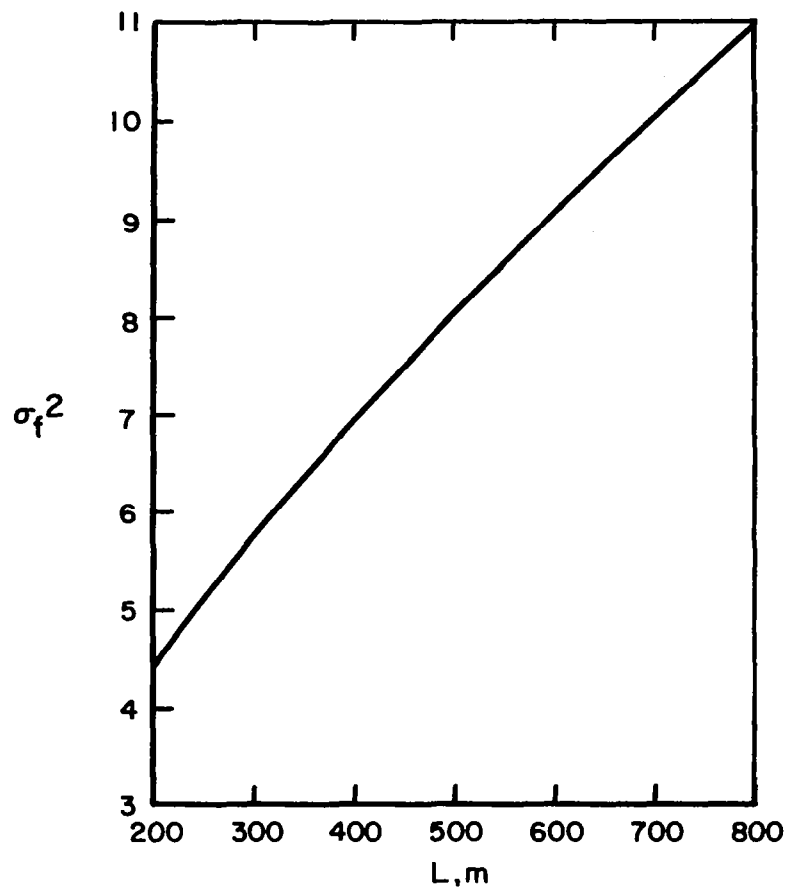


FIG. 6. CONSTRAINT BETWEEN σ_f^2 AND L FOR CONSTRAINED LEAST-SQUARES ESTIMATION PROCEDURE APPLIED TO LATERAL RECORD SHOWN IN FIG. 4.

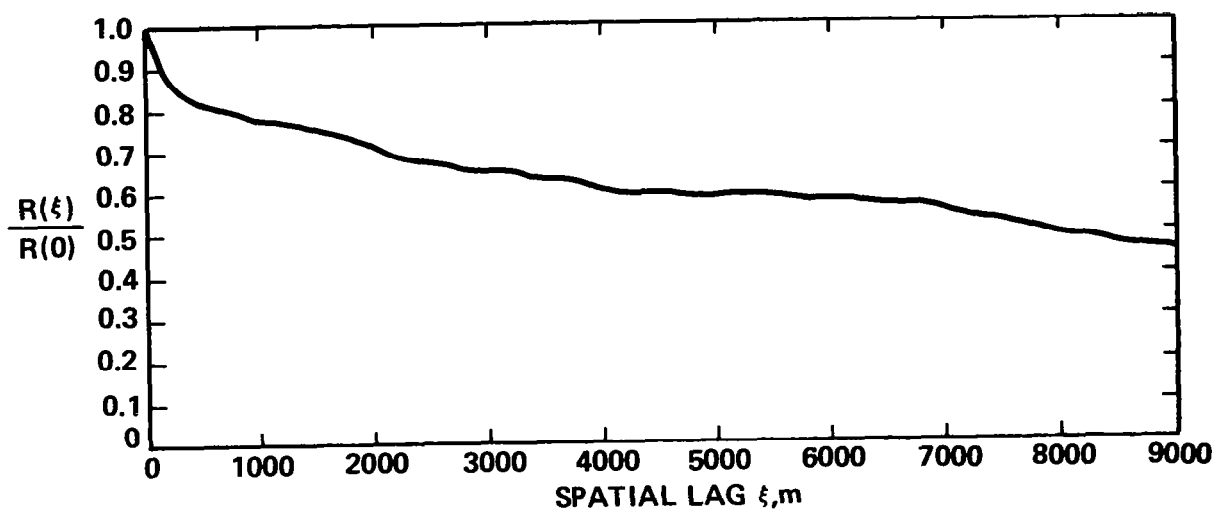


FIG. 7. AUTOCORRELATION FUNCTION OF LATERAL RECORD SHOWN IN FIG. 4.

for the autocorrelation function shown in Fig. 7, the method was exercised a number of times for the values of ξ_H and m listed in Table 1.

Four autocorrelation function representations $\phi(\xi)$ given by Eq. (3.2) for four different sets of values of ξ_H and m are plotted on expanded scales in Fig. 8 along with the empirical autocorrelation function $R(\xi)$ of the lateral record shown in Fig. 4. The von Karman transverse spectrum of Eq. (2.1) for one of these cases ($\xi_H = 5998.1$ m and $m = 3$) is plotted in Fig. 5 for comparison with the "high-frequency" portion of the empirical spectrum. The values of L and $\sigma_f^2 = \sigma^2$ used in the evaluation of Eq. (2.1) shown in Fig. 5 are those given by Eq. (3.1), which were taken from Table 1. From Fig. 5, we see that the asymptotic slope of the empirical spectrum is somewhat steeper than the $-5/3$ slope of the von Karman spectrum. Hence, for this record, one of the basic assumptions in the constrained least-squares fit method is not well satisfied. Because of this discrepancy in slopes, none of four curves $\phi(\xi)$ shown in Fig. 8 fits well the "von Karman" region of the autocorrelation function near $\xi = 0$. Nevertheless, the knee of the von Karman spectrum shown in Fig. 5 would appear to have about the right position.

Figure 9 displays the autocorrelation function of the lateral component of the wind shear record with the von Karman autocorrelation component $\sigma_f^2 \phi_K(\xi; L)$ removed. That is, the solid curve in Fig. 8 is the autocorrelation function $R(\xi)$ of the lateral record shown in Fig. 4 after subtraction of the von Karman autocorrelation function component

$$\phi_{KT}(\xi) \equiv \sigma_f^2 \phi_K(\xi; L) \quad (3.3)$$

evaluated from Eq. (2.4) with the values of L and σ^2 given by Eq. (3.1). The dashed curve is the cubic (3rd degree of polynomial) that best represents the solid curve in an integral least-squares sense over the lag region from 0 to 10,000 meters. We see from Fig. 9 that a third-degree polynomial represents very nicely the autocorrelation function of the slow turbulence component $w_s(t)$ over a 10,000 meter lag interval. Such polynomial representations are the characterizations suggested in Sec. 1 of Ref. 1 for describing the "slow" turbulence component $w_s(t)$ for aircraft response calculations.

TABLE 1. CONSTRAINED LEAST-SQUARES ESTIMATION OF AUTOCORRELATION FUNCTION PARAMETERS FOR WIND-SHEAR LATERAL RECORD

ξ_H m	m	σ_f^2 m ² /sec ²	L m	$\phi(0)$	a_0	a_1	a_2	a_3
1004.5	1	5.518	281.5	52.28	46.77	$-.582 \times 10^{-2}$		
1999.6	1	6.486	361.1	51.68	45.19	$-.341 \times 10^{-2}$		
3004.2	1	6.046	324.1	51.76	45.14	$-.390 \times 10^{-2}$		
3004.2	2	5.793	303.6	51.81	46.02	$-.431 \times 10^{-2}$	$.115 \times 10^{-6}$	
4496.9	2	5.428	274.7	51.96	46.53	$-.517 \times 10^{-2}$	$.409 \times 10^{-6}$	
5998.9	2	5.265	261.8	51.99	46.73	$-.542 \times 10^{-2}$	$.468 \times 10^{-6}$	
5998.9	3	5.315	265.5	51.98	46.66	$-.533 \times 10^{-2}$	$.434 \times 10^{-6}$	$.353 \times 10^{-11}$
7998.6	3	4.459	201.0	52.29	47.83	$-.723 \times 10^{-2}$	$.116 \times 10^{-5}$	$-.739 \times 10^{-10}$
9998.2	3	4.742	221.8	52.20	47.46	$-.674 \times 10^{-2}$	$.102 \times 10^{-5}$	$-.624 \times 10^{-10}$
9998.2	4	4.530	206.3	52.28	47.75	$-.721 \times 10^{-2}$	$.122 \times 10^{-5}$	$-.913 \times 10^{-10}$
								$a_4 = .141 \times 10^{-14}$

Exact value of $R(0)$ is $53.66 \text{ m}^2/\text{sec}^2$.

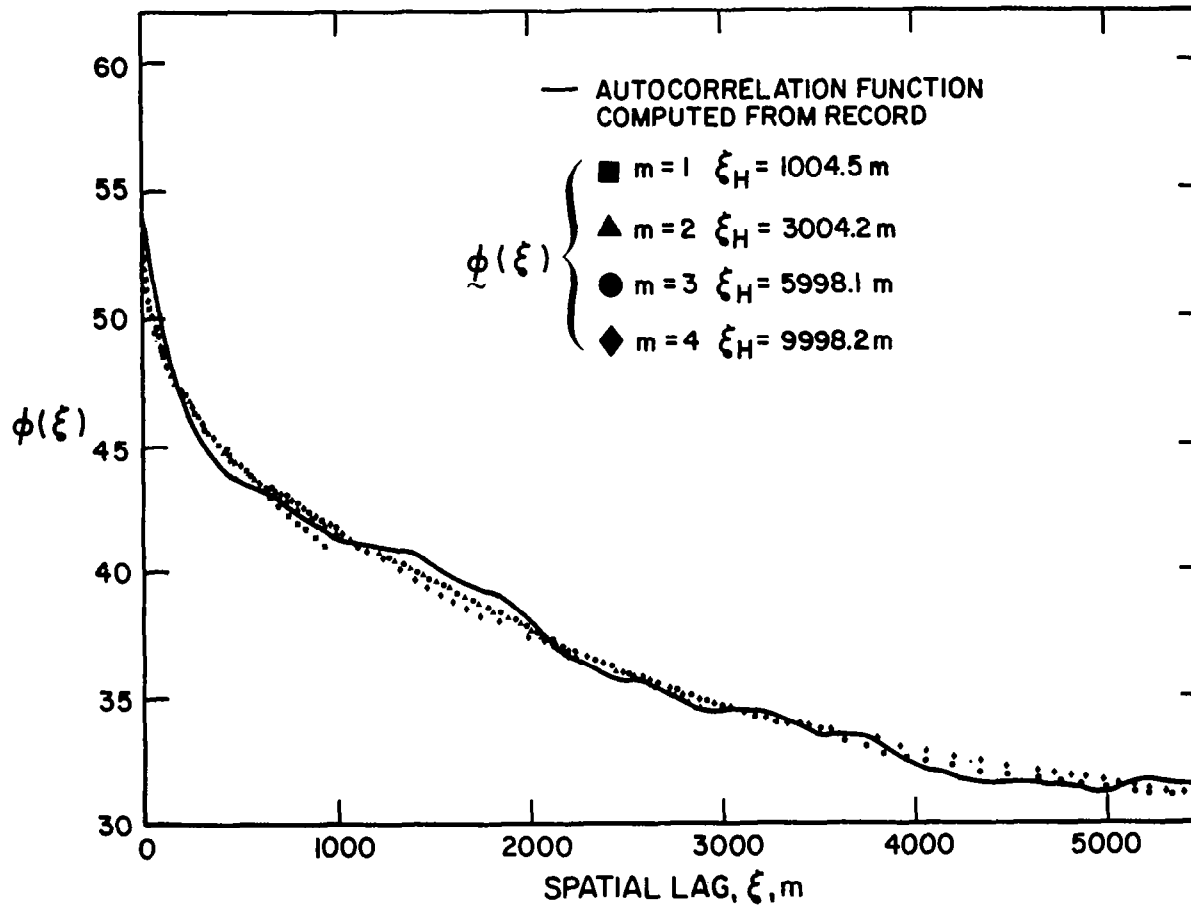


FIG. 8. COMPARISON OF AUTOCORRELATION FUNCTION COMPUTED FROM LATERAL RECORD SHOWN IN FIG. 4 AND CONSTRAINED LEAST-SQUARES FIT OF AUTOCORRELATION MODEL OF EQ. (3.2).

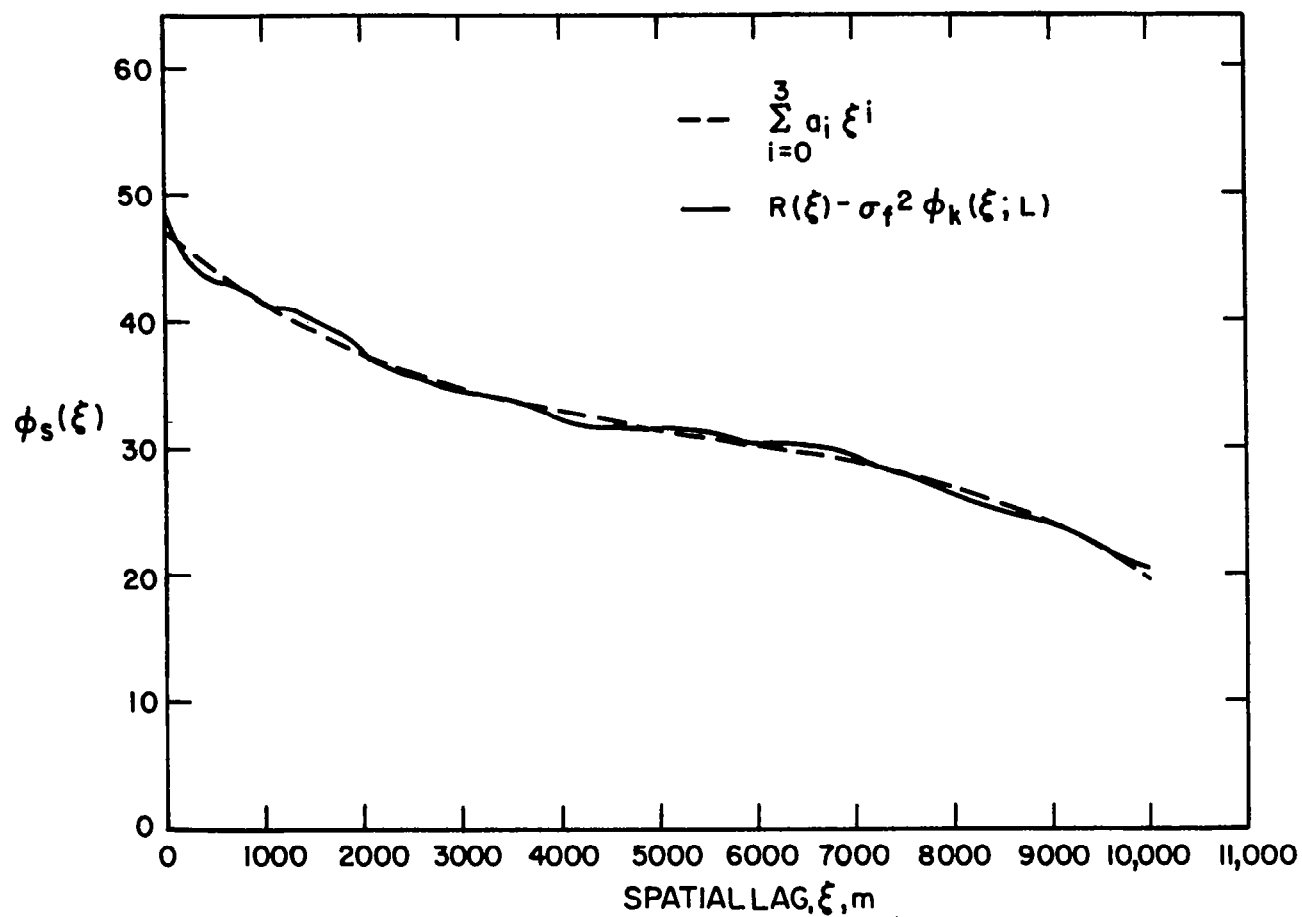


FIG. 9. COMPARISON OF AUTOCORRELATION FUNCTION $R(\xi)$ OF LATERAL RECORD SHOWN IN FIG. 4 MINUS AUTOCORRELATION FUNCTION $\sigma_f^2 \phi_k(\xi; L)$ OF von KARMAN COMPONENT AND INTEGRAL LEAST-SQUARES THIRD-DEGREE POLYNOMIAL APPROXIMATION.

CONSTRAINED LEAST-SQUARES ESTIMATION OF AUTOCORRELATION FUNCTION PARAMETERS OF LATERAL RECORD FROM FLIGHT 30 RUN 8 (MOUNTAIN-WAVE CONDITIONS)

The lateral record shown in Fig. 10 illustrates a record with an exceptionally strong low-frequency component $w_s(t)$ relative to the "fast" component $w_f(t)$. Figures 11 to 15 illustrate, respectively, the same quantities for the lateral record shown in Fig. 10 that Figs. 5 to 9 displayed for the lateral record in Fig. 4. Similarly, Table 2 displays for the lateral record in Fig. 10 quantities comparable to the quantities displayed in Table 1 for the lateral record in Fig. 4. Computations of the material in Table 2 and Figs. 11 to 15 were carried out using the same methods as in the case of Table 1 and Figs. 5 to 9.

The value of M used in computing the empirical spectrum in Fig. 11 was 10,089 m which corresponds to 1024 temporal sample points. The von Karman transverse spectrum plotted in Fig. 11 was computed using the parameter values

$$L = 128.9 \text{ m}, \sigma^2 = 0.684 \text{ m}^2/\text{sec}^2 \quad (4.1)$$

which correspond to the case $\xi_H = 2295.6$ meters and a 2nd degree polynomial ($m=2$) in the autocorrelation function representation of Eq. (3.2).

In computing the constraint relationship between σ_f^2 and L displayed in Fig. 12, the lower and upper wavenumbers used were $k_l = 10^{-3}\text{m}^{-1}$ and $k_u = 4 \times 10^{-2}\text{m}^{-1}$.

Discussion. The four fits to the empirical autocorrelation function shown in Fig. 14(a) illustrates misleading results that the method described in Sec. 4 of Ref. 1 can yield when it is not used properly. Although each of the four fits provided by Eq. (3.2) to the empirical autocorrelation function appears reasonable to the eye, reference to Table 2 shows that the largest integral scale obtained for these four cases is that corresponding to $\xi_H = 699.5$ meters and $m = 1$ which yielded $L = 118.6$ meters, whereas the next largest value of L for the four cases is $L = 69.9$ meters for the case $\xi_H = 896.6$ meters and $m = 2$. The discrepancy between these two values of L is quite large. The problem here is that the polynomial in the right-band of Eq. (3.2) is actually representing part of the von Karman portion of the empirical

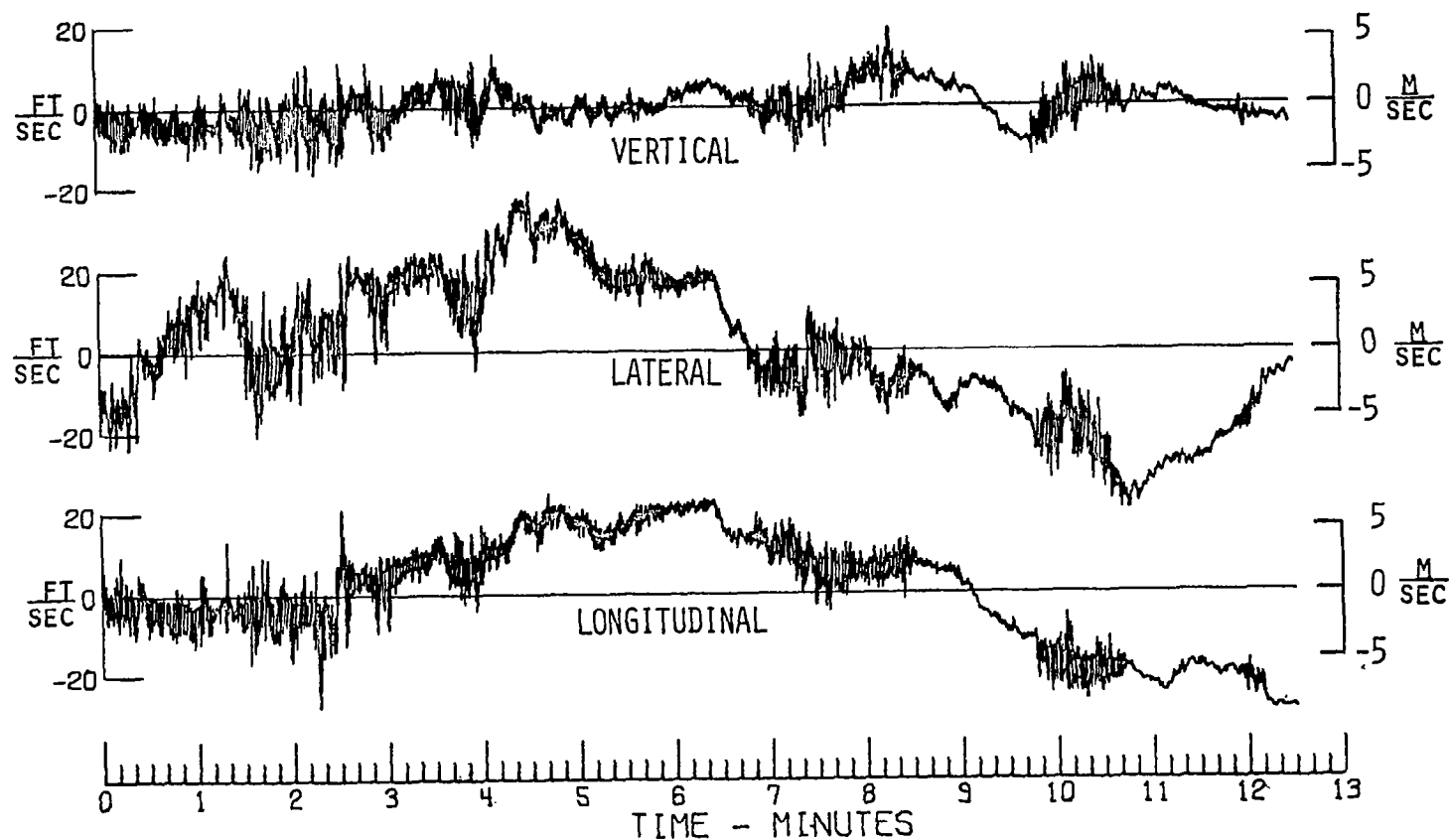


FIG. 10. TURBULENCE RECORDS CONTAINING EXCEPTIONALLY STRONG "SLOW" COMPONENTS $w_s(t)$.
 [MOUNTAIN WAVE CONDITIONS. AIRCRAFT SPEED 197 m/sec (646 ft/sec).] (Ref. 3,
 FIG. 10, p. 285.)

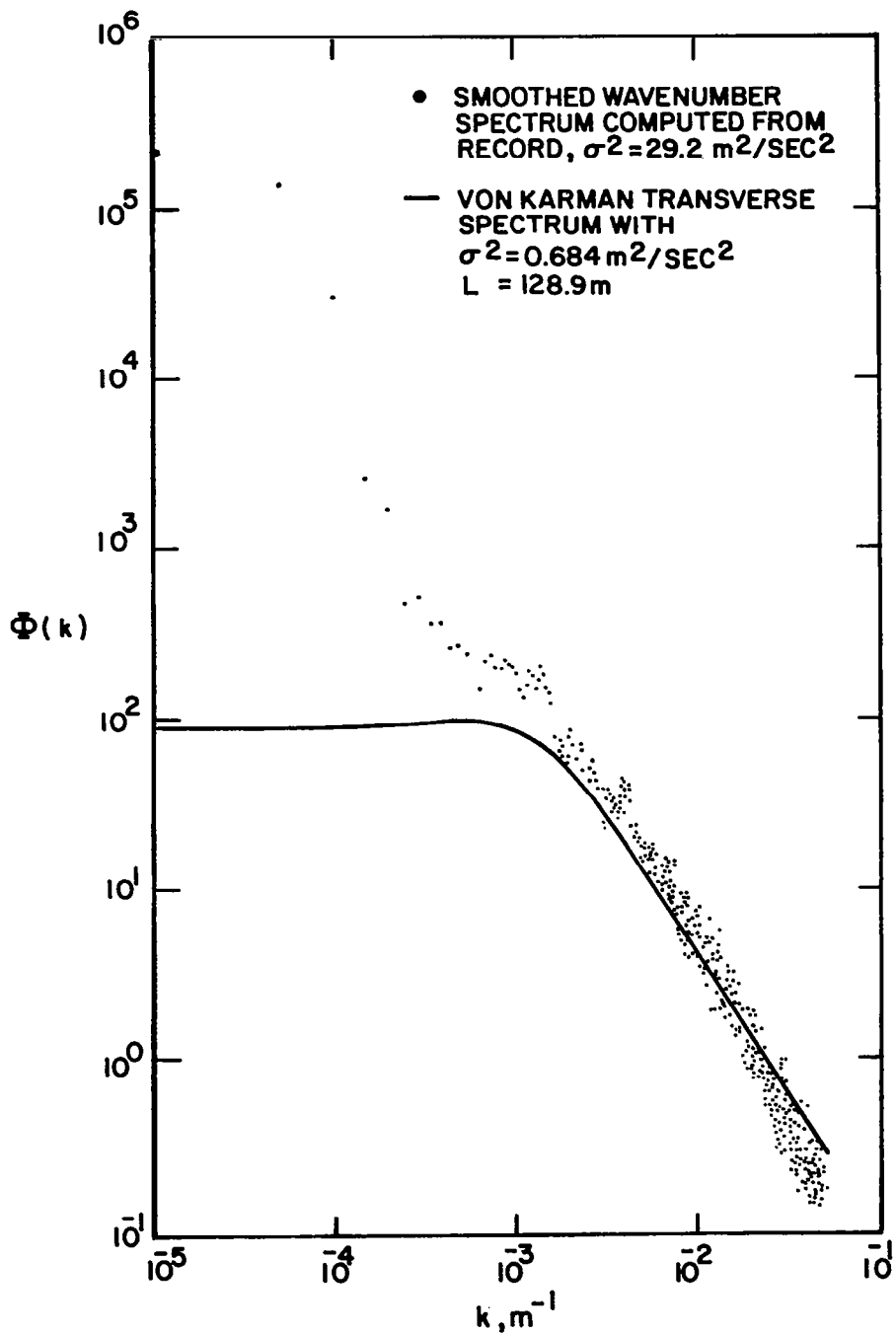


FIG. 11. COMPARISON OF SMOOTHED WAVENUMBER SPECTRUM COMPUTED FROM LATERAL RECORD SHOWN IN FIG. 10 AND von KARMAN TRANSVERSE SPECTRUM OBTAINED BY CONSTRAINED LEAST-SQUARES FIT TO THE (EMPIRICAL) AUTO-CORRELATION FUNCTION. Von KARMAN SPECTRUM CHARACTERIZES "FAST" TURBULENCE COMPONENT ONLY.

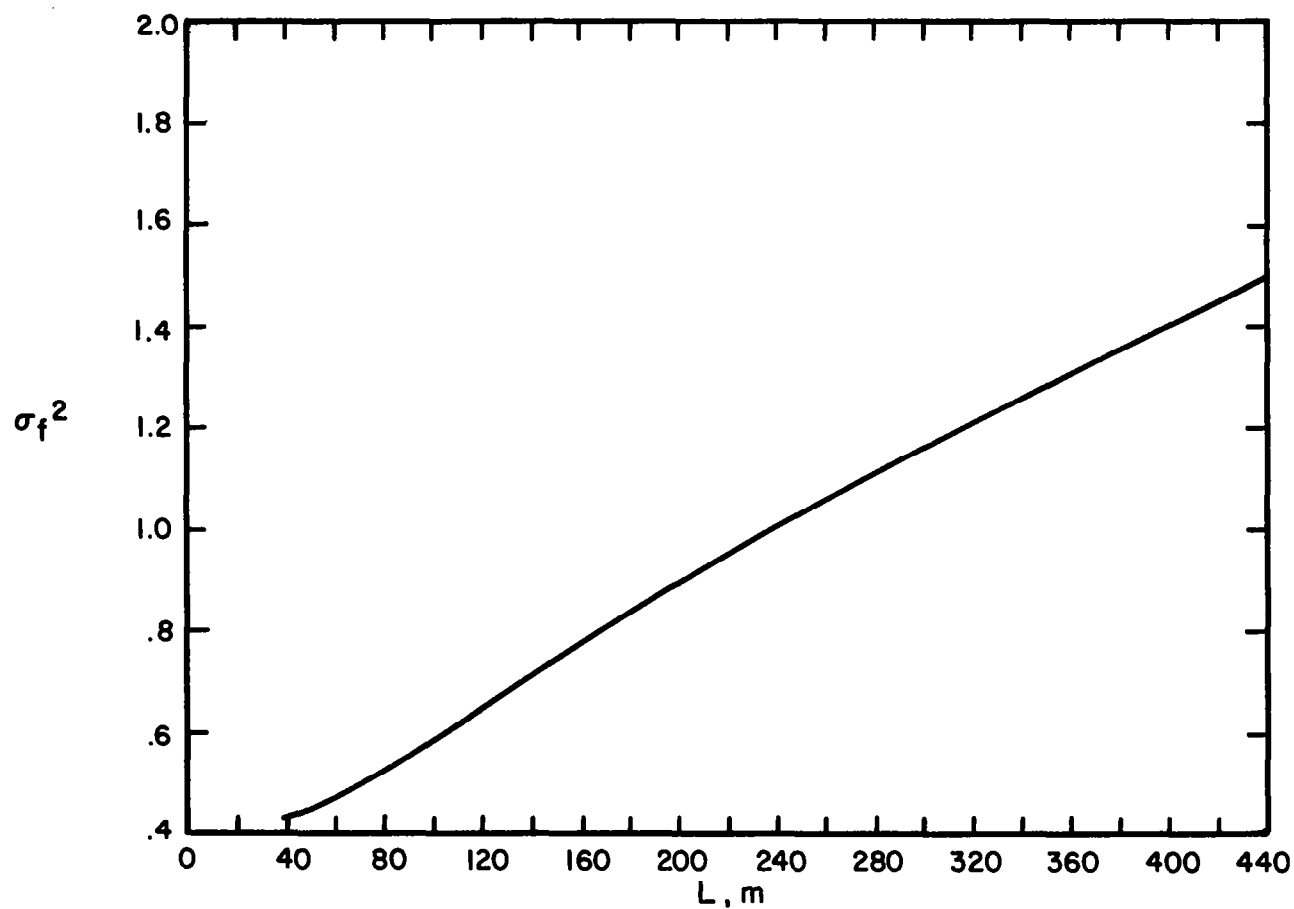


FIG. 12. CONSTRAINT BETWEEN σ_f^2 AND L FOR CONSTRAINED LEAST-SQUARES ESTIMATION PROCEDURE APPLIED TO LATERAL RECORD SHOWN IN FIG. 10.

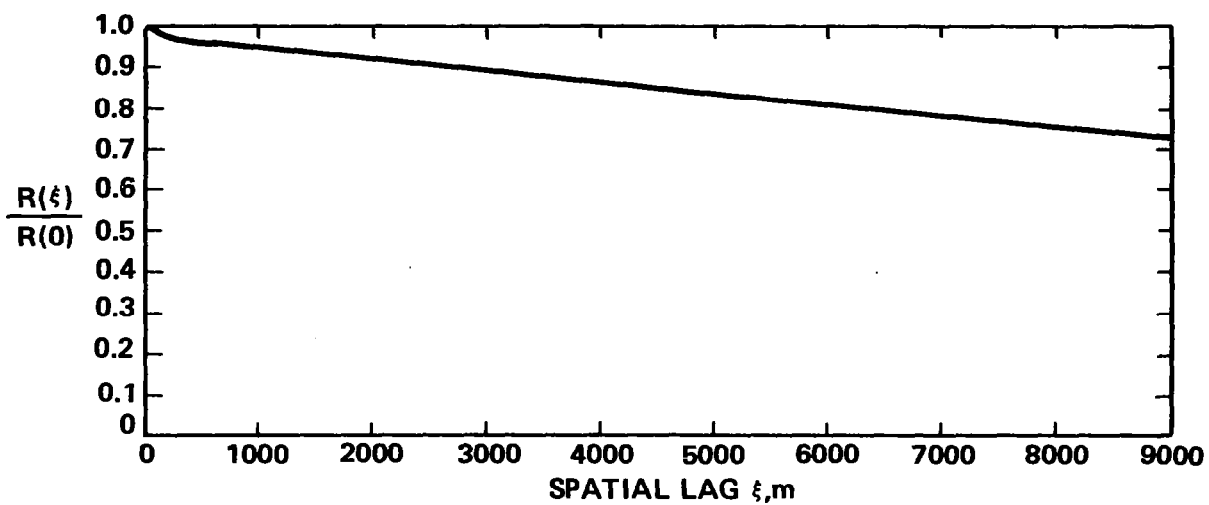


FIG. 13. AUTOCORRELATION FUNCTION OF LATERAL RECORD SHOW IN FIG. 10.
[FROM MAT PROJECT, NASA LANGLEY RESEARCH CENTER.]

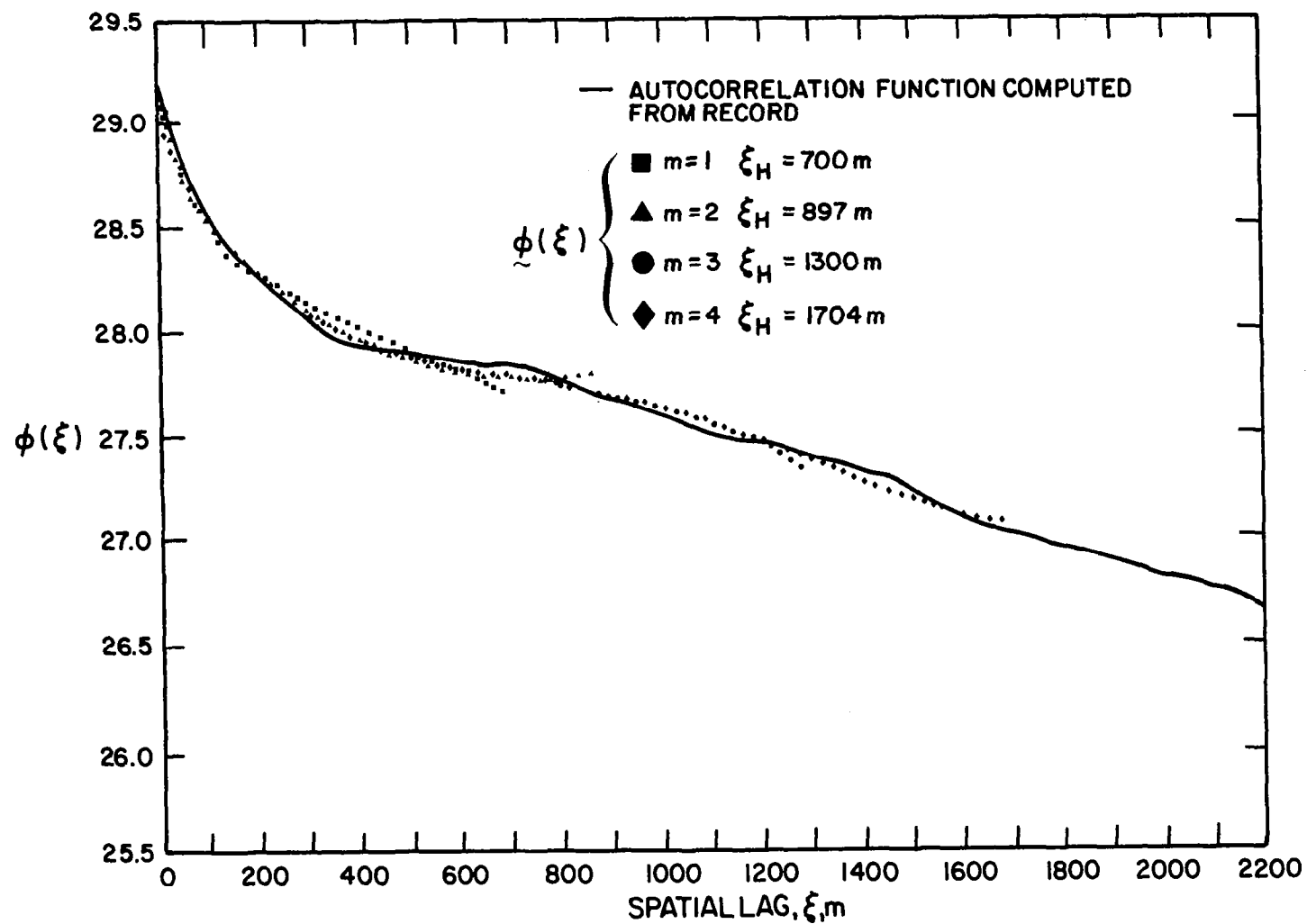
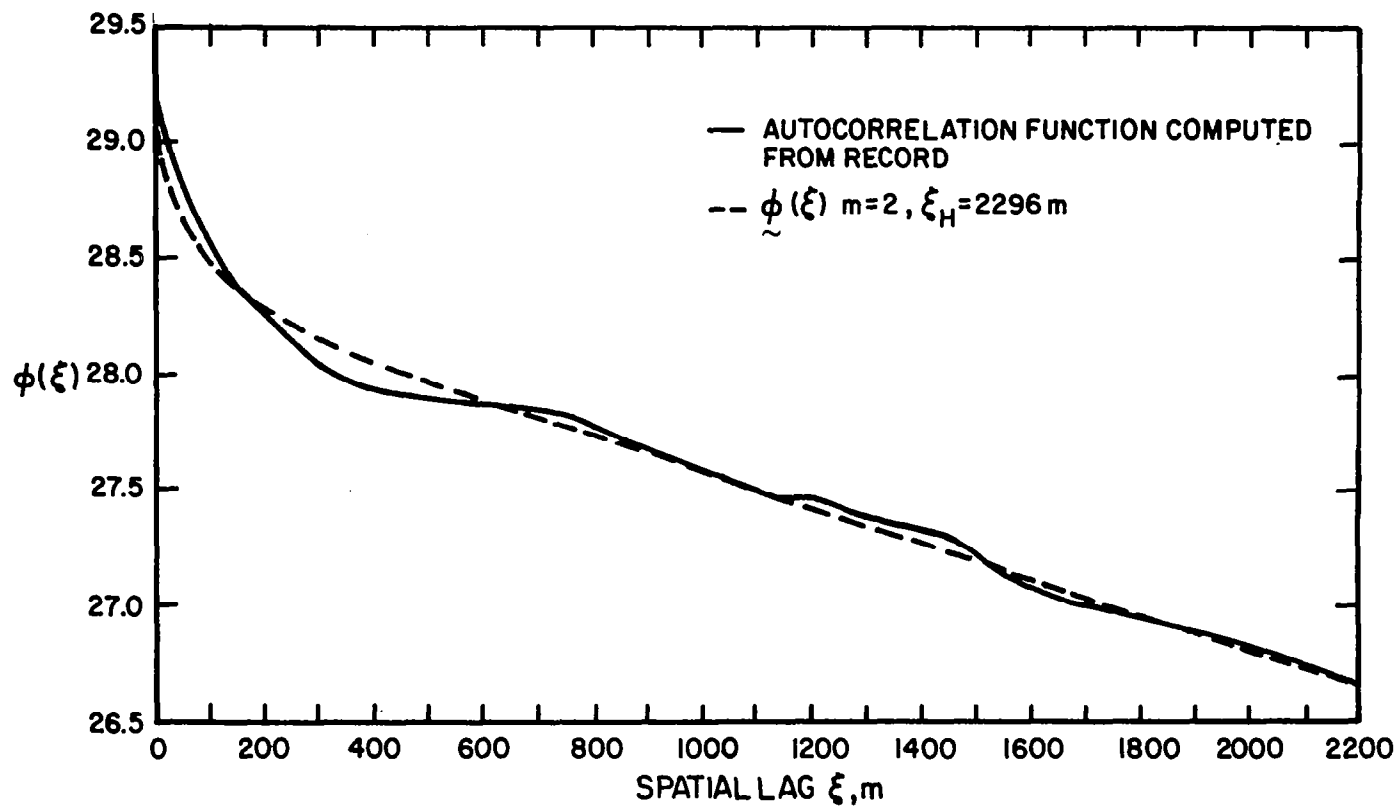


FIG. 14(a). COMPARISON OF AUTOCORRELATION FUNCTION COMPUTED FROM LATERAL RECORD SHOWN IN FIG. 10 AND CONSTRAINED LEAST-SQUARES FIT OF AUTOCORRELATION MODEL OF EQ. (3.2). VALUES OF ξ_H USED IN OBTAINING THE ABOVE RESULTS WERE TOO SMALL.



3

FIG. 14(b). COMPARISON OF AUTOCORRELATION FUNCTION COMPUTED FROM LATERAL RECORD SHOWN IN FIG. 10 AND CONSTRAINED LEAST-SQUARES FIT OF AUTOCORRELATION MODEL OF EQ. (3.2).

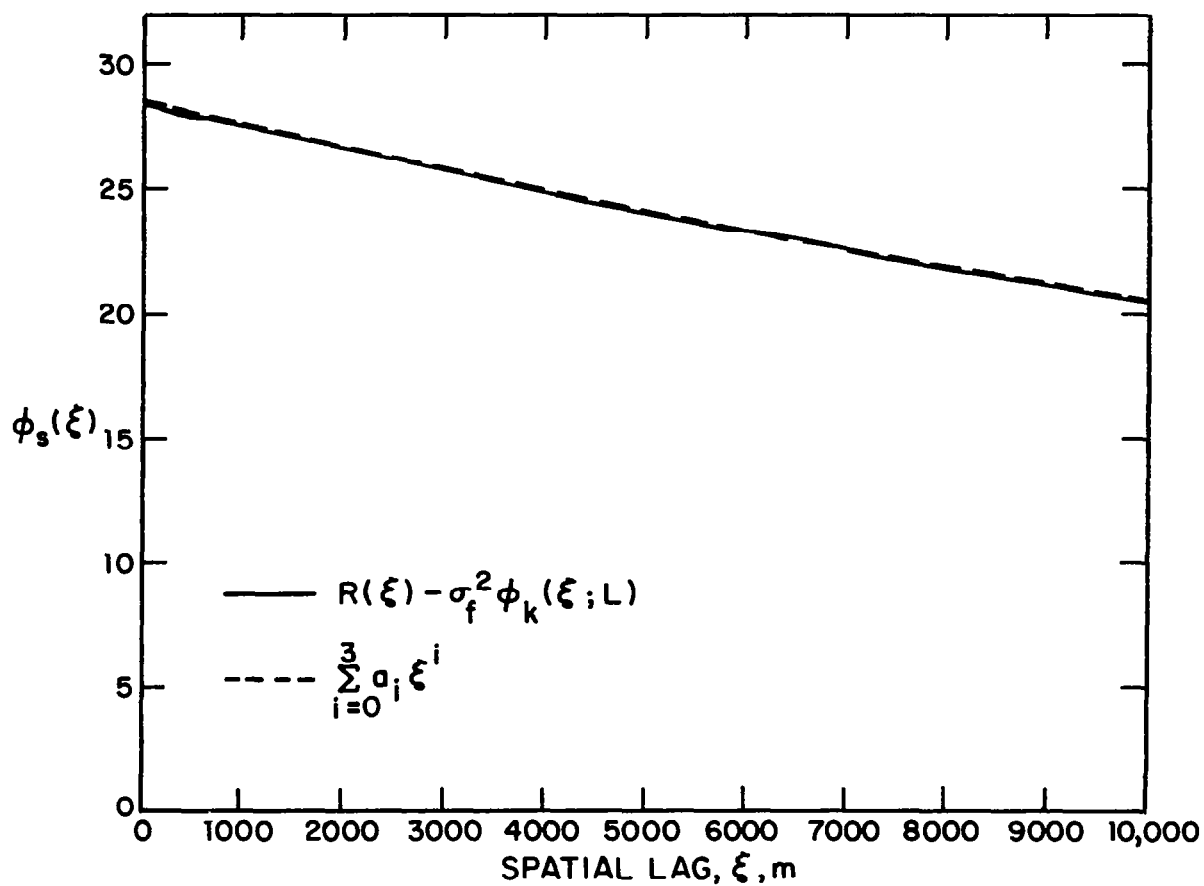


FIG. 15. COMPARISON OF AUTOCORRELATION FUNCTION $R(\xi)$ OF LATERAL RECORD SHOWN IN FIG. 10 MINUS AUTOCORRELATION FUNCTION $\sigma_f^2(\xi; L)$ OF von KARMAN COMPONENT AND INTEGRAL LEAST-SQUARES THIRD-DEGREE POLYNOMIAL APPROXIMATION.

TABLE 2. CONSTRAINED LEAST-SQUARES ESTIMATION OF AUTOCORRELATION FUNCTION PARAMETERS FOR MOUNTAIN-WAVE LATERAL RECORD

ξ_H m	m	σ_f^2 m ² /sec ²	L m	$\phi(0)$	a_0	a_1	a_2	a_3
502.5	0	.944	217.2	28.97	28.03			
699.5	1	.651	118.6	29.12	28.47	$-.109 \times 10^{-2}$		
699.5	2	.467	57.1	29.35	28.89	$-.376 \times 10^{-2}$	$.340 \times 10^{-5}$	
896.6	2	.495	69.9	29.26	28.76	$-.272 \times 10^{-2}$	$.187 \times 10^{-5}$	
1103.5	3	.455	50.8	29.38	28.92	$-.439 \times 10^{-2}$	$.612 \times 10^{-5}$	$-.305 \times 10^{-8}$
1300.5	3	.490	67.4	29.30	28.81	$-.331 \times 10^{-2}$	$.373 \times 10^{-5}$	$-.159 \times 10^{-8}$
1704.5	4	.474	60.2	29.37	28.89	$-.431 \times 10^{-2}$	$.670 \times 10^{-5}$	$-.485 \times 10^{-8}$
								$a_4 = .119 \times 10^{-11}$
2295.6	2	.684	128.9	29.07	28.39	$-.821 \times 10^{-3}$	$.171 \times 10^{-7}$	
2502.5	3	.653	119.1	29.12	28.47	$-.117 \times 10^{-2}$	$.360 \times 10^{-6}$	$-.933 \times 10^{-10}$

Exact value of $R(0)$ is $29.22 \text{ m}^2/\text{sec}^2$

autocorrelation function that occurs near the origin $\xi = 0$. When this misleading behavior takes place, the value of σ_f^2 in Eq. (3.2) obtained by the least integral-squared fit is somewhat smaller than it should be; hence, the value obtained for the integral scale L also is too small as can be seen from Fig. 12. *This misleading behavior can largely be avoided by choosing ξ_H as large as possible and m as small as possible consistent with achieving a "reasonable" representation of the autocorrelation function of the "slow" component $w_s(t)$ by the polynomial in Eq. (3.2).** Following this rule, we would never expect to have to choose m larger than 3. Mathematically, this misleading behavior is a consequence of the fact that the polynomial in the right-hand side of Eq. (3.2) is not orthogonal to the von Karman autocorrelation function which is the first term in the right-hand side.

In the present example, this problem is aggravated by the fact that the asymptotic slope of the empirical spectrum shown in Fig. 11 is somewhat steeper than the $-5/3$ asymptotic slope of the von Karman spectrum also shown in Fig. 11. (Thus, in this case also, one of the basic assumptions used in developing the method is not satisfied by the turbulence data.)

If we follow the above italicized rule, we see from the empirical autocorrelation function shown in Fig. 14(a) that for ξ larger than about 750 meters, we cannot reasonably use $m = 1$ [which is a linear approximation to the autocorrelation function of the "slow" component $w_s(t)$]; however, if we let m increase to a value of 2 (quadratic approximation) then we can reasonably choose ξ_H to be 2200 meters. Using $\xi_H = 2295.6$ meters and $m = 2$, we obtained the fit shown in Fig. 14(b) which yielded the values of L and σ_f^2 given by Eq. (4.1).

Upon first inspection of Fig. 11, the value of $L = 128.9$ m given by Eq. (4.1) appears too small. However, closer inspection shows several (about 4) weak resonances between $k = 7.5 \times 10^{-4} \text{m}^{-1}$ and $k = 1.43 \times 10^{-3} \text{m}^{-1}$, with the spectrum dropping off fairly abruptly beyond the latter value until it almost reaches the von Karman curve. Hence, we should expect several weak oscillations in the autocorrelation function with periods equal to the reciprocal values of the above frequencies - i.e., periods ranging from $0.70 \times 10^3 = 700$ m to $0.13 \times 10^4 = 1300$ m. Referring to Fig. 14(b), we see that our least-squares fit underestimates the empirical autocorrelation function at the origin, overestimates at

*Trade-off criteria between choices of ξ_H and m are discussed in detail in Appendix G of Ref. 1.

$\xi = 350$ m and again underestimates at $\xi = 700$ m, which according to the above is one period of the oscillation with the smallest period. At $\xi = 750$ m we see another peak in the empirical autocorrelation function, which is the period of the resonance located in the spectrum slightly to the left of the resonance at $k = 1.43 \times 10^{-3} \text{m}^{-1}$. Looking back in the region of $\xi \approx 350$ to 400 m in Fig. 14(b), we see that these two oscillations have added in phase in that region to produce a relatively large discrepancy between the empirical autocorrelation function and our least squares fit. Finally, in the region of Fig. 14(b) between $\xi = 1200$ m and $\xi = 1450$ m we observe all 4 of the above oscillations adding almost in phase in this region - as expected from the appearance of the spectrum. Hence, some destructive interference must have occurred in the region near $\xi = 750$ m which further explains why the discrepancy between the empirical and least-squares curves is less near $\xi = 750$ m than near $\xi = 350$ m. Hence, the main discrepancies between the empirical and least-squares fit of the autocorrelation function shown in Fig. 14(b) can be explained by the approximately 4 weak resonances between $k = 7.5 \times 10^{-4} \text{m}^{-1}$ and $1.43 \times 10^{-3} \text{m}^{-1}$ in Fig. 11. If these resonances were removed from the empirical spectrum shown in Fig. 11, the knee of the von Karman curve would appear to be in about the right position. Hence, the values of L and σ^2 given by Eq. (4.1) - which characterize the von Karman component of the turbulence - appear to be about right.

Finally, we note that the presence of spectral peaks or "resonances" such as those discussed above will tend to bias the values of L and σ^2 obtained using the maximum likelihood method developed in Sec. 3 of Ref. 1 since this method assumes that the turbulence obeys the von Karman spectral form. In contrast, such spectral peaks produce oscillations in the autocorrelation function, and the constrained least-squares estimation procedure developed in Sec. 4 of Ref. 1 tends to average out - i.e., ignore - these oscillations. The constrained least-squares procedure therefore should produce better estimates of the integral scale and intensity of the von Karman component of turbulence records in these situations.

CONSTRAINED LEAST-SQUARES AND MAXIMUM LIKELIHOOD ESTIMATION OF AUTOCORRELATION FUNCTION PARAMETERS OF VERTICAL RECORD FROM FLIGHT 30 RUN 8 (MOUNTAIN-WAVE CONDITIONS)

Figures 16 to 20 illustrate, respectively, the same quantities for the vertical record shown in Fig. 10 that Figs. 11 to 15 displayed for the lateral record in Fig. 10. Similarly, Table 3 displays for the vertical record in Fig. 10 quantities comparable to the quantities displayed in Table 2 for the lateral record in Fig. 10. The computations for the material in Table 3 and Figs. 16 to 20 were carried out using the same methods as used in the cases of Tables 1 and 2 and Figs. 5 to 15.

The value of M used in computing the empirical spectrum in Fig. 16 was 10,089 m which corresponds to 1024 temporal sample points. This is the same value of M as used in computation of the spectrum in Fig. 11 for the lateral component.

The von Karman transverse spectrum plotted in Fig. 16 was computed using the parameter values

$$L = 68.4 \text{ m}, \sigma^2 = 0.470 \text{ m}^2/\text{sec}^2 \quad (5.1)$$

which correspond to the case $\xi_H = 1202$ meters and a second degree polynomial ($m=2$) in the autocorrelation function representation of Eq. (3.2).

In computing the constraint relationship between σ_f^2 and L displayed in Fig. 17, the lower and upper wavenumbers used were the same values as those used in the previous two examples — namely, $k_l = 10^{-3} \text{ m}^{-1}$ and $k_u = 4 \times 10^{-2} \text{ m}^{-1}$.

Discussion. The vertical record displayed in Fig. 10 is much better behaved than either of the previous two records studied, and this improved behavior is reflected in our results. The record is better behaved for four reasons: (1) the asymptotic slope of the empirical spectrum shown in Fig. 16 agrees very well with the asymptotic slope of $-5/3$ of the von Karman spectrum; (2) the empirical spectrum in Fig. 16 has a better developed "knee" than the empirical spectra shown in Figs. 5 and 11; (3) the fractional energy in the "resonances" in the spectrum in Fig. 16 is less than in the previous two cases as may be seen by comparing the

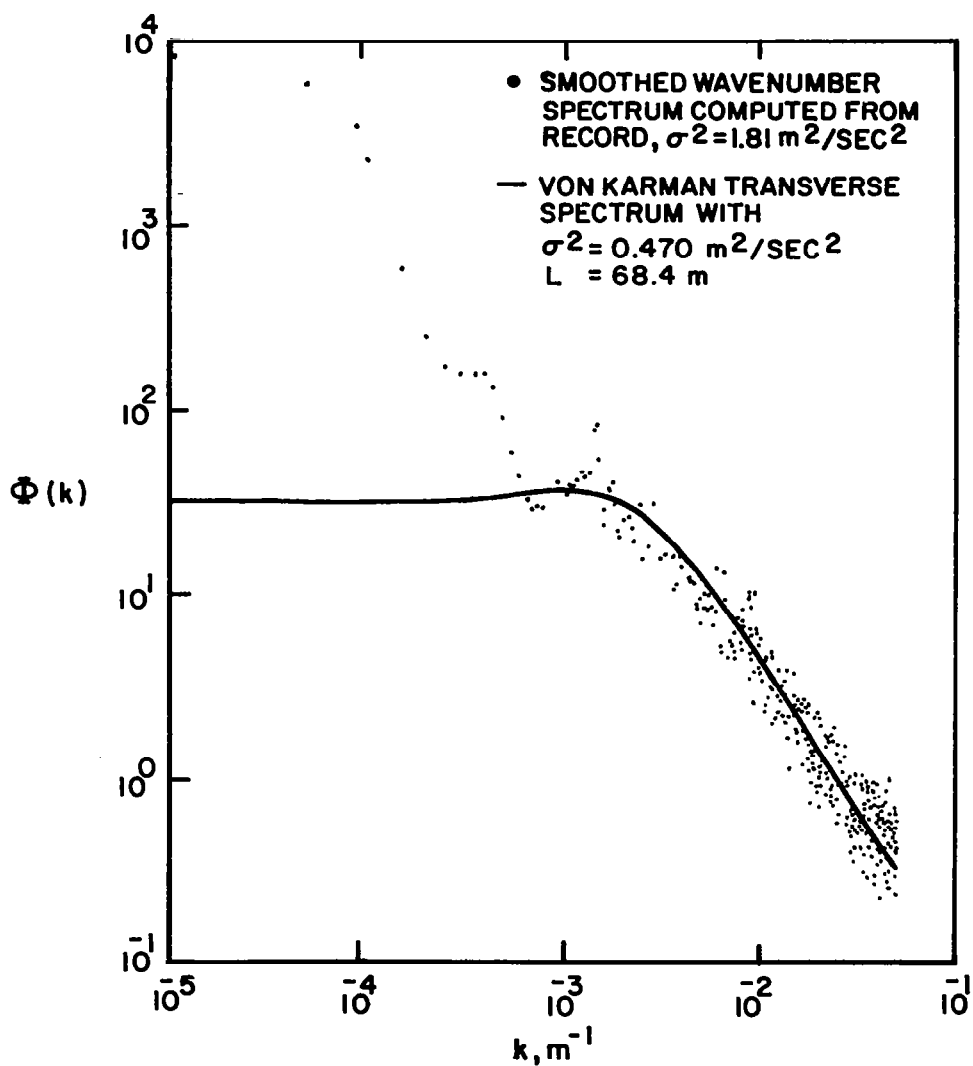


FIG. 16. COMPARISON OF SMOOTHED WAVENUMBER SPECTRUM COMPUTED FROM VERTICAL RECORD SHOWN IN FIG. 10 AND von KARMAN TRANSVERSE SPECTRUM OBTAINED BY CONSTRAINED LEAST-SQUARES FIT TO THE (EMPIRICAL) AUTOCORRELATION FUNCTION. Von KARMAN SPECTRUM CHARACTERIZES "FAST" TURBULENCE COMPONENT ONLY.

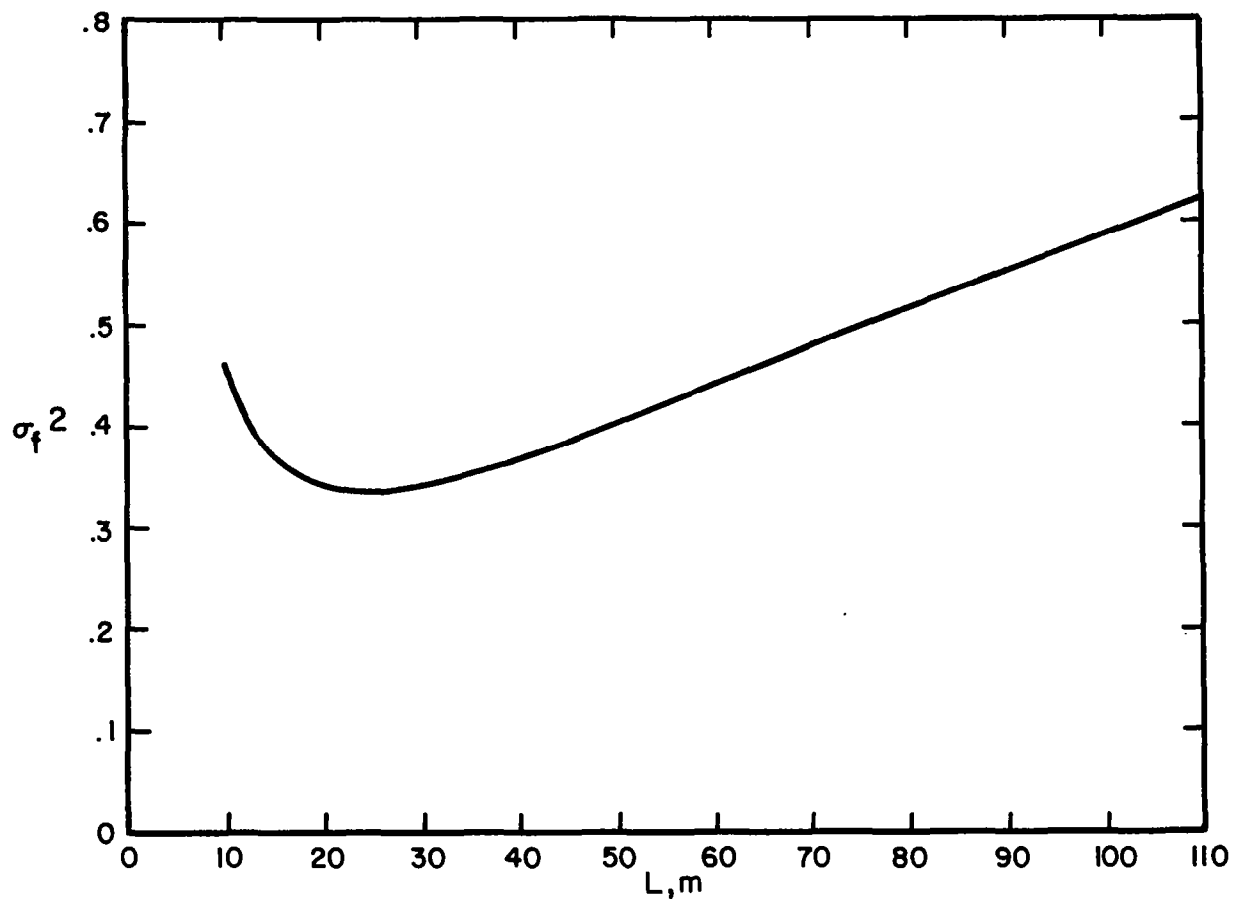


FIG. 17. CONSTRAINT BETWEEN σ_f^2 AND L FOR CONSTRAINED LEAST-SQUARES ESTIMATION PROCEDURE APPLIED TO VERTICAL RECORD SHOWN IN FIG. 10.

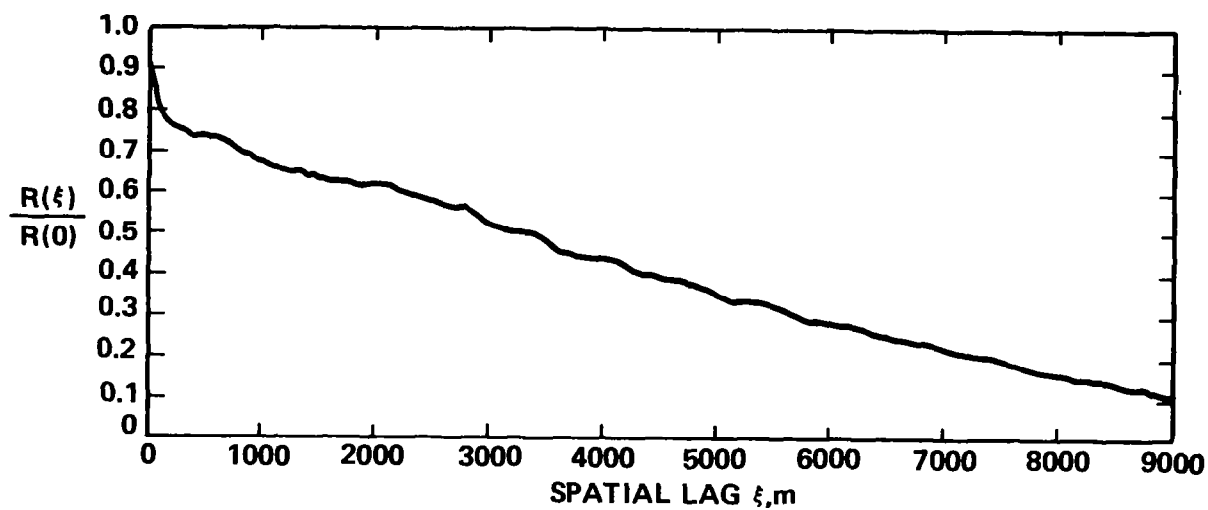


FIG. 18. AUTOCORRELATION FUNCTION OF VERTICAL RECORD SHOWN IN FIG. 10.
[FROM MAT PROJECT, NASA LANGLEY RESEARCH CENTER.]

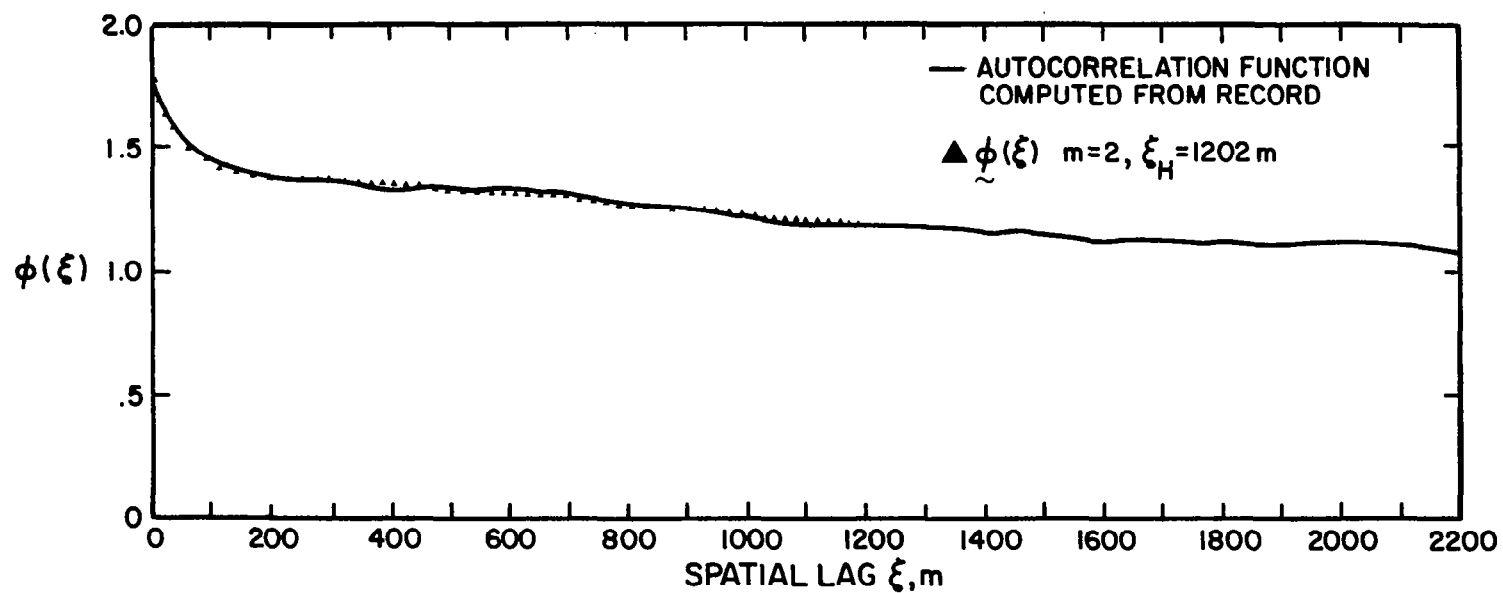


FIG. 19. COMPARISON OF AUTOCORRELATION FUNCTION COMPUTED FROM VERTICAL RECORD SHOWN IN FIG. 10 AND CONSTRAINED LEAST-SQUARES FIT OF AUTOCORRELATION MODEL OF EQ. (3.2).

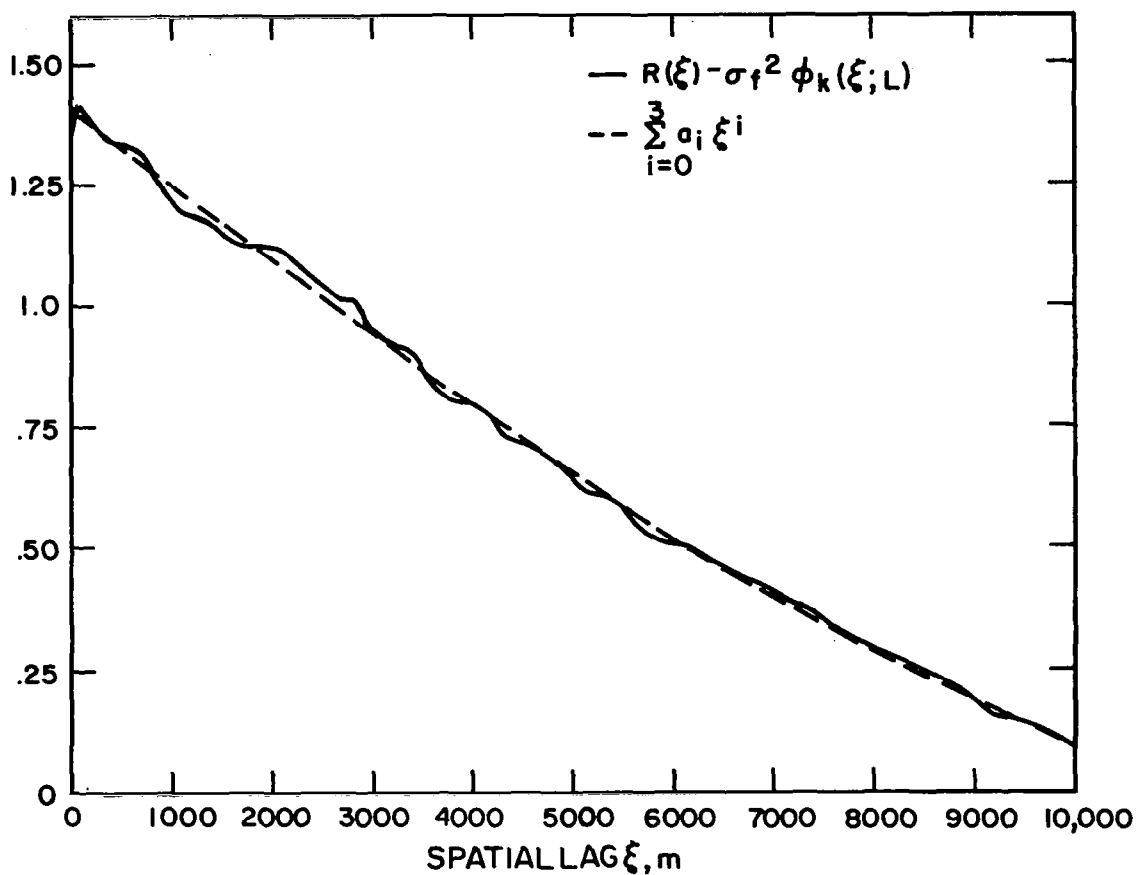


FIG. 20. COMPARISON OF AUTOCORRELATION FUNCTION $R(\xi)$ OF VERTICAL RECORD SHOWN IN FIG. 10 MINUS AUTOCORRELATION FUNCTION $\sigma_f^2 \phi_k(\xi; L)$ OF von KARMAN COMPONENT AND INTEGRAL LEAST-SQUARES THIRD-DEGREE POLYNOMIAL APPROXIMATION.

TABLE 3. CONSTRAINED LEAST-SQUARES ESTIMATION OF AUTOCORRELATION FUNCTION PARAMETERS FOR MOUNTAIN-WAVE VERTICAL RECORD

ξ_H m	m	σ_f^2 m ² /sec ²	L m	$\phi(0)$	a_0	a_1	a_2	a_3
295.6	0	.486	72.7	1.872	1.39			
798.1	1	.462	66.4	1.877	1.42	$-.157 \times 10^{-3}$		
798.1	2	.423	56.0	1.877	1.45	$-.350 \times 10^{-3}$	$.202 \times 10^{-5}$	
995.1	1	.446	62.2	1.874	1.43	$-.192 \times 10^{-3}$		
995.1	2	.474	69.8	1.876	1.40	$-.745 \times 10^{-4}$	$-.101 \times 10^{-6}$	
1202.0	1	.437	59.8	1.874	1.44	$-.210 \times 10^{-3}$		
1202.0	2	.470	68.4	1.876	1.41	$-.949 \times 10^{-4}$	$-.855 \times 10^{-7}$	
1399.0	1	.443	61.1	1.875	1.43	$-.201 \times 10^{-3}$		
1399.0	2	.443	61.4	1.875	1.43	$-.198 \times 10^{-3}$	$-.194 \times 10^{-5}$	
2502.5	1	.488	73.2	1.881	1.39	$-.148 \times 10^{-3}$		
2502.5	2	.429	57.6	1.874	1.45	$-.255 \times 10^{-3}$	$.409 \times 10^{-7}$	
2502.5	3	.422	55.9	1.874	1.45	$-.280 \times 10^{-3}$	$.632 \times 10^{-7}$	$-.565 \times 10^{-11}$
2502.5	4	.492	74.4	1.870	1.38	$.110 \times 10^{-3}$	$-.522 \times 10^{-6}$	$.323 \times 10^{-9}$
								$a_4 = -.616 \times 10^{-13}$

Exact value of $R(0)$ is $1.812 \text{ m}^2/\text{sec}^2$.

"oscillations" in Fig. 19 with those in Figs. 8 and 14(b), and finally (4) the fraction of the "power" of the record in the von Karman component is larger than in the previous two cases as is most easily seen by comparing Fig. 18 with Figs. 7 and 13.

In view of the above considerations, it is not surprising to see relatively less spread in the values of L shown in Table 3 than found in either of the previous two cases. Examination of the empirical autocorrelation function in Fig. 19 shows that the largest value of ξ_H for which we can expect a 2nd degree polynomial to represent well the autocorrelation function component from the "slow" turbulence component $w_s(t)$ is about $\xi_H = 1200$ m. Hence, the values of L and σ_f^2 from Table 3 that we choose to use for our von Karman curve in Fig. 16 were those given by Eq. (5.1) which were obtained using $\xi_H = 1202$ meters and $m = 2$. Values of L and σ_f^2 from the case $\xi_H = 798$ meters and $m = 1$ also would have provided reasonable choice ($L = 66.4$ meters and $\sigma_f^2 = 0.462 \text{ m}^2/\text{sec}^2$ as would have been the values obtained from the case $\xi_H = 995$ meters and $m = 2$ ($L = 69.8$ meters and $\sigma_f^2 = 0.474 \text{ m}^2/\text{sec}^2$). Very little discrepancy is found in the values of L from these three choices.

Maximum Likelihood Estimation of Integral Scale and Intensity of von Karman Component. The knee in the empirical spectrum shown in Fig. 16 is sufficiently well developed to apply the maximum likelihood method described in Sec. 3 of Ref. 1 to estimate the values of L and σ^2 of the von Karman component of the record. The likelihood equations were solved using spectrum sample points in the range from $k = 6.5 \times 10^{-4} \text{ m}^{-1}$ to $k = 3 \times 10^{-2} \text{ m}^{-1}$, which the reader may verify from Fig. 16 as the range dominated by the von Karman part of the spectrum. Values of L and σ^2 obtained using the maximum likelihood method were

$$L = 70.0 \text{ m}, \sigma^2 = 0.456 \text{ m}^2/\text{sec}^2. \quad (5.2)$$

Very little difference is observed among the various values of the integral scale cited above.

WAVENUMBER SPECTRAL DENSITY OF INSTANTANEOUS VARIANCE OF "FAST" COMPONENT OF VERTICAL RECORD FROM FLIGHT 30 RUN 8 (MOUNTAIN-WAVE CONDITIONS)

In Sec. 6.2 of Ref. 5, a method was developed for estimating the wavenumber spectrum of the instantaneous variance $\sigma_f^2(t)$ of the fast component $w_f(t)$ of a turbulence record, where time is translated to the spatial variable x using the relationship $x = Vt$ where V is the aircraft speed. The procedure used to compute the wavenumber spectrum of $\sigma_f^2(t)$ is as follows:

(1) High-pass filter the record to eliminate the low frequency component $w_s(t)$. In our computations with the vertical record from Flight 30 Run 8, we used first and second order digital Butterworth filters as described in Chapter 12 and Appendix C of Ref. 6.

(2) Find the wavenumber spectral density and autocorrelation function of the high-pass filtered record using the method described in Appendix B of Ref. 2.

(3) Square the high-pass filtered record and find the wavenumber spectral density and autocorrelation function of the squared high-pass filtered record using the method described in Appendix B of Ref. 2.

(4) Compute the wavenumber spectrum of $\sigma_f^2(t)$ using the formula:

$$\Phi_{\sigma_f^2}(k) = \{E[\sigma_f^2]\}^2 \left\{ \delta(k) + \int_{-M}^M p_0(\xi) \left(\frac{R_{w_h^2}(\xi) - [R_{w_h}(0)]^2 - 2[R_{w_h}(\xi)]^2}{[R_{w_h}(0)]^2 + 2[R_{w_h}(\xi)]^2} \right) e^{-i2\pi k \xi} d\xi \right\}, \quad (6.1)$$

where $p_0(\xi)$ is the Papoulis window function [7]

$$p_0(\xi) = \begin{cases} \frac{1}{\pi} \left| \sin \frac{\pi \xi}{M} \right| + \left(1 - \frac{|\xi|}{M} \right) \cos \frac{\pi \xi}{M}, & |\xi| \leq M \\ 0 & , |\xi| > M \end{cases} \quad (6.2)$$

see Appendix B of Ref. 2. Equation (6.1) is essentially the same as Eq. (6.49) of Ref. 5 except for translation of time t to distance x using $x = Vt$ where V is aircraft speed, and inclusion of the Papoulis window $p_0(\xi)$ to smooth the fluctuations normally associated with power spectral estimates of empirical waveforms. The aircraft speed for Flight 30 Run 8 was 197.05 m/sec and the value of M used was 10,089 m which corresponded to 1024 sample points.

The wavenumber spectra obtained by the above procedure are shown on log-log coordinates in Fig. 21 for three different high-pass filter cutoff frequencies (wavenumbers), where NS designates the number of filter sections used [6], and k_c designates the (3 dB) cutoff frequency (wavenumber) of the filters.

Except for statistical fluctuations associated with the finite length of the record, the wavenumber spectra $\Phi_{\sigma_f^2}(k)$ should be independent of k_c if our basic turbulence model

$$w_f(t) = \sigma_f^2(t) z(t) \quad (6.3)$$

is valid. From Fig. 21, we see that good agreement among the three spectra is obtained for values of $k < 10^{-3} \text{m}^{-1}$.

For values of $k > 10^{-3} \text{m}^{-1}$, the three spectra are shown vertically separated to avoid confusion among the three spectra. The agreement in the wavenumber range $k > 10^{-3} \text{m}^{-1}$ is poorer. However, in this wavenumber region, the locally stationary assumption leading to Eq. (6.34) of Ref. 5 is not valid — that is, a basic assumption in the derivation of Eq. (6.1) above is that the spectral content of σ_f^2 is negligible in the region $k > 10^{-3} \text{m}^{-1}$. Hence, the spectra in Fig. 21 in the region $k > 10^{-3} \text{m}^{-1}$ cannot be believed.

The same three spectra of Fig. 21 are plotted on linear coordinates in Fig. 22 for the wavenumber region $k < 10^{-3} \text{m}^{-1}$ where the spectra are believed to be valid.

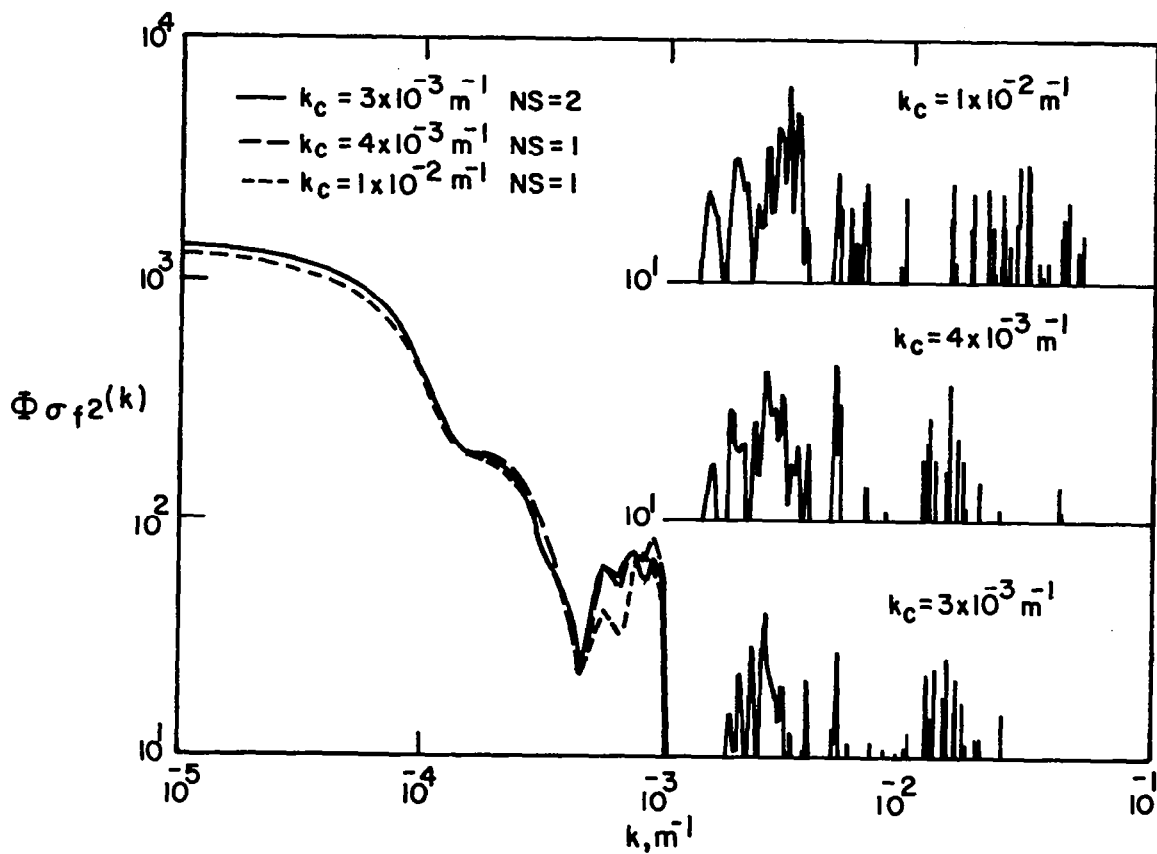


FIG. 21. WAVENUMBER SPECTRA OF INSTANTANEOUS VARIANCE $\sigma_f^2(t)$ OF "FAST" COMPONENT $w_f(t)$ OF VERTICAL RECORD SHOWN IN FIG. 10. THE THREE SPECTRA SHOWN WERE COMPUTED FROM THE SAME RECORD USING THREE DIFFERENT HIGH-PASS CUTOFF WAVENUMBERS.

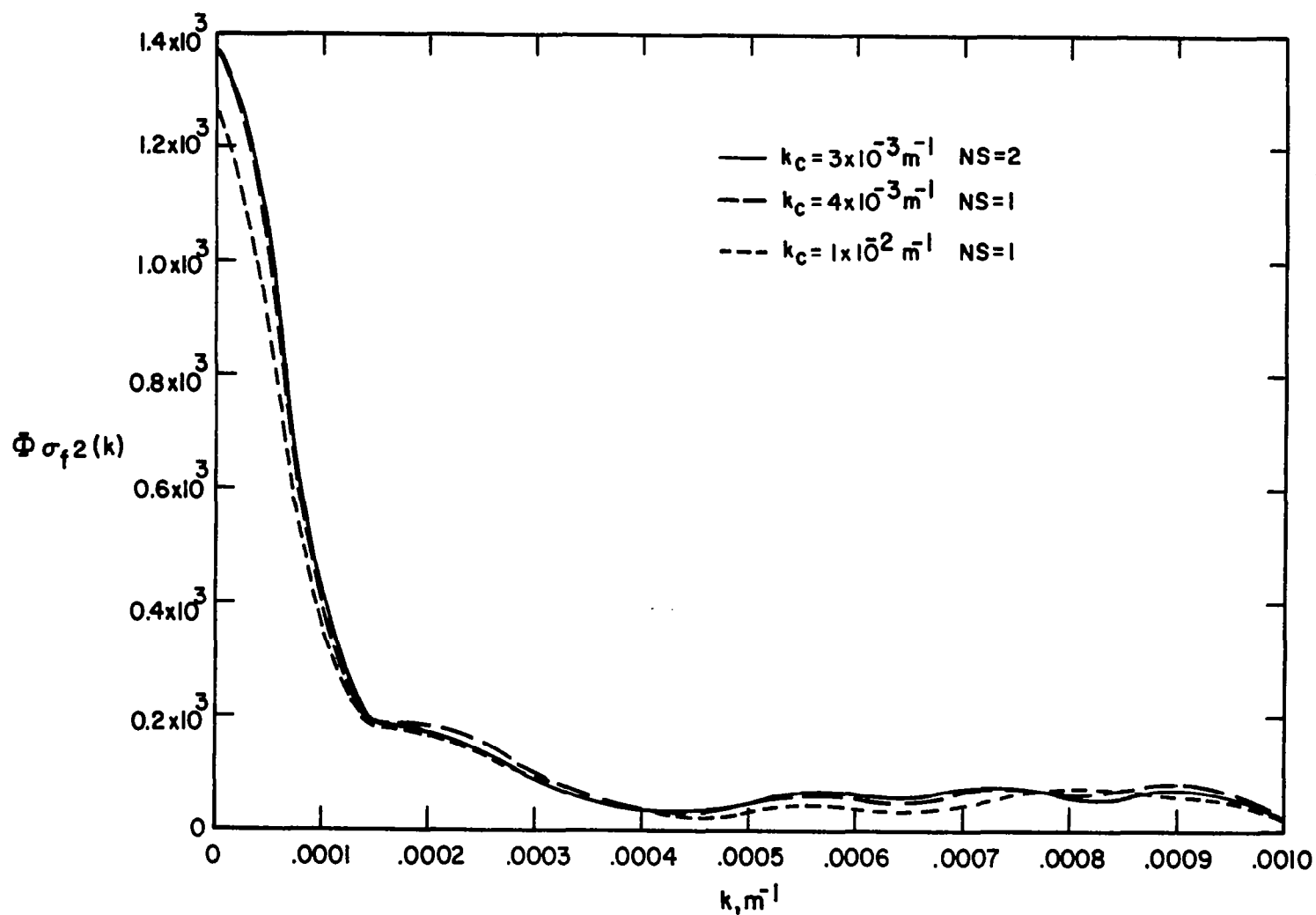


FIG. 22. WAVENUMBER SPECTRA OF INSTANTANEOUS VARIANCE $\sigma_f^2(t)$ OF "FAST" COMPONENT $w_f(t)$ OF VERTICAL RECORD SHOWN IN FIG. 10. THE THREE SPECTRA SHOWN WERE COMPUTED FROM THE SAME RECORD USING THREE DIFFERENT HIGH-PASS CUTOFF WAVENUMBERS.

The three high-pass filter cutoff wavenumbers used in obtaining the spectra of Figs. 21 and 22 may be compared with the spectrum of the vertical record $w(t)$ for Flight 30 Run 8 shown in Fig. 16.

PROBABILITY DENSITY FUNCTIONS OF INSTANTANEOUS VARIANCE $\sigma_f^2(t)$
AND SLOW TURBULENCE COMPONENT $w_s(t)$ OF VERTICAL RECORD FROM
FLIGHT 30 RUN 8 (MOUNTAIN-WAVE CONDITIONS)

Probability density of $\sigma_f^2(t)$. In Sec. 6.3 of Ref. 5, a method is developed for estimating the probability density of the instantaneous variance $\sigma_f^2(t)$ of the "fast component" $w_f(t)$ of our turbulence model. Using that method, the probability density of $\sigma_f^2(t)$ for the vertical record shown in Fig. 10 was estimated using moments of $\sigma_f^2(t)$ through orders 3, 4, 5, and 6. A high-pass two stage digital Butterworth filter [6] with cut-off wavenumber $k_c = 3 \times 10^{-3} \text{m}^{-1}$ was used in this procedure. The resulting probability densities obtained using Eq. (6.77) of Ref. 5 are plotted in Fig. 23 for the cases where moments through the third and sixth were used. The cases using moments through the fourth and fifth fell between the two curves shown in Fig. 23. All four approximations are sufficiently close to one another so that little practical significance can be ascribed to their differences.

Each of the four computed density functions had an integrable singularity at the origin. Thus, neither of the two curves shown in Fig. 23 has a finite value at $\sigma_f^2 = 0$. The reason for the integrable singularity at the origin is apparent when we examine the vertical record shown in Fig. 10. In the region between about 9 min 0 sec and 9 min 45 sec, the fast turbulence component $w_f(t)$ is virtually absent; hence, $\sigma_f^2(t)$ is very nearly zero during this time interval. This behavior is undoubtedly responsible for the singularity in $p(\sigma_f^2)$ at $\sigma_f^2 = 0$.

The probability density function of $\sigma_f^2(t)$ is useful for calculation of aircraft-response mean exceedance rates. See Eq. (4.8) of Ref. 5.

Probability density of $w_s(t)$. In Sec. 6.4 of Ref. 5, a method is developed for estimating the probability density of the "slow" turbulence component $w_s(t)$ from a turbulence time history. Using that method, the probability density of $w_s(t)$ for the vertical record shown in Fig. 10 was estimated using moments of $w_s(t)$ through the fourth. This technique also requires the moments through the fourth of the fast turbulence component $w_f(t)$ which were obtained using a high-pass two stage digital Butterworth filter [6] with cut-off wavenumber $k_c = 3 \times 10^{-3} \text{m}^{-1}$ as before. The probability

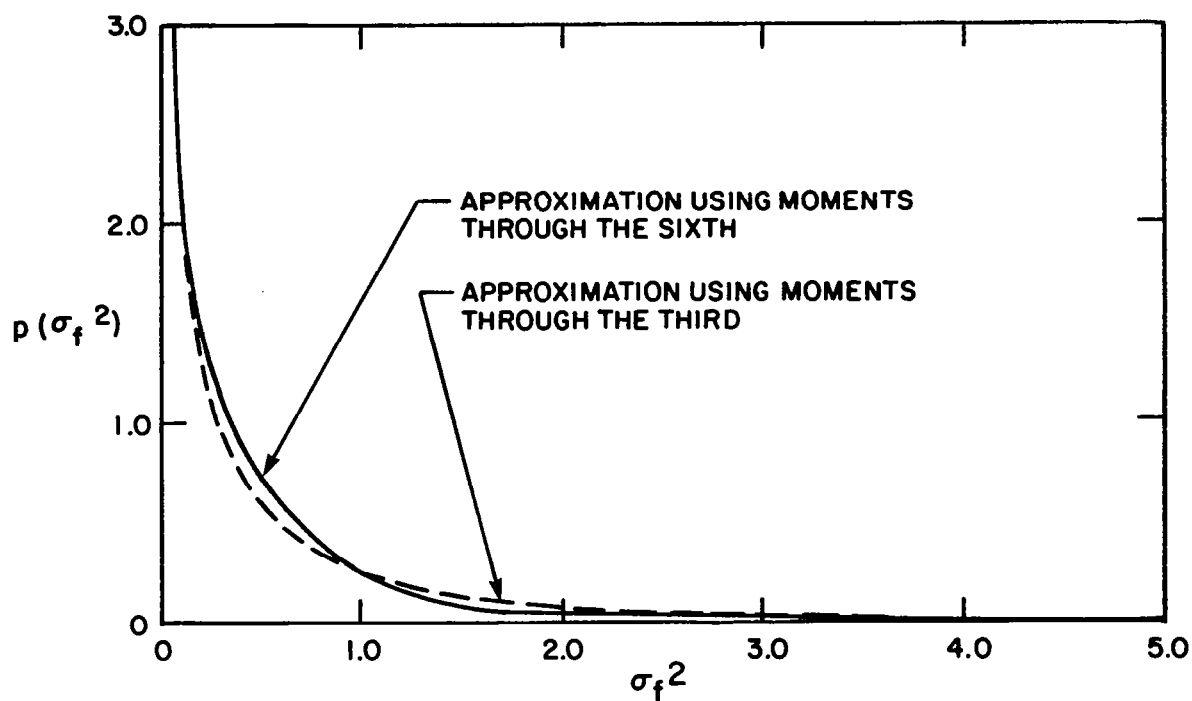


FIG. 23. PROBABILITY DENSITY FUNCTIONS OF INSTANTANEOUS VARIANCE $\sigma_f^2(t)$ OF THE "FAST" COMPONENT $w_f(t)$ OF VERTICAL RECORD SHOWN IN FIG. 10.

density of $w_s(t)$ is actually computed using Eq. (6.93) of Ref. 5, and is displayed in Fig. 24 for the vertical record shown in Fig. 10. Also shown there is the Gaussian density function with the same mean and variance.

The main purpose in estimating the probability density of $w_s(t)$ is to make a first check on the assumption that $w_s(t)$ is a stationary Gaussian process. Considering the obviously small number of statistical degrees-of-freedom in the component $w_s(t)$ of the vertical record shown in Fig. 10, we conclude from Fig. 24 that the deviation of the probability density of $w_s(t)$ from the Gaussian curve is not statistically significant.

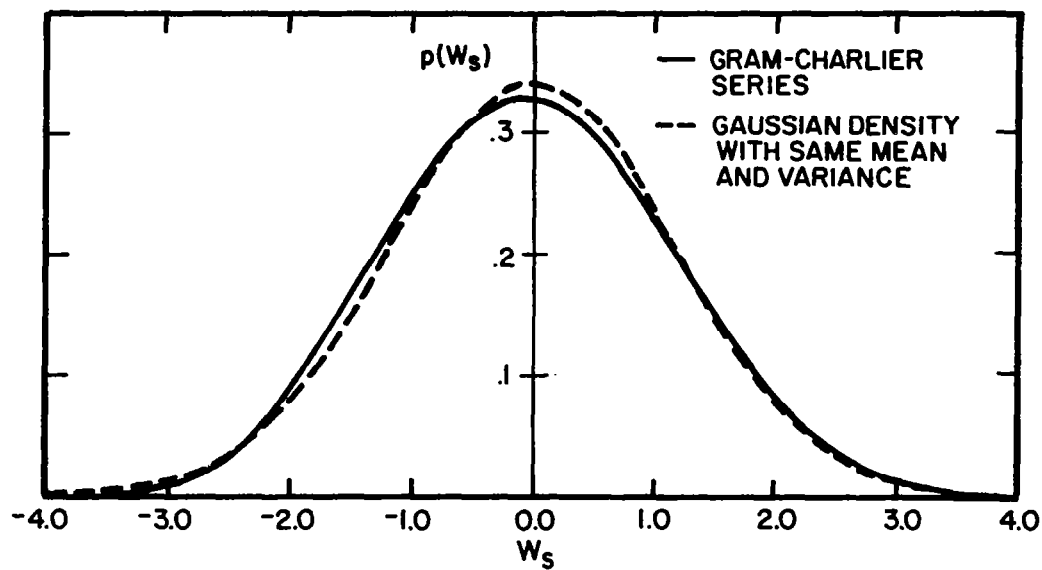


FIG. 24. ESTIMATE OF THE PROBABILITY DENSITY OF THE "SLOW" COMPONENT $w_s(t)$ OF THE VERTICAL RECORD SHOWN IN FIG. 10 USING THE GRAM-CHARLIER EXPANSION AND MOMENTS THROUGH THE FOURTH.

CONSTRAINED LEAST-SQUARES ESTIMATION OF AUTOCORRELATION
FUNCTION PARAMETERS OF LONGITUDINAL RECORD FROM FLIGHT 30
RUN 8 (MOUNTAIN-WAVE CONDITIONS) AND VERTICAL AND
LONGITUDINAL RECORDS FROM FLIGHT 32 RUN 4
(WIND-SHEAR CONDITIONS)

Here, we discuss collectively the results from three records, the longitudinal record from Flight 30 Run 8 (mountain-wave conditions) and the vertical and longitudinal records from Flight 32 Run 4 (wind-shear conditions). This completes all three records from Flight 30 Run 8 and Flight 32 Run 4.

The power spectral density computed from the longitudinal record of Fig. 10 is shown by the solid dots in Fig. 25, and the power spectral densities computed from the vertical and longitudinal records of Fig. 4 are shown by the solid dots in Figs. 26 and 27, respectively. All three of these spectra were computed by the method described in Appendix B of Ref. 2. The value of M used in the computation involving the mountain-wave record was 10,089 m as before, whereas the value of M used in the computations involving the wind-shear records was 9613.3 m. These values of M correspond in each case to 1024 sample points. Before computing these power spectra, the mean values of the records were computed and removed.

Von Karman longitudinal spectra are also plotted in Figs. 25 and 27 using solid lines, and the von Karman transverse spectrum of Eq. (2.1) is shown by the solid line in Fig. 26. The von Karman longitudinal power spectrum is described by

$$\Phi_{KL}(k) = 2\sigma^2 L \frac{1}{[1+70.78L^2k^2]^{5/6}} . \quad (8.1)$$

The values of σ^2 and L for the von Karman spectra are given in each of the three figures. These values of σ^2 and L were arrived at using the constrained least-squares estimation method described in Sec. 4 of Ref. 1, which postulates that the autocorrelation functions are of the form of Eq. (3.2), where $\sigma^2 \phi_K(\xi; L)$ is the appropriate form (transverse or longitudinal) of the von Karman autocorrelation function. In obtaining the von Karman spectra shown in Figs. 25 and 27, we used a value of $m = 3$, whereas $m = 4$ was used in

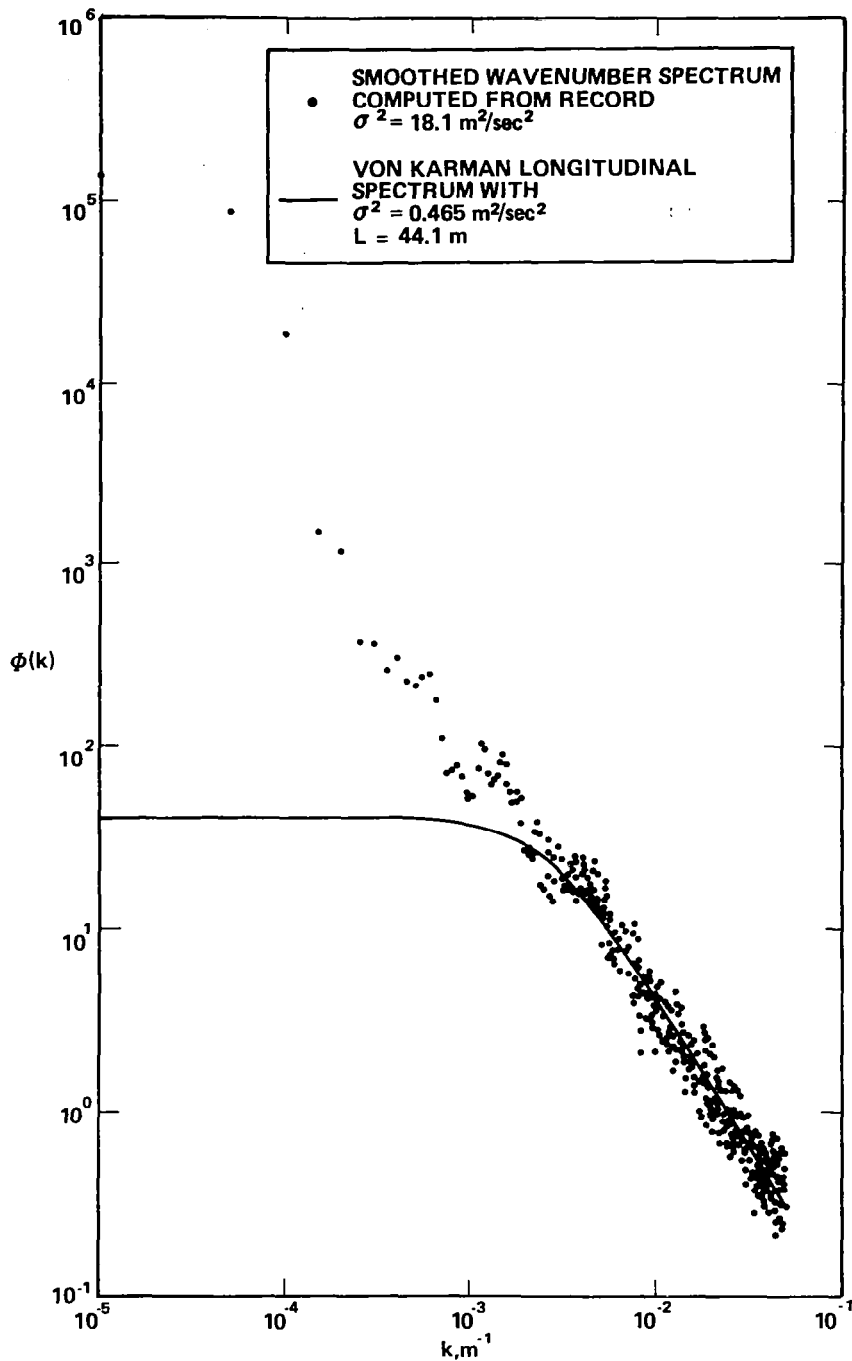


FIG. 25. COMPARISON OF SMOOTHED WAVENUMBER SPECTRUM COMPUTED FROM LONGITUDINAL RECORD SHOWN IN FIG. 10 AND VON KARMAN LONGITUDINAL SPECTRUM OBTAINED BY CONSTRAINED LEAST-SQUARES FIT TO THE (EMPIRICAL) AUTOCORRELATION FUNCTION. VON KARMAN SPECTRUM CHARACTERIZES "FAST" TURBULENCE COMPONENT ONLY.

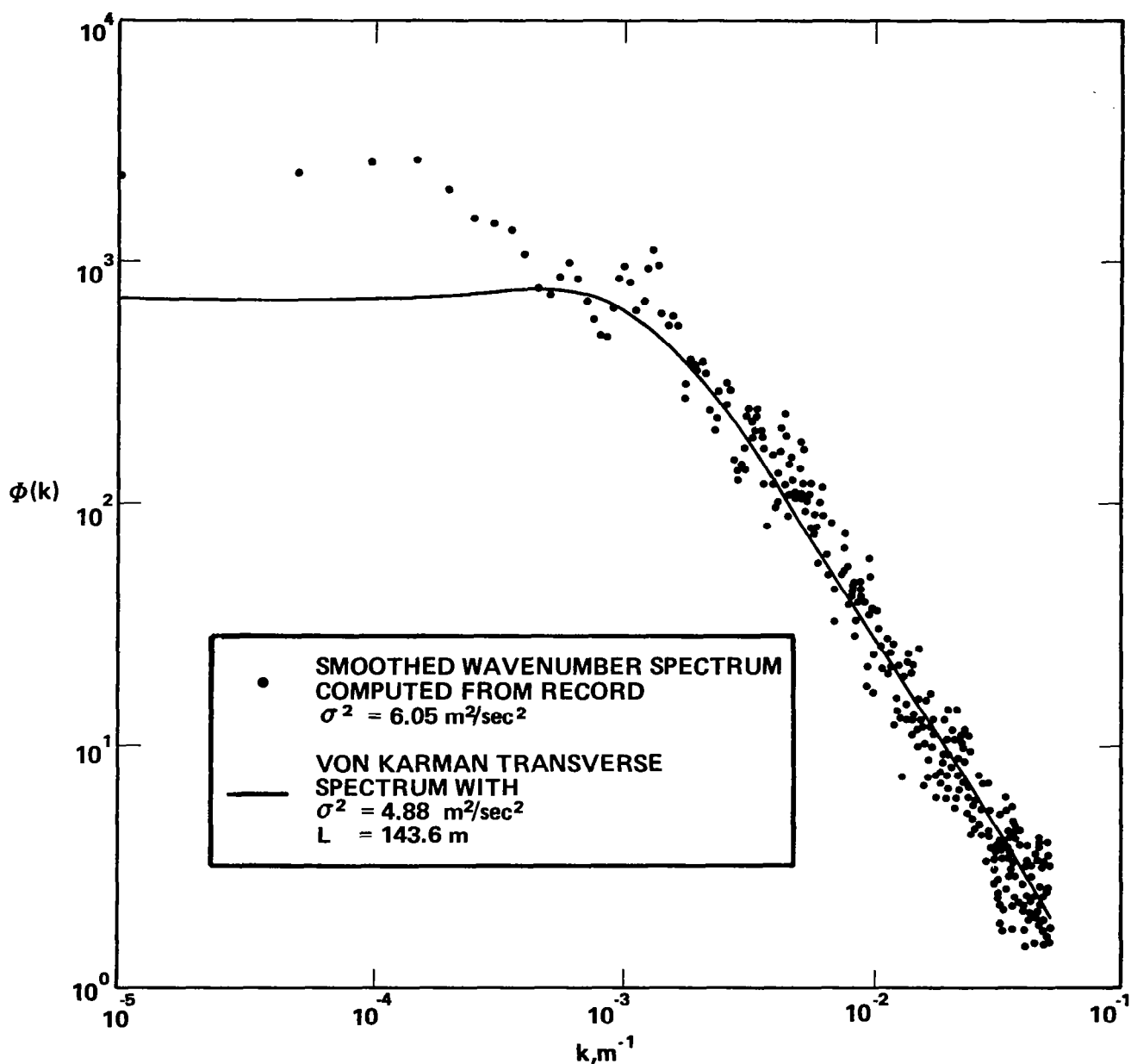


FIG. 26. COMPARISON OF SMOOTHED WAVENUMBER SPECTRUM COMPUTED FROM VERTICAL RECORD SHOWN IN FIG. 4 AND VON KARMAN TRANSVERSE SPECTRUM OBTAINED BY CONSTRAINED LEAST-SQUARES FIT TO THE (EMPIRICAL) AUTOCORRELATION FUNCTION. VON KARMAN SPECTRUM CHARACTERIZES "FAST" TURBULENCE COMPONENT ONLY.

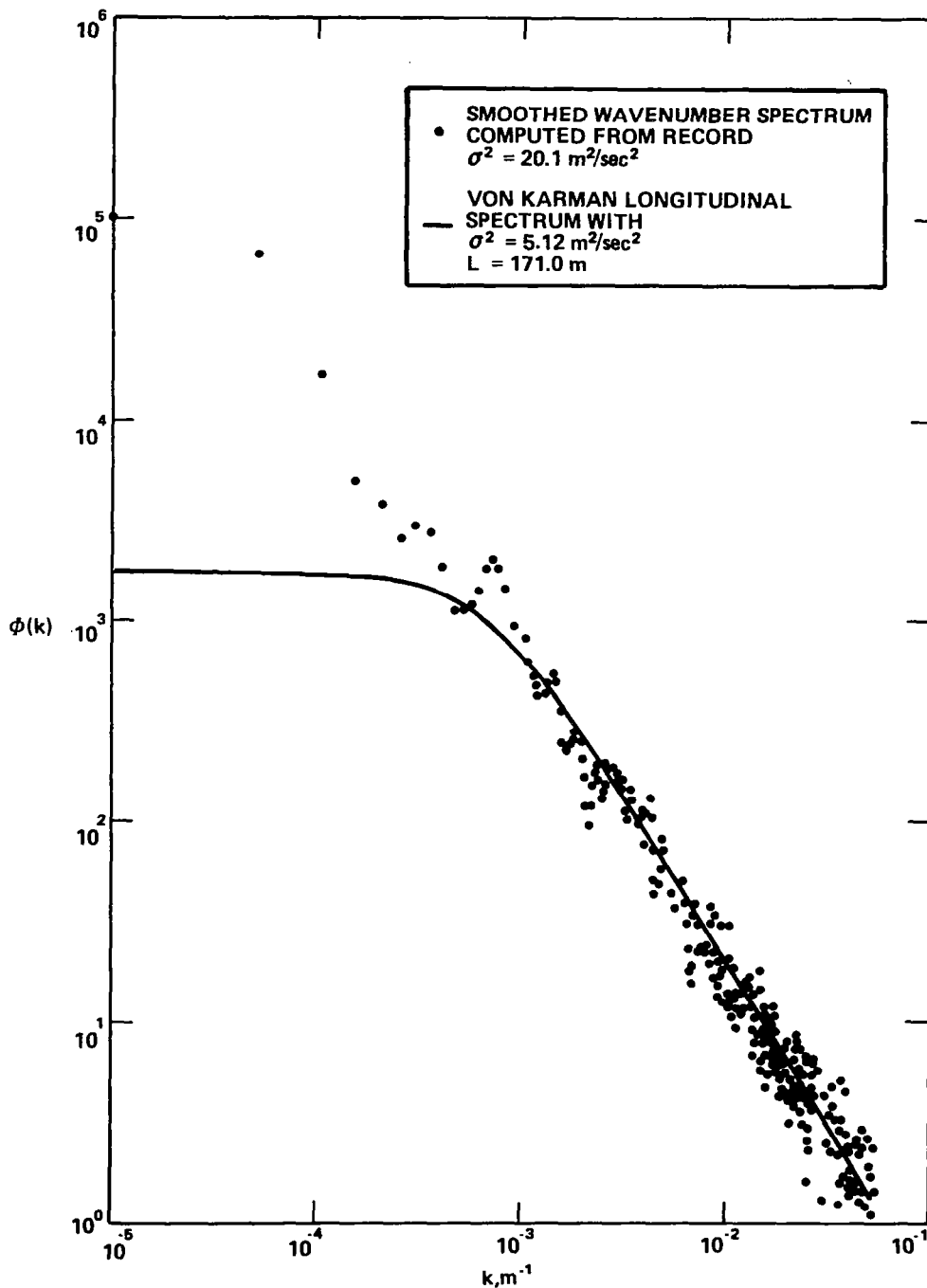


FIG. 27. COMPARISON OF SMOOTHED WAVENUMBER SPECTRUM COMPUTED FROM LONGITUDINAL RECORD SHOWN IN FIG. 4 AND VON KARMAN LONGITUDINAL SPECTRUM OBTAINED BY CONSTRAINED LEAST-SQUARES FIT TO THE (EMPIRICAL) AUTOCORRELATION FUNCTION. VON KARMAN SPECTRUM CHARACTERIZES "FAST" TURBULENCE COMPONENT ONLY.

obtaining the solid curve in Fig. 26. The values of ξ_H used in Figs. 25 through 27 were, respectively, $\xi_H = 1300.5$ m, $\xi_H = 7501.0$ m, and $\xi_H = 2497.0$ m. Values of k_ℓ and k_u used in every case were $k_\ell = 10^{-3} \text{m}^{-1}$ and $k_u = 4 \times 10^{-2} \text{m}^{-1}$. The equation of constraint used in the least-squares method for the mountain-wave longitudinal component is shown in Fig. 28. The equations of constraint for the other two components were not plotted.

Figures 29 through 31 show the autocorrelation functions of the three records under consideration. The autocorrelation function shown in Fig. 29 is dominated by the "slow" component, whereas the "slow" component contributes relatively little to the autocorrelation function shown in Fig. 30. The relative contribution of the "slow" component to the autocorrelation function shown in Fig. 31 is midway between the other two cases.

The autocorrelation function representation, Eq. (3.2), provided by the constrained least-squares estimation procedure of Ref. 1 is shown plotted in Figs. 32 through 34 for the three records under consideration. Each figure contains the results of several pairs of the parameters ξ_H and m . We may observe from these figures that the constrained least-squares fits $\phi(\xi)$ of Eq. (3.2) to the autocorrelation functions computed directly from the records are, in general, quite good.

Figures 35 through 37 show the autocorrelation function representations of the "slow" turbulence component

$$\phi_s(\xi) = \sum_{i=0}^m a_i \xi^i \quad (8.2)$$

obtained after removal of the von Karman component $\sigma_f^2 \phi_K(\xi; L)$ obtained in the constrained least-squares procedure. The value $m = 3$ was used in all three fits. Figures 35 through 37 were computed by first subtracting from the empirical autocorrelation functions $R(\xi)$ the von Karman autocorrelation functions with parameters σ^2 and L as given in Figs. 25 through 27 respectively. The resulting differences, $R(\xi) - \sigma_f^2 \phi_K(\xi, L)$ are shown by the solid lines in Figs. 35 through 37. We then computed an integral least-squares fit

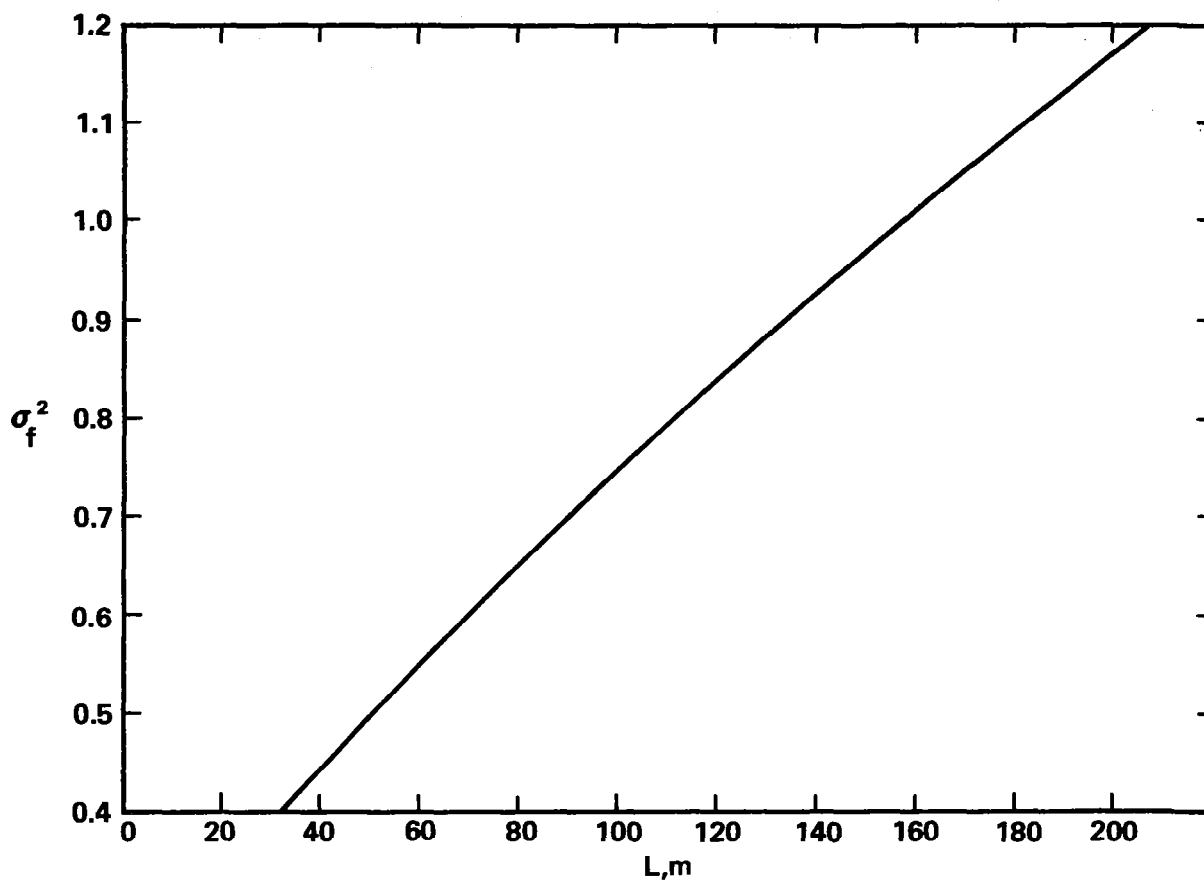


FIG. 28. CONSTRAINT BETWEEN σ_f^2 AND L FOR CONSTRAINED LEAST-SQUARES ESTIMATION PROCEDURE APPLIED TO LONGITUDINAL RECORD SHOWN IN FIG. 10.

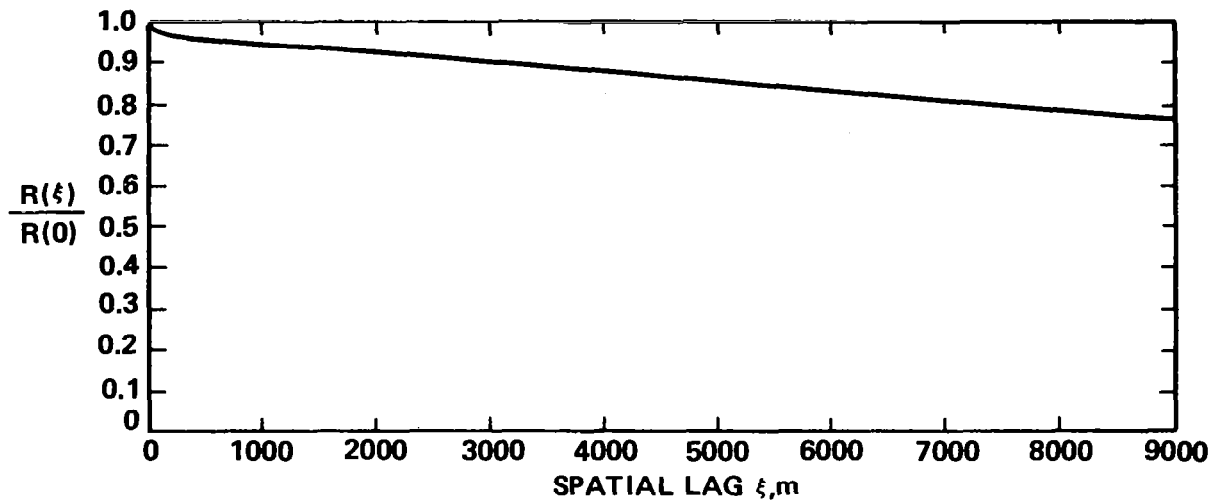


FIG. 29. AUTOCORRELATION FUNCTION OF LONGITUDINAL RECORD SHOWN IN FIG. 10 (MOUNTAIN-WAVE CONDITIONS). [FROM MAT PROJECT, NASA LANGLEY RESEARCH CENTER.]

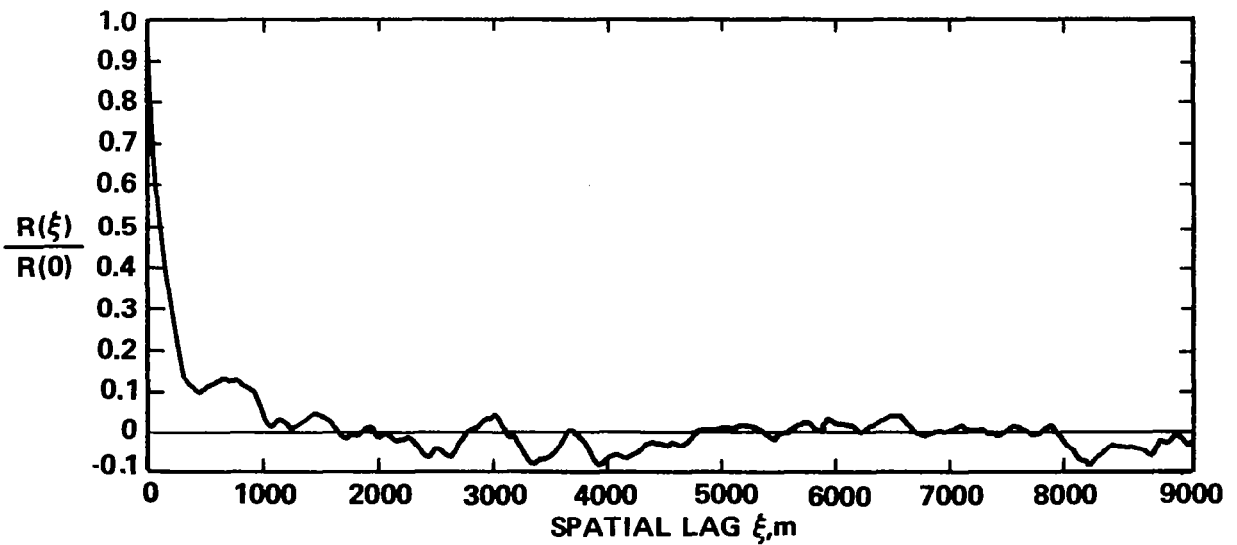


FIG. 30. AUTOCORRELATION FUNCTION OF VERTICAL RECORD SHOWN IN FIG. 4 (WIND-SHEAR CONDITIONS). [FROM MAT PROJECT, NASA LANGLEY RESEARCH CENTER.]

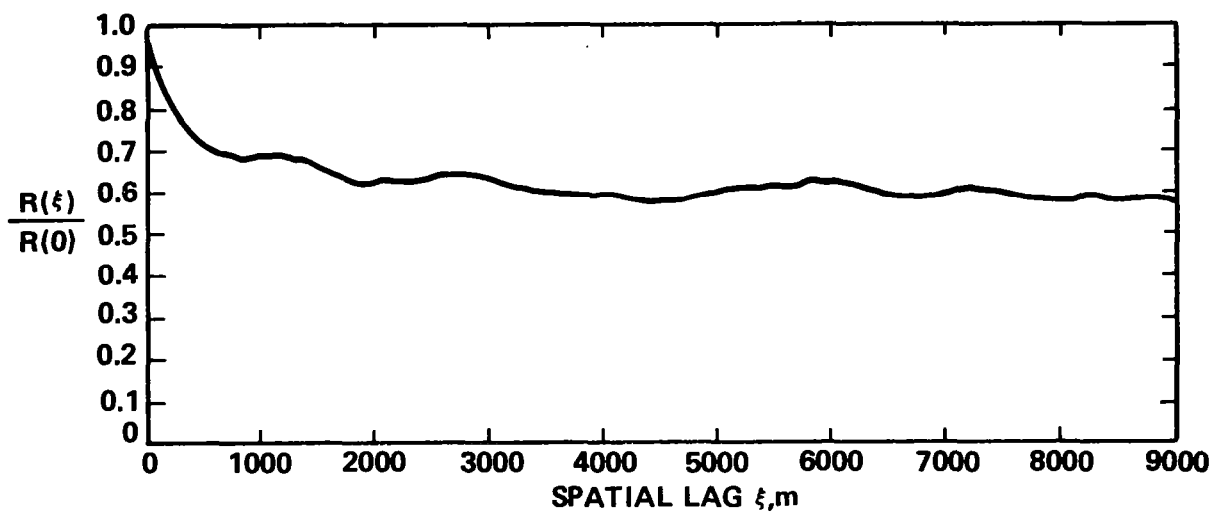


FIG. 31. AUTOCORRELATION FUNCTION OF LONGITUDINAL RECORD SHOWN IN FIG. 4 (WIND-SHEAR CONDITIONS). [FROM MAT PROJECT, NASA LANGLEY RESEARCH CENTER.]

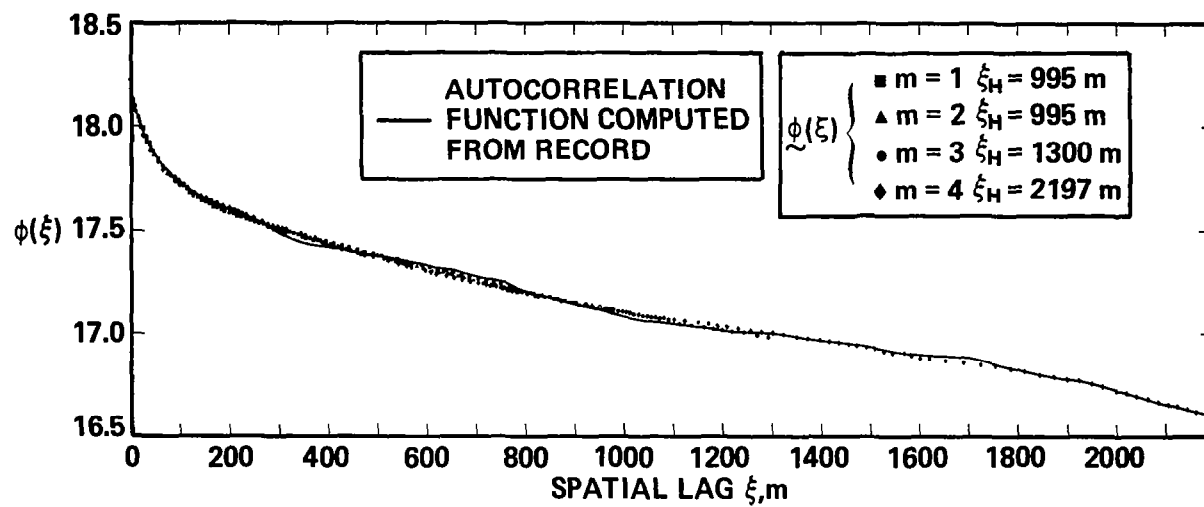


FIG. 32. COMPARISON OF AUTOCORRELATION FUNCTION COMPUTED FROM LONGITUDINAL RECORD SHOWN IN FIG. 10 AND CONSTRAINED LEAST-SQUARES FIT OF AUTOCORRELATION MODEL OF EQ. (3.2).

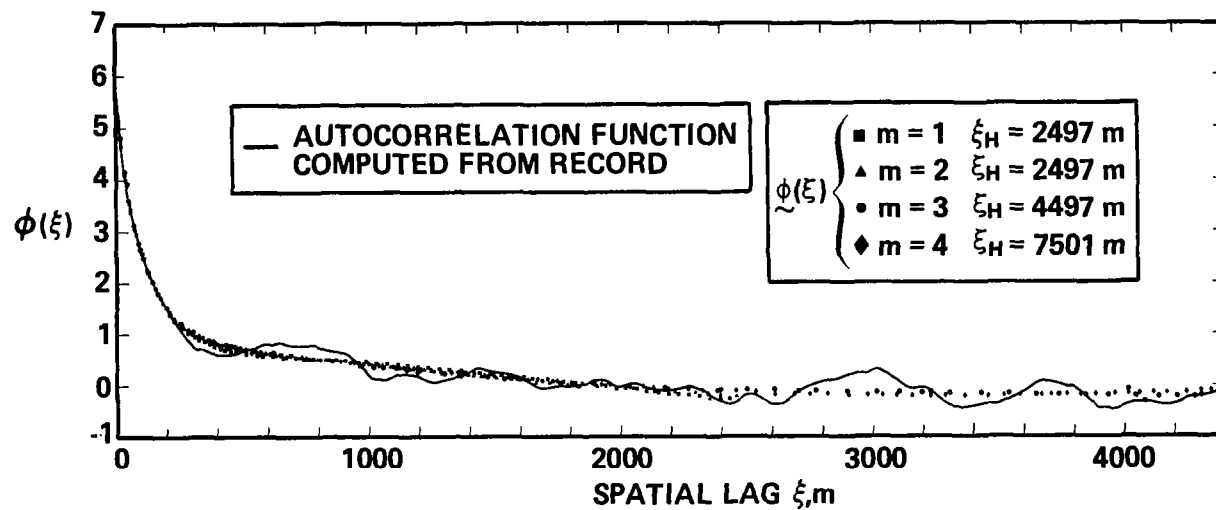


FIG. 33. COMPARISON OF AUTOCORRELATION FUNCTION COMPUTED FROM VERTICAL RECORD SHOWN IN FIG. 4 AND CONSTRAINED LEAST-SQUARES FIT OF AUTOCORRELATION MODEL OF EQ. (3.2).

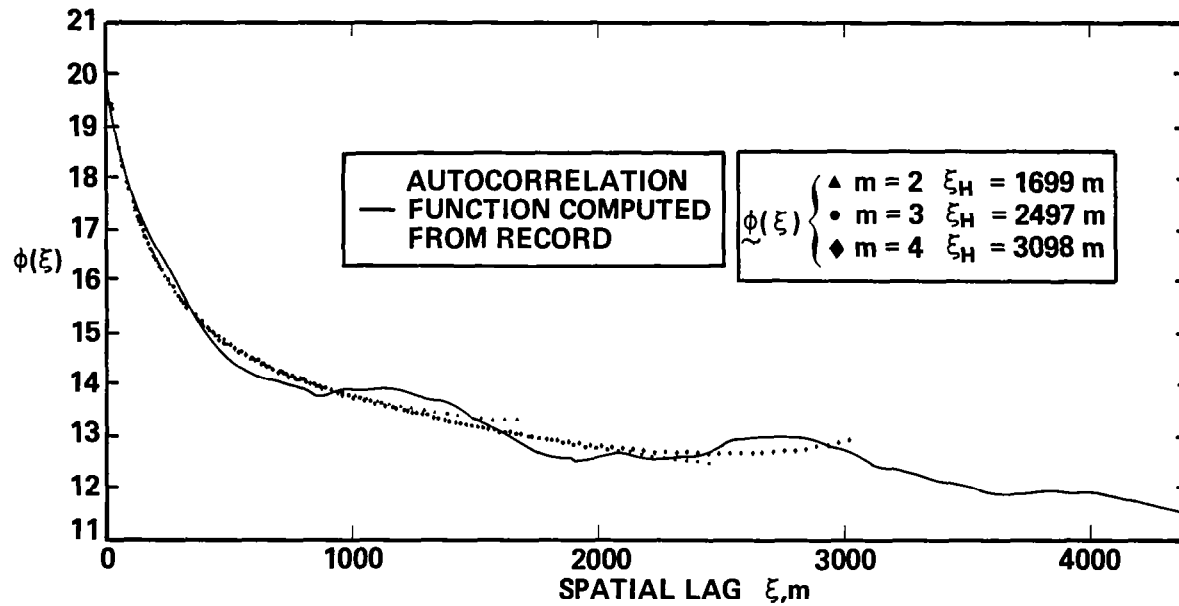


FIG. 34. COMPARISON OF AUTOCORRELATION FUNCTION COMPUTED FROM LONGITUDINAL RECORD SHOWN IN FIG. 4 AND CONSTRAINED LEAST-SQUARES FIT OF AUTOCORRELATION MODEL OF EQ. (3.2).

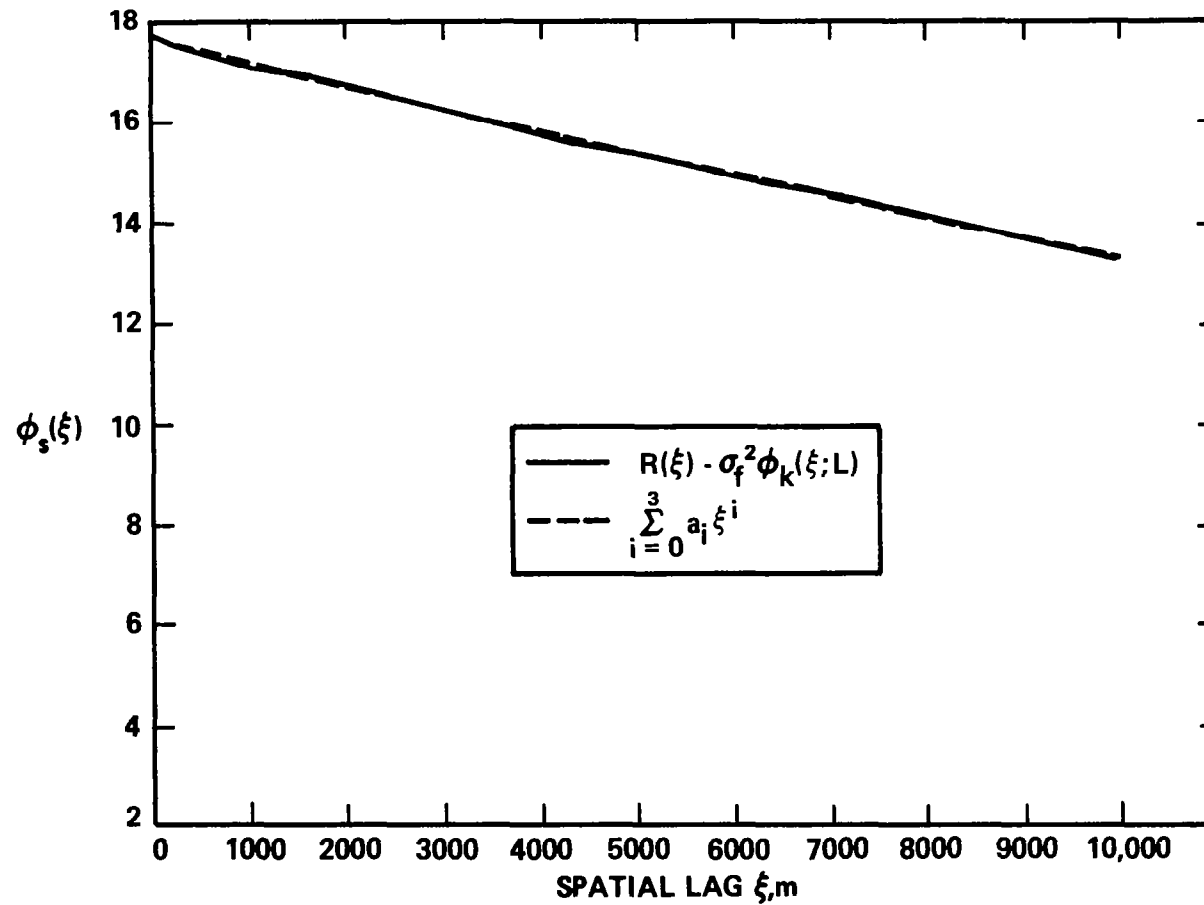


FIG. 35. COMPARISON OF AUTOCORRELATION FUNCTION $R(\xi)$ OF LONGITUDINAL RECORD SHOWN IN FIG. 10 MINUS AUTOCORRELATION FUNCTION $\sigma_f^2 \phi_k(\xi; L)$ OF VON KARMAN COMPONENT AND INTEGRAL LEAST-SQUARES THIRD-DEGREE POLYNOMIAL APPROXIMATION.

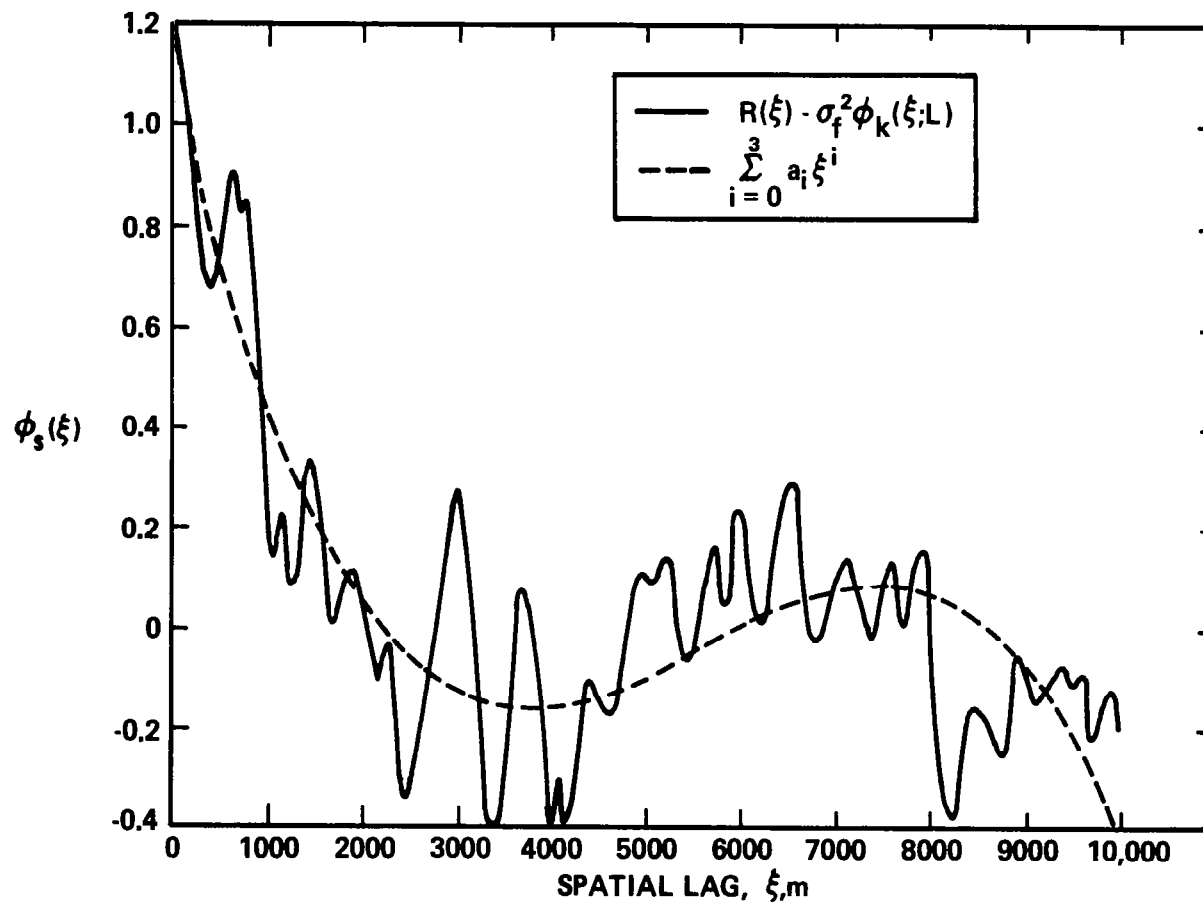


FIG. 36. COMPARISON OF AUTOCORRELATION FUNCTION $R(\xi)$ OF VERTICAL RECORD SHOWN IN FIG. 4 MINUS AUTOCORRELATION FUNCTION $\sigma_f^2 \phi_k(\xi; L)$ OF VON KARMAN COMPONENT AND INTEGRAL LEAST-SQUARES THIRD-DEGREE POLYNOMIAL APPROXIMATION.

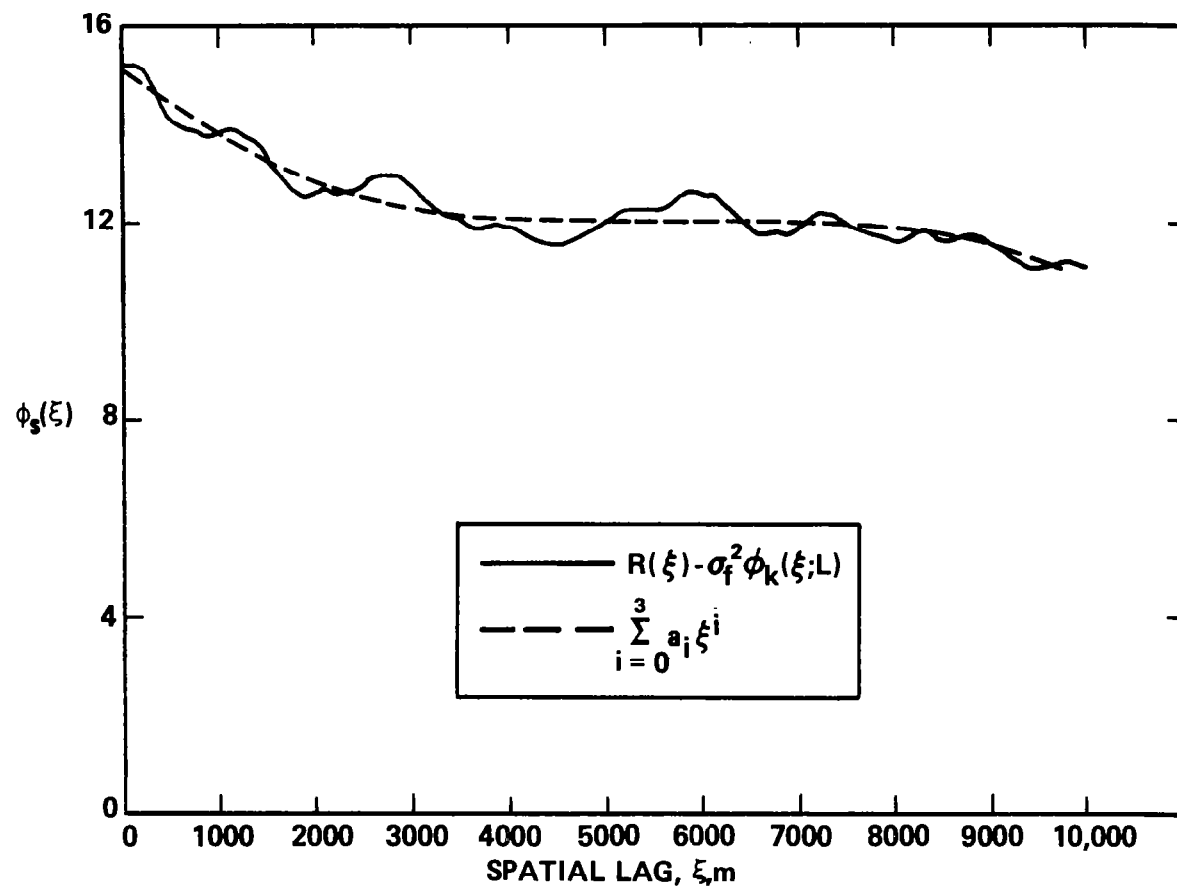


FIG. 37. COMPARISON OF AUTOCORRELATION FUNCTION $R(\xi)$ OF LONGITUDINAL RECORD SHOWN IN FIG. 4 MINUS AUTOCORRELATION FUNCTION $\sigma_f^2 \phi_k(\xi; L)$ OF VON KARMAN COMPONENT AND INTEGRAL LEAST-SQUARES THIRD-DEGREE POLYNOMIAL APPROXIMATION.

of Eq. (8.2) to the functions $R(\xi) - \sigma_f^2 \phi_K(\xi, L)$ over the lag intervals $0 < \xi < 10,000$ m using the third-degree polynomial ($m = 3$) of Eq. (8.2). These third-degree polynomial representations are shown by the dashed lines in Figs. 35 through 37.

Tables 4 through 6 show the values of σ_f^2 , L , $\phi(0)$, and a_0 through a_m for each combination of values of ξ_H and m used in the constrained least-squares fit of Eq. (3.2) to the autocorrelation functions of each of the three records under consideration. With the exception of the cases where $m = 1$, we observe relatively little spread in the values of the integral scale L of the von Karman component in each of Tables 4 through 6.

TABLE 4. CONSTRAINED LEAST-SQUARES ESTIMATION OF AUTOCORRELATION FUNCTION
PARAMETERS FOR MOUNTAIN-WAVE LONGITUDINAL RECORD

ξ_H m	m	σ_f^2 m^2/sec^2	L m	$\phi(0)$	a_0	a_1	a_2	a_3
995.1	1	.550	60.4	18.221	17.671	$-.582 \times 10^{-3}$		
995.1	2	.481	47.2	18.221	17.740	$-.829 \times 10^{-3}$	$.198 \times 10^{-6}$	
1300.5	2	.475	46.1	18.220	17.744	$-.837 \times 10^{-3}$	$.193 \times 10^{-6}$	
1300.5	3	.465	44.1	18.220	17.755	$-.889 \times 10^{-3}$	$.266 \times 10^{-6}$	$-.310 \times 10^{-10}$
2197.1	3	.414	34.5	18.223	17.809	-1.146×10^{-3}	$.589 \times 10^{-6}$	$-.143 \times 10^{-9}$
2197.1	4	.443	39.8	18.221	17.778	$-.977 \times 10^{-3}$	$.310 \times 10^{-6}$	$.303 \times 10^{-10}$
								$a_4 = -.365 \times 10^{-13}$

Exact value of $R(0)$ is $18.147 m^2/sec^2$

TABLE 5. CONSTRAINED LEAST-SQUARES ESTIMATION OF AUTOCORRELATION FUNCTION
PARAMETERS FOR WIND-SHEAR VERTICAL RECORD

ξ_H	m	σ_f^2	L	$\phi(0)$	a_0	a_1	a_2	a_3	a_4
m		m ² /sec ²	m						
2497.2	1	5.102	154.3	6.090	.989	-.518×10 ⁻³			
2497.2	2	4.858	142.6	6.066	1.208	-.909×10 ⁻³	.139×10 ⁻⁵		
4496.9	2	5.036	151.1	6.095	1.058	-.712×10 ⁻³	.986×10 ⁻⁵		
4496.9	3	4.710	135.6	6.070	1.360	-1.303×10 ⁻³	.388×10 ⁻⁶	-.40×10 ⁻¹⁰	
7501.0	3	4.775	138.6	6.066	1.291	-1.106×10 ⁻³	.258×10 ⁻⁶	-.18×10 ⁻¹⁰	
7501.0	4	4.881	143.6	6.073	1.192	-.909×10 ⁻³	.154×10 ⁻⁶	.25×10 ⁻¹¹	-.13×10 ⁻¹⁴
9003.1	4	4.709	135.5	6.065	1.356	-1.249 10 ⁻³	.340×10 ⁻⁶	-.34×10 ⁻¹⁰	.11×10 ⁻¹⁴

Exact value of R(0) is 6.053 m²/sec².

TABLE 6. CONSTRAINED LEAST-SQUARES ESTIMATION OF AUTOCORRELATION FUNCTION
PARAMETERS FOR WIND-SHEAR LONGITUDINAL RECORD.

ξ_H	m	σ_f^2	L	$\phi(0)$	a_0	a_1	a_2	a_3	a_4
m		m^2/sec^2	m						
1699.2	2	5.200	175.3	20.430	15.230	$-.197 \times 10^{-2}$	$.486 \times 10^{-6}$		
2102.9	3	4.262	128.7	20.705	16.443	$-.536 \times 10^{-2}$	$.370 \times 10^{-5}$	$-.10 \times 10^{-8}$	
2497.0	3	5.120	171.0	20.406	15.286	$-.190 \times 10^{-2}$	$.384 \times 10^{-6}$	$-.29 \times 10^{-10}$	
3098.0	4	4.965	163.2	20.482	15.517	$-.258 \times 10^{-2}$	$.108 \times 10^{-5}$	$-.35 \times 10^{-9}$	$.6 \times 10^{-13}$
3605.0	4	5.016	166.0	20.347	15.330	$-.149 \times 10^{-2}$	$-.599 \times 10^{-6}$	$.56 \times 10^{-9}$	-1.0×10^{-13}

Exact value of $R(0)$ is $20.072 \text{ m}^2/\text{sec}^2$.

METHODS FOR COMPUTATION OF THE INTEGRAL SCALE AND INTENSITY OF THE "SLOW" TURBULENCE COMPONENT

If $\phi(\xi)$ is the autocorrelation function of a turbulence record which is a function of the spatial lag variable ξ , and

$$\sigma^2 \equiv \phi(0) \quad (9.1)$$

is the mean-square value of the record, then the integral scale L is defined as [p. 43 of Ref. 8]

$$L \triangleq \frac{1}{\sigma^2} \int_0^{\infty} \phi(\xi) d\xi. \quad (9.2)$$

Let us define the wavenumber spectrum of the record as

$$\Phi(k) \triangleq \int_{-\infty}^{\infty} \phi(\xi) e^{-i2\pi k\xi} d\xi. \quad (9.3)$$

Hence,

$$\Phi(0) = \int_{-\infty}^{\infty} \phi(\xi) d\xi = 2 \int_0^{\infty} \phi(\xi) d\xi, \quad (9.4)$$

since $\phi(\xi)$ is an even function of ξ . From Eqs. (9.2) and (9.4), we may express the integral scale in terms of $\Phi(0)$ as

$$L = \frac{1}{2\sigma^2} \Phi(0). \quad (9.5)$$

In the case of the von Karman transverse wavenumber spectrum which we have chosen to characterize the "fast" turbulence component, we have instead of Eq. (9.5),

$$\Phi_{KT}(0) = \sigma_f^2 L_f , \quad (9.6)$$

according to Eq. (3.18) on p. 80 of Ref. 1, where in Eq. (9.6) we have used subscripts f to denote characteristics of the fast turbulence component. From Eqs. (9.5) and (9.6), we see that there is a factor of two between the definition L_f for the von Karman transverse integral scale and the definition given by Eq. (9.5). This is customary - e.g., see Ref. 9.

Let us denote the wavenumber spectra of $w(t)$, $w_s(t)$ and $w_f(t)$ in our turbulence model of Eq. (1.1) by $\Phi_w(k)$, $\Phi_s(k)$, and $\Phi_f(k)$. Then, from the assumed statistical independence of $\{w_s(t)\}$ and $\{w_f(t)\}$, it follows that

$$\Phi_w(k) = \Phi_s(k) + \Phi_f(k) ; \quad (9.7)$$

hence,

$$\Phi_s(0) = \Phi_w(0) - \Phi_f(0) . \quad (9.8)$$

Applying Eq. (9.5) to the slow turbulence component, and using subscripts s to denote "slow," we have

$$L_s = \frac{1}{2\sigma_s^2} \Phi_s(0) . \quad (9.9)$$

If we now recognize that the fast turbulence component whose spectrum is $\Phi_f(k)$ has been modeled by the von Karman transverse spectrum; hence, $\Phi_f(k) = \Phi_{KT}(k)$, we may combine Eqs. (9.6), (9.8), and (9.9) to yield the desired expression for the integral scale of the slow turbulence component:

$$L_s = \frac{1}{2\sigma_s^2} [\Phi_w(0) - \sigma_f^2 L_f] , \quad (9.10)$$

where all quantities on the right-hand side are easily estimated from methods discussed earlier. The mean-square value σ_s^2 of the slow component is obtained from the mean-square value of the original record and the mean-square value of the fast component by subtraction:

$$\sigma_s^2 = \sigma_w^2 - \sigma_f^2 \quad (9.11)$$

as may be seen from Eq. (9.7).

Initial evaluations of L_s for the first three records discussed in this report. The parameters required to evaluate L_s for the first three records with a "slow" component discussed earlier in this report are listed below in metric system units.

	Flt. 32 Run 4 (Wind Shear Lat.)		Flt. 30 Run 8 (Mt. Wave Lat.)		Flt. 30 Run 8 (Mt. Wave Vert.)	
$\Phi_w(0)$	$3.001 \times 10^5 \text{ m}^3/\text{sec}^2$		$2.187 \times 10^5 \text{ m}^3/\text{sec}^2$		$8.502 \times 10^3 \text{ m}^3/\text{sec}^2$	
σ_w^2	53.66	m^2/sec^2	29.23	m^2/sec^2	1.812	m^2/sec^2
σ_f^2	5.315	m^2/sec^2	0.684	m^2/sec^2	0.470	m^2/sec^2
L_f	265.5	m	128.9	m	68.4	m

Substitution of the above parameters into Eqs. (9.10) and (9.11) yields the following values for the integral scale L_s of the slow turbulence components of the three records:

$$\text{wind-shear lateral: } L_s = 3090 \text{ m} \quad (9.12a)$$

$$\text{mountain-wave lateral: } L_s = 3829 \text{ m} \quad (9.12b)$$

$$\text{mountain-wave vertical: } L_s = 3156 \text{ m.} \quad (9.12c)$$

Discussion. Rough checks may be made for each of the above values of L_s by comparing them with plots of the autocorrelation functions of the slow turbulence components shown in Figs. 9, 15, and 20 respectively. From the definition of L given by Eq. (9.2), it is evident from the above mentioned figures that each of the values of L_s given by Eq. (9.12) is too small. The reason that these computed values of L_s are too small may be seen from Eq. (9.10). To compute them, we used the values of the smoothed spectra $\Phi_w(0)$ evaluated at $k = 0$. The process of smoothing the wavenumber spectra to get improved statistical reliability reduced the values of the wavenumber spectra evaluated at

$k = 0$. Hence, we require an alternative method to estimate the integral scales L_S . Such an improved method has been developed in Appendix H of Ref. 1. This method is based directly on integration of the autocorrelation function representation of Eq. (8.2) after extrapolating its "tail" using a decaying exponential with continuous slope at $\xi = \xi_H$. The resulting formula for L_S derived in Appendix H of Ref. 1 [Eq. (H.14)] is

$$L_S = \frac{1}{a_0} \left[\sum_{j=0}^m \frac{a_j}{j+1} \xi_H^{j+1} - \frac{\left(\sum_{j=0}^m a_j \xi_H^j \right)^2}{\sum_{j=1}^m j a_j \xi_H^{j-1}} \right] \quad (9.13)$$

where a_0, \dots, a_m are the coefficients of the autocorrelation representation of Eq. (8.2), and $\xi_H = 10,000$ m is the upper limit of the lag parameter used in the integral least-squares fit.

Improved evaluations of L_S for the three records from Flight 30 Run 8 and the three records from Flight 32 Run 4. Using the autocorrelation function representation of Eq. (8.2) with $m = 3$ over the interval $0 \leq \xi \leq 10,000$ m that is displayed for the six records of interest by the dashed lines in Figs. 9, 15, 20, 35, 36, and 37, we have computed L_S using Eq. (9.13). The resulting values of integral scale are shown in Table 7.

TABLE 7. VALUES OF INTEGRAL SCALE L_S OF THE SLOW TURBULENCE COMPONENT COMPUTED USING EQ. (9.13).

wind-shear lateral	Fig. 9	$L_S = 8,511$ m
mountain-wave lateral	Fig. 15	$L_S = 37,094$ m
mountain-wave vertical	Fig. 20	$L_S = 4,997$ m
mountain-wave longitudinal	Fig. 35	$L_S = 32,384$ m
wind-shear vertical	Fig. 36	$L_S = 825$ m*
wind-shear longitudinal	Fig. 37	$L_S = 17,052$ m

*The value of $L_S = 825$ m for the wind-shear vertical case is completely unreliable because of the behavior of the autocorrelation function representation of Eq. (8.2) shown in Fig. 36.

From examination of the dashed autocorrelation function representations in each of the above cited figures, we can conclude that the method of Eq. (9.13) should provide reasonable values of L_s for all records except the wind-shear vertical record whose autocorrelation function is shown in Fig. 36. This lack of reliability is noted in Table 7.

APPENDIX A

INTRODUCTION TO COMPUTER PROGRAMS

The Appendices to this report document the following computer programs developed to characterize nonGaussian atmospheric turbulence records relevant to aircraft response calculations:

- a. maximum likelihood estimation of the integral scale and variance of von Karman turbulence,
- b. constrained least-squares estimation of turbulence autocorrelation function parameters,
- c. power spectral density of the instantaneous variance $\sigma_f^2(t)$,
- d. probability density estimation of the instantaneous variance $\sigma_f^2(t)$ and the "slow" turbulence component $w_s(t)$.

The turbulence models used to develop these programs are described in BBN Report 4319, Characterization, Parameter Estimation, and Aircraft Response Statistics of Atmospheric Turbulence (Ref. 1).

These Appendices explain the program usage, program inputs and outputs, and give typical teletype printouts. Included are source program listings and printouts of typical output data files. The programs were written in FORTRAN IV. The notation used in this report is consistent with that of the above cited reference.

The source program and subroutine listings are contained in Appendix F in alphabetic order by program name. The source programs and subroutines contain comment statements to aid the user in following the flow of the computations.

Table 8 contains a summary of the purpose of each program, subroutines required, and the form of the input/output of each program.

TABLE 8. ATMOSPHERIC TURBULENCE MODELING PROGRAMS.

Appendix	Main Program	Subroutines	Purpose	Inputs	Outputs*
B.1.1	ATURB2	CFFT SIMP	compute $\phi_{\ell}(k)$, $R_{\ell}(\xi)$, $\phi_{\rho}(k)$	turbulence samples, TTY inputs	(PHILK): $\phi_{\ell}(k)$ (AUTO): $R_{\ell}(\xi)$ (DSPS): $\phi_{\rho}(k)$
B.1.2	PART2	AK AKDAT GAM PARAB SIMQ	computes $E(L)$, $LG(k_1, L)$, $\phi_K(\xi)$ and $\phi_K(k)$	data file (PHILK) and TTY inputs	(LG): $LG(k_1)$ (PHIXI): $\phi_K(\xi)$ (PHIK): $\phi_K(k)$
C.1.1	ATURB3	CFFT SIMP	computes proven spectrum auto-correlation for nonGaussian turbulence samples - similar to ATURB2	turbulence	(PHILK): $\phi_{\ell}(k)$ (AUTO): $R(\xi)$ (DSPS): $\phi_{\rho}(k)$
C.1.2	PART5	--	computes $\sigma^2(L_j)$, $j=1, \dots, 15$	(PHILK) TTY inputs	TTY: $\sigma^2(L_j)$
C.1.3	FINAL	AK1 AKDAT ANRP1 FnDECT DGELG GAM PAR1&2 SET SIMP2 SIMQ TRAP3&6	computes $\phi(\xi)$	TTY inputs	(ITM2): $\phi(\xi)$ (ITM2L): $\sigma_f^2 L \bar{\phi}_K(kL)$
D.1.1	ATURB4	CFFT HPDES SIMP	Filter data, produce power spectrum & autocorrelation fn	turbulence samples & TTY inputs	(PHILK): $\phi_{\ell}(k)$ (AUTO): $R_{w_h}(\xi)$ TTY inputs
D.1.1 (cont.)	ATUR4A	CFFT1 HPDES SIMP	Filter data, square, produce power spectrum and auto-correlation	same as above	(FPSD2): $\phi_{\ell}(k)$ (AUTF2): $R_{w_h}(\xi)$
D.1.2	ITEM3	CFFT1	compute $R_{\sigma_f^2}(\xi)$, $\phi_{\sigma_f^2}(k)$	(AUTO) (AUTF2) & TTY inputs	(RSIGF): $R_{\sigma_f^2}(\xi)$ (PHIF): $\phi_{\sigma_f^2}(k)$
E.1.1	MOMENT	BIN	computes $\alpha_w^{[k]}$, $\alpha_{w_h}^{[k]}$, $\alpha_h^{[k]}$, and $\alpha_{\sigma_f^2}^{[k]}$ $k=1, \dots, 8$	turbulence samples & TTY input	TTY: moments listed to left
E.1.2	GDIST6	GAM	probability density $\rho_{\sigma_f^2}$	TTY input	TTY output
E.1.3	ITEM4	FAC1	probability density ρ_{w_s}	TTY input	(PROB): ρ_{w_s} & TTY output

*Data files in parentheses.

APPENDIX B

MAXIMUM LIKELIHOOD ESTIMATION OF THE INTEGRAL SCALE AND VARIANCE OF von KARMAN TURBULENCE

Two main programs, ATURB2.F4 and PART2.F4, are used in computing the integral scale and variance of von Karman turbulence. The development of the maximum likelihood estimation technique is discussed in Sec. 3 of Ref. 1 and Appendix F of Ref. 1. The first program, ATURB2, computes the two sided power spectrum, S_j , the autocorrelation function, $R_0(\xi)$, and the two sided smoothed power spectrum, $\Phi_p(k)$, of the turbulence record, $w(t)$, being processed. The second program, PART2, calculates the integral scale of the stationary turbulence sample and its variance.

A fast Fourier transform subroutine is utilized in computing the power spectra and autocorrelation function. The positive frequency domain values of the spectra and the values of the autocorrelation function are stored in three separate output data files. The data file PHILK, containing S_j values, is used by the program PART2. In addition to computing the integral scale and variance, PART2 computes the von Karman autocorrelation function $\phi_K(\xi)$ and von Karman spectrum $\phi_{KT}(k)$ using values of σ^2 and length scale L determined by the program.

Program Outlines and Usage

Program ATURB2.F4: Computes the two sided power spectrum, S_j , of turbulence data, the autocorrelation function ϕ , and smoothed power spectrum Φ .

a. Subroutines

- i. CFFT - fast Fourier transform routine
- ii. SIMP - integration by Simpson's rule.

b. Inputs* [from the teletype (TTY) unless otherwise noted]

- i. speed of craft in m/sec - $(V)^\dagger$
- ii. number of points in Fourier transform (NPTS) and power of two of that number (MPWRN)
- iii. number of points in turbulence record (NOPTS) and sampling rate of data (SRATE)

*Unless otherwise noted a "G" format is used for numerical inputs from the teletype.

[†]Alphanumerics in parenthesis represents equivalent variable name in source program.

- iv. number of points for smoothed power spectrum (MPTS) and power of two of that number (MPWRM)
- v. name of data file (A5 format) containing turbulence values $w(t)$ and number of points in the array (NXRAY; input file has 0 to NXRAY-1 points divided into 4 columns with a format of 4 (E15.7); input values in units of ft/sec - program converts values to m/sec); this data file is read by the program.
- vi. answer yes (Y) if program check on integration is desired (value printed on TTY should equal value printed in data file) or no (N) if program check is not desired.

c. Outputs (to various data files or to the TTY)

- i. data file, PHILK, containing positive frequency (k) domain values of power spectrum S_j , equation 3.27, Ref. 3.
- ii. data file, AUTO, containing autocorrelation values
- iii. data file, DSPS, containing values of the smoothed power spectrum

d. Example

A typical teletype printout of the execution of this program is shown in Fig. B.1. The user supplied information discussed in b is underlined in the example. The first page of each output file mentioned is shown in Figs. B.2 through B.4.

Program PART2.F4: Computes the integral scale, L , of a stationary turbulence sample and its variance (Eq. 3.25, Ref. 1).

a. Subroutines:

- i. GAM - computes gamma function Γ
- ii. AK and AKDAT - compute modified Bessel functions of fractional orders $1/3$ and $2/3$, AK uses data stored in AKDAT
- iii. SIMQ - simultaneous linear equation routine from the IBM Scientific Subroutine Package
- iv. PARAB - curve fitting routine

```

@LOADER
+ATURB2,CFFT,SIMP$

ATURB2 36K CORE, 212 WORDS FREE
LOADER USED 39+5K CORE

EXIT.
^C
@SAVE (CORE FROM) 20 (TO) 777777 (ON)      ATURB2.SAV [NEW FILE]
@ST

```

```

INPUT SPEED OF CRAFT (M/SEC)
187.76

```

```

INPUT TOTAL NO. OF POINTS TO BE USED IN 2L M. OF DATA
AND POWER OF TWO OF THAT NO.
16384
14

```

```

INPUT NO. OF POINTS OF W(X) TO BE READ
AND SAMPLING RATE OF DATA 14592.,.05

```

```

INPUT VALUE OF MPTS
AND POWER OF TWO OF THAT NO.
1024
10

```

```

INPUT DATA FILE NAME THAT CONTAINS SAMPLES
OF W(X) AND NO. POINTS NXRAY
NASA1
14592

```

```

PERFORM INTEGRATION CHECK (Y OR N)   Y

```

```

                                program output should match
                                value of  $\sigma^2$  of data files
INTEGRAL OF PHI OF L (K) = 0.4802E+02
0.3594628E-02  0.0000000  1478369.  0.0000000 } intermediate
848732.1      0.0000000  244530.9  0.0000000 } output, useful
118152.3      26228.80   0.0000000 } for debugging
0.4658256                                only
0.7216021      0.7061144
848688.2      244518.3   118146.2

```

```

CPU TIME: 4:0.23      ELAPSED TIME: 10:5.85
NO EXECUTION ERRORS DETECTED

```

```

EXIT.
^C

```

FIG. B.1. TTY PRINTOUT FOR RUNNING PROGRAM ATURB2.

DATA FILE CREATED BY PROGRAM ATURB2

POWER SPECTRUM OF PHI OF L(K)

DATA TAKEN FROM FILE VERT

32768 DATA POINTS WERE USED IN 2L = 210894.8460 METER
1024 DATA POINTS WERE USED IN M = 6590.4639 METER

9968 ZEROS WERE ADDED TO DATA

MEAN VALUE OF W(X) = 0.83463E-01 M/SEC
MEAN SQ. VALUE = 0.13305E+01 (M/SEC)**2

<W OF L(X)**2> = 1.3305

PRINTOUT OF THE VALUES OF THE POWER SPECTRUM

K	PS VALUE	K CONTD	PS VALUE CONTD
0.000000	0.1419E-09	0.038844	0.5791E+00
0.000005	0.1622E+04	0.038849	0.1750E+01
0.000009	0.6052E+03	0.038853	0.4770E+00
0.000014	0.2032E+02	0.038858	0.2895E+00
0.000019	0.4472E+02	0.038863	0.1043E-01
0.000024	0.5508E+03	0.038868	0.7540E+00
0.000028	0.1108E+03	0.038872	0.4492E-01
0.000033	0.7429E+02	0.038877	0.1940E+00
0.000038	0.5229E+03	0.038882	0.4078E+00
0.000043	0.2987E+03	0.038887	0.6840E-01
0.000047	0.4791E+03	0.038891	0.8488E+00
0.000052	0.3334E+03	0.038896	0.1444E+00
0.000057	0.1399E+03	0.038901	0.2073E-01
0.000062	0.1206E+03	0.038906	0.1683E+00
0.000066	0.1233E+03	0.038910	0.9916E+00
0.000071	0.1368E+03	0.038915	0.8517E+00
0.000076	0.5743E+03	0.038920	0.1654E-02
0.000081	0.9769E+02	0.038925	0.2768E+00
0.000085	0.3835E+03	0.038929	0.8511E+00
0.000090	0.8038E+02	0.038934	0.6148E+00
0.000095	0.5548E+03	0.038939	0.7542E+00
0.000100	0.7102E+03	0.038944	0.1934E+01
0.000104	0.3008E+03	0.038948	0.8529E-01
0.000109	0.3210E+03	0.038953	0.5054E+00
0.000114	0.7278E+03	0.038958	0.1431E+00
0.000119	0.1500E+03	0.038963	0.4173E+00
0.000123	0.1513E+04	0.038967	0.2629E+01

FIG. B.2. OUTPUT DATA FILE PHILK.

↑
 AUTO renamed to AUTWSV

DATA FILE CREATED BY PROGRAM ATURB2

AUTOCORRELATION OF STATIONARY SAMPLE
DATA TAKEN FROM FILE NASV

16384 DATA POINTS WERE USED IN 2L = 153812.9920 METER
 1024 DATA POINTS WERE USED IN M = 9613.3120 METER

1796 ZEROS WERE ADDED TO DATA

MEAN VALUE OF W(X) = -0.43757E-01 M/SEC
 MEAN SQ. VALUE = 0.60534E+01 (M/SEC)**2

<W OF L(X)**2> = 6.0533

TRUNCATION POINT WAS 13000.00 METERS
 WHICH CONTAINS 1384 POINTS

PRINTOUT OF THE VALUES OF THE AUTOCORRELATION

X	RL	RL/R0
0.0000000	6.053274	1.0000000
9.388000	5.531511	0.9138248
18.77600	5.046111	0.8336167
28.16400	4.635600	0.7658203
37.55200	4.277285	0.7066268
46.94000	3.941774	0.6511804
56.32800	3.612217	0.5967377
65.71600	3.323236	0.5489980
75.10400	3.083437	0.5093834
84.49200	2.885883	0.4767475
93.88000	2.686956	0.4438848
103.2680	2.529167	0.4178180
112.6560	2.397351	0.3960421
122.0440	2.274019	0.3756676
131.4320	2.177932	0.3597940
140.8200	2.081459	0.3438567
150.2080	1.979035	0.3269362
159.5960	1.874740	0.3097068
168.9840	1.796849	0.2968393
178.3720	1.730489	0.2858766
187.7600	1.661303	0.2744470
197.1480	1.577359	0.2605794
206.5360	1.491435	0.2463848
215.9240	1.401096	0.2314689
225.3120	1.312728	0.2168625

FIG. B.3. OUTPUT DATA FILE AUTO.

↑ Renamed from DSPS

DATA FILE CREATED BY PROGRAM ATURB2

SMOOTHED POWER SPECTRUM PHI OF P(K)
DATA TAKEN FROM DATA FILE NASV

16384 DATA POINTS WERE USED IN 2L = 153812.9920 METER
1024 DATA POINTS WERE USED IN M = 9613.3120 METER

MEAN VALUE OF W(X) = -0.43757E+01 M/SEC
MEAN SQ. VALUE = 0.60534E+01 (M/SEC)**2

<W OF L(X)**2> = 6.0533

1796 ZEROS WERE ADDED TO DATA

RL(0) = 0.6053E+01

PRINTOUT OF VALUES OF THE SMOOTHED POWER SPECTRUM
K SPS VALUE K CONTD SPS VALUE CONTD

0.000000	0.2211E+04	0.026630	0.5582E+01
0.000052	0.2324E+04	0.026682	0.5629E+01
0.000104	0.2601E+04	0.026734	0.5344E+01
0.000156	0.2601E+04	0.026786	0.5022E+01
0.000208	0.2030E+04	0.026838	0.5740E+01
0.000260	0.1521E+04	0.026890	0.6989E+01
0.000312	0.1446E+04	0.026942	0.6953E+01
0.000364	0.1371E+04	0.026994	0.5436E+01
0.000416	0.1086E+04	0.027046	0.4300E+01
0.000468	0.7951E+03	0.027098	0.4341E+01
0.000520	0.7422E+03	0.027150	0.4698E+01
0.000572	0.8925E+03	0.027202	0.5071E+01
0.000624	0.9883E+03	0.027254	0.5294E+01
0.000676	0.8745E+03	0.027306	0.5114E+01
0.000728	0.6968E+03	0.027358	0.4857E+01
0.000780	0.5919E+03	0.027410	0.5411E+01
0.000832	0.4981E+03	0.027462	0.6069E+01
0.000884	0.4930E+03	0.027514	0.5290E+01
0.000936	0.6537E+03	0.027566	0.4402E+01
0.000988	0.8731E+03	0.027618	0.4795E+01
0.001040	0.9670E+03	0.027670	0.5412E+01
0.001092	0.8431E+03	0.027722	0.5693E+01
0.001144	0.6466E+03	0.027774	0.6767E+01

FIG. B.4. OUTPUT DATA FILE DSPS.

b. Inputs and teletype outputs

- i. number of data points (NPTS); step size of k (DELK); and step size of ξ (DELX)
- ii. program automatically calls in data file "PHILK" containing S_j values
- iii. summation index $N(N)$; initial estimate of L (ALV(1)); and step size of L ($\Delta L \approx STPL$)
- iv. program prints computed values of $E(L)$, $E(L+\Delta L)$ and questions whether the values are positive and negative (opposite signs) – answer yes (Y) or no (N); the program then outputs the interpolated value of L ; this value, L' , is used to compute the new values of $E(L')$ and $E(L'+\Delta L/10)$
- v. again the program questions whether the values of $E(L)$ and $E(L+\Delta L/10)$ are of opposite sign; if the answer is yes (Y) the program computes the final value of L , if the answer is no (N) the program has to be restated with a new value of L [abort (A) program]
- vi. after program computes and prints out value of σ^2 the program can be aborted (A) or continued (C) for the computation of $LG(k_j, L)$ Eq. 3.35, Ref. 1, and the von Karman spectrum and autocorrelation.

c. Outputs

- i. values of L and σ^2 (Eq. 3.25, Ref. 1) are printed out on TTY
- ii. data file LG containing values of $LG(k_j, L)$ (Eq. 3.35, Ref. 1)
- iii. data file PHIXI containing values of the von Karman autocorrelation function:

$$\phi_K(\xi) = \sigma^2 \frac{2^{2/3}}{\Gamma(\frac{1}{3})} (\beta \xi / L)^{1/3} K_{1/3}(\beta \xi / L) - \frac{\beta \xi}{2L} K_{-2/3}(\beta \xi / L)$$

- iv. which is similar to Eq. 4.48, Ref. 1.
data file PHIK containing values of the von Karman spectrum $\phi_{KT}(k)$ or $F_j(L)$, Eq. 3.20, Ref. 1.

d. Example

A typical teletype printout of this interactive program is shown in Fig. B.5. Again, the user supplied information is underscored. Examples of the first page of the output data files are shown in Figs. B.6 through B.8.

```

@LOADER
♦PART2,GAM,AK,AKDAT,PARAB,SIMQ$

PART2 24K CORE, 899 WORDS FREE
LOADER USED 26+5K CORE

EXIT.
^C
@SAVE (CORE FROM) 20 (TO) 777777 (ON) PART2.SAV (NEW FILE)
@ST

INPUT NO. OF POINTS TO BE READ, DELK,DELX
16450,4.7416995E-06,6.436,
INPUT N,L, & STEP SIZE OF L
6326,305.,30.5 ←  $\Delta L \approx \cdot | * L$ 

ON 1 PASS E = -0.8844976E-03
ON 2 PASS E = 0.4644040E-02
ARE THERE POS. AND NEG. E"S (Y OR N) Y ← If answer is no restart
program with another
FOR 2 PASS INTERPOLATTED L = 309.8796 estimate of L
ON 3 PASS E = 0.9391967E-04
ON 4 PASS E = -0.5123543E-03
ARE THERE POS. AND NEG. E"S (Y OR N) Y

FOR 4 PASS INTERPOLATTED L = 309.4071
CONTINUE OR ABORT? (C OR A) C

SIG2 = 1.926393
CONTINUE OR ABORT? (C OR A) C
EXIT

```

FIG. B.5. TTY PRINTOUT FOR RUNNING PROGRAM PART2.

OUTPUT OF ITEM 1, PART 4, PHASE II ATMO. TURB.
CREATED BY PROGRAM PART2

FOR L = 309.4071 METERS

K	LG	K CONTO	LG CONTO
0.4741700E-05	-0.2538761E-03	0.1500274E-01	-1.665575
0.9483399E-05	-0.1015040E-02	0.1500748E-01	-1.665575
0.1422510E-04	-0.2282104E-02	0.1501222E-01	-1.665576
0.1896680E-04	-0.4052757E-02	0.1501696E-01	-1.665577
0.2370850E-04	-0.6323783E-02	0.1502170E-01	-1.665577
0.2845020E-04	-0.9091070E-02	0.1502645E-01	-1.665578
0.3319190E-04	-0.1234963E-01	0.1503119E-01	-1.665579
0.3793360E-04	-0.1609363E-01	0.1503593E-01	-1.665579
0.4267530E-04	-0.2031638E-01	0.1504067E-01	-1.665580
0.4741700E-04	-0.2501044E-01	0.1504541E-01	-1.665581
0.5215869E-04	-0.3016757E-01	0.1505015E-01	-1.665581
0.5690039E-04	-0.3577880E-01	0.1505490E-01	-1.665582
0.6164209E-04	-0.4183448E-01	0.1505964E-01	-1.665583
0.6638379E-04	-0.4832431E-01	0.1506438E-01	-1.665584
0.7112549E-04	-0.5523736E-01	0.1506912E-01	-1.665584
0.7586719E-04	-0.6256217E-01	0.1507386E-01	-1.665585
0.8060889E-04	-0.7028672E-01	0.1507860E-01	-1.665586
0.8535059E-04	-0.7839854E-01	0.1508335E-01	-1.665586
0.9009229E-04	-0.8688474E-01	0.1508809E-01	-1.665587
0.9483399E-04	-0.9573203E-01	0.1509283E-01	-1.665588
0.9957569E-04	-0.1049268	0.1509757E-01	-1.665588
0.1043174E-03	-0.1144552	0.1510231E-01	-1.665589
0.1090591E-03	-0.1243030	0.1510705E-01	-1.665590
0.1138008E-03	-0.1344560	0.1511180E-01	-1.665590
0.1185425E-03	-0.1448997	0.1511654E-01	-1.665591
0.1232842E-03	-0.1556195	0.1512128E-01	-1.665592
0.1280259E-03	-0.1666008	0.1512602E-01	-1.665592
0.1327676E-03	-0.1778290	0.1513076E-01	-1.665593
0.1375093E-03	-0.1892895	0.1513550E-01	-1.665594
0.1422510E-03	-0.2009678	0.1514025E-01	-1.665594
0.1469927E-03	-0.2128494	0.1514499E-01	-1.665595
0.1517344E-03	-0.2249200	0.1514973E-01	-1.665596
0.1564761E-03	-0.2371655	0.1515447E-01	-1.665596
0.1612178E-03	-0.2495720	0.1515921E-01	-1.665597
0.1659595E-03	-0.2621257	0.1516396E-01	-1.665598
0.1707012E-03	-0.2748132	0.1516870E-01	-1.665598
0.1754429E-03	-0.2876214	0.1517344E-01	-1.665599
0.1801846E-03	-0.3005372	0.1517818E-01	-1.665600
0.1849263E-03	-0.3135481	0.1518292E-01	-1.665600
0.1896680E-03	-0.3266419	0.1518766E-01	-1.665601
0.1944097E-03	-0.3398067	0.1519241E-01	-1.665602
0.1991514E-03	-0.3530309	0.1519715E-01	-1.665602
0.2038931E-03	-0.3663033	0.1520189E-01	-1.665603
0.2086348E-03	-0.3796131	0.1520663E-01	-1.665604
0.2133765E-03	-0.3929498	0.1521137E-01	-1.665604
0.2181182E-03	-0.4063035	0.1521611E-01	-1.665605
0.2228599E-03	-0.4196644	0.1522086E-01	-1.665606

FIG. B.6. OUTPUT DATA FILE LG.

DATA FILE CREATED BY PROGRAM PART2

OUTPUT FOR PART 7 OF ITEM I, PHASE II
 CALCULATIONS OF VON-KARMEN AUTOCORRELATION FN

SIGMA SQBD = 1.326393

ξ	$\beta\xi/L$	$\phi_K(\xi)$	Comment: headings not correct in main program format
41.42915	0.1000000	0.9711831	
82.85830	0.2000000	0.7809889	
124.2874	0.3000000	0.6363430	
165.7166	0.4000000	0.5221149	
207.1457	0.5000000	0.4288443	
248.5749	0.6000000	0.3515248	
290.0040	0.7000000	0.2870403	
331.4332	0.8000000	0.2331573	
372.8623	0.9000000	0.1881397	
414.2915	1.0000000	0.1502596	
455.7206	1.1000000	0.1185698	
497.1498	1.2000000	0.9201595E-01	
538.5789	1.3000000	0.6982034E-01	
580.0081	1.4000000	0.5133690E-01	
621.4372	1.5000000	0.3597379E-01	
662.8664	1.6000000	0.2325935E-01	
704.2955	1.7000000	0.1281303E-01	
745.7247	1.8000000	0.4265060E-02	
787.1538	1.9000000	-0.2655269E-02	
828.5830	2.0000000	-0.8212838E-02	
870.0121	2.1000000	-0.1262164E-01	
911.4412	2.2000000	-0.1603264E-01	
952.8704	2.3000000	-0.1864146E-01	
994.2995	2.4000000	-0.2054586E-01	
1035.729	2.5000000	-0.2190808E-01	
1077.158	2.6000000	-0.2277357E-01	
1118.587	2.7000000	-0.2324410E-01	
1160.016	2.8000000	-0.2340022E-01	
1201.445	2.9000000	-0.2329913E-01	
1242.874	3.0000000	-0.2298791E-01	
1284.304	3.1000000	-0.2250897E-01	
1325.733	3.2000000	-0.2189903E-01	
1367.162	3.3000000	-0.2119117E-01	
1408.591	3.4000000	-0.2040945E-01	
1450.020	3.5000000	-0.1958042E-01	
1491.449	3.6000000	-0.1871502E-01	
1532.878	3.7000000	-0.1783107E-01	
1574.308	3.8000000	-0.1694023E-01	
1615.737	3.9000000	-0.1604555E-01	
1657.166	4.0000000	-0.1517245E-01	
1698.595	4.1000000	-0.1431294E-01	
1740.024	4.2000000	-0.1347734E-01	
1781.453	4.3000000	-0.1267253E-01	

FIG. B.7. OUTPUT DATA FILE PHIXI.

DATA FILE CREATED BY PROGRAM PART2

OUTPUT FOR PART 8 OF ITEM I, PHASE II

CALCULATIONS OF PHI K WITH L = 309.4071

AND SIG SORD = 1.326393

**

Comment:

K	PHIK	K CCNTD	PHIK CONTD	
0.4741699E-05	410.4475	0.1500274E-01	2.432225	1) headings should
0.9483399E-05	410.6035	0.1500748E-01	2.430946	line up with
0.1422510E-04	410.8626	0.1501222E-01	2.429668	columns
0.1896680E-04	411.2235	0.1501696E-01	2.428391	2) PHIK = $F_j(L)$
0.2370850E-04	411.6845	0.1502170E-01	2.427115	
0.2845020E-04	412.2432	0.1502645E-01	2.425840	
0.3319190E-04	412.8968	0.1503119E-01	2.424566	
0.3793360E-04	413.6421	0.1503593E-01	2.423293	
0.4267529E-04	414.4755	0.1504067E-01	2.422021	
0.4741699E-04	415.3928	0.1504541E-01	2.420751	
0.5215869E-04	416.3897	0.1505015E-01	2.419481	
0.5690039E-04	417.4612	0.1505490E-01	2.418213	
0.6164209E-04	418.6022	0.1505964E-01	2.416945	
0.6638379E-04	419.8074	0.1506438E-01	2.415679	
0.7112549E-04	421.0711	0.1506912E-01	2.414414	
0.7586719E-04	422.3873	0.1507386E-01	2.413150	
0.8060889E-04	423.7500	0.1507860E-01	2.411886	
0.8535059E-04	425.1530	0.1508335E-01	2.410624	
0.9009229E-04	426.5900	0.1508809E-01	2.409363	
0.9483399E-04	428.0545	0.1509283E-01	2.408103	
0.9957569E-04	429.5402	0.1509757E-01	2.406844	
0.1043174E-03	431.0406	0.1510231E-01	2.405586	
0.1090591E-03	432.5493	0.1510705E-01	2.404329	
0.1138008E-03	434.0599	0.1511180E-01	2.403073	
0.1185425E-03	435.5663	0.1511654E-01	2.401819	
0.1232842E-03	437.0623	0.1512128E-01	2.400565	
0.1280259E-03	438.5420	0.1512602E-01	2.399313	
0.1327676E-03	439.9994	0.1513076E-01	2.398061	
0.1375093E-03	441.4290	0.1513550E-01	2.396810	
0.1422510E-03	442.8254	0.1514025E-01	2.395561	
0.1469927E-03	444.1835	0.1514499E-01	2.394312	
0.1517344E-03	445.4982	0.1514973E-01	2.393065	
0.1564761E-03	446.7650	0.1515447E-01	2.391818	
0.1612178E-03	447.9793	0.1515921E-01	2.390573	
0.1659595E-03	449.1372	0.1516395E-01	2.389329	
0.1707012E-03	450.2347	0.1516870E-01	2.388085	
0.1754429E-03	451.2684	0.1517344E-01	2.386843	
0.1801846E-03	452.2349	0.1517818E-01	2.385602	
0.1849263E-03	453.1312	0.1518292E-01	2.384361	
0.1896680E-03	453.9548	0.1518766E-01	2.383122	
0.1944097E-03	454.7033	0.1519241E-01	2.381884	
0.1991514E-03	455.3744	0.1519715E-01	2.380647	
0.2038931E-03	455.9665	0.1520189E-01	2.379411	

FIG. B.8. OUTPUT DATA FILE PHIK.

APPENDIX C

CONSTRAINED LEAST-SQUARES ESTIMATION OF TURBULENCE AUTOCORRELATION FUNCTION PARAMETERS

This series of programs computes the maximum likelihood estimate of σ_f^2 and length scale L for nonGaussian atmospheric turbulence. The method used is generally described in Sec. 4 of Ref. 1. The computer method used to compute σ^2 vs L is similar to that already discussed in Sec. B.1. The program, PART5, is used to compute up to 15 pairs of $\sigma^2(L); L$ values.

These values are used by program FINAL according to the method outlined in Sec. 4 of Ref. 1 to determine σ_f and L . There are several versions of FINAL tailored to particular turbulence cases. Only one example is covered by this report.

Program Outlines and Usage

Program ATURB3.F4: Computes the two sided power spectrum, S_j , the autocorrelation function, and smoothed power spectrum of the nonGaussian turbulence samples. Usage of this program is similar to that outlined in Sec. B.1.1 of this report.

Program PART5.F4: Computes σ^2 for up to 15 different values of length scale L (Eq. 4.46, Ref. 1).

a. Subroutines — none

b. Inputs

- i. number of data points ($M41$); step size of k ($DELK$); and step size of ξ ($DELX$)
- ii. whether turbulence record is transverse (T) or longitudinal (L)
- iii. program reads in data file ($PHILK$) containing values of S_1 computed by ATURB3.
- iv. lower and upper bounds of k (KL, KU) and number of data points between the two limits (N)
- v. index counter (J) [from 1 to a limit of 15 successively] and value of integral scale L (AL)
- vi. option to continue or abort program after each computation of $\sigma^2(L)$.

c. Outputs

- i. index counter (J) and calculated value of σ^2 for the L chosen (Eq. 4.14, Ref. 1), printed out on TTY.

d. Example

A typical TTY interaction is shown in Fig. C.1.

Program FINAL.F4: Computes constrained least-squares estimation of turbulence autocorrelation function $\phi(\xi)$ (Eq. 4.1, Ref. 1)

a. Subroutines

- i. AK1 and AKDAT - similar to AK and AKDAT described in Sec. B.1.1 except more data points are contained in AKDAT
- ii. GAM and SIMQ - described in Sec. B.1.1
- iii. PAR1 and PAR2 - curve fitting routines
- iv. DGELG - simultaneous linear equation routine using double precision - from the IBM Scientific Subroutine Package
- v. TRAP3 and TRAP6 - subroutines for integration by trapezoidal rule
- vi. ANRP1 - interpolation routine
- vii. SIMP2 - integration by Simpson's rule
- viii. Function DECT and subroutine SET are incorporated in the main program body of FINAL.

b. Inputs and TTY printout

- i. if the program has been executed once and the contents of all registers saved answer yes (Y) and the program will skip the input data phase (skip to step vi), otherwise answer no (N)
 - ii. input step size of ξ (DELX) and number of combinations of σ^2 and L pairs that have been computed previously and *edited* into the main program FINAL (maximum of 11, indexed from 0 to 10). [Since these values are placed in the main body of the program there are different versions of FINAL for each different turbulence record]
 - iii. is turbulence record transverse (T) or longitudinal (L)?
 - iv. integer value of ξ_H/L -(NXIH), $\xi_H(EH)$, and value of m (MM) used in the summation of Eq. 4.1, Ref. 1.
 - v. program automatically reads in data from file AUTO which contains values of the autocorrelation function computed by program ATURB3
- [program prints out value of $\phi(\xi_H)$]

```

@LOADER
♦PART5$

PART5 16K CORE, 1022 WORDS FREE
LOADER USED 18+5K CORE

EXIT.
^C
@ST

INPUT NO. OF POINTS TO BE READ, DELK,DELX
12916,3.09744513E-06,9 .8525

IS RECORD TRANSVERSE OR LONGITUDINAL (T OR L) I

INPUT KL,KU,N
324,12916,12592

INPUT J,L 1,20.

FOR J = 1 $I62 = 0.3392710
PICK ANOTHER J,L (Y OR N) Y

INPUT J,L 1,25.

FOR J = 1 $I62 = 0.3340475
PICK ANOTHER J,L (Y OR N) Y

INPUT J,L 3,35.

FOR J = 3 $I62 = 0.3517353
PICK ANOTHER J,L (Y OR N) Y

INPUT J,L 4,45.

FOR J = 4 $I62 = 0.3832800
PICK ANOTHER J,L (Y OR N) Y

INPUT J,L 5,55.

FOR J = 5 $I62 = 0.4192905
PICK ANOTHER J,L (Y OR N) Y

INPUT J,L 6,65.

FOR J = 6 $I62 = 0.4566356
PICK ANOTHER J,L (Y OR N) Y

INPUT J,L 7,75.

FOR J = 7 $I62 = 0.4940900
PICK ANOTHER J,L (Y OR N) Y

INPUT J,L 8,85.

FOR J = 8 $I62 = 0.5311373
PICK ANOTHER J,L (Y OR N) Y

INPUT J,L 9,95.

FOR J = 9 $I62 = 0.5675591
PICK ANOTHER J,L (Y OR N) N

CPU TIME: 4:18.85      ELAPSED TIME: 14:46.23
NO EXECUTION ERRORS DETECTED

EXIT.
^C
@

```

FIG. C.1. TELETYPE PRINTOUT FOR RUNNING PART5.

- vi. choose an initial value of σ^2 and its associated array index i ; i.e., which index i in the array, $L_i(\sigma^2)$, σ^2 is closest to; the array $L_i(\sigma^2)$ was discussed in Sec. C.1.3b ii.

[program then prints out interpolated value of L for σ^2 value chosen; the value of $dL/d\sigma^2$ (Eq. 4.47, Ref. 1); and IER should equal ϕ or there is an error in the simultaneous linear equation routine]

[program outputs solution of σ^2 as explained in Sec. 4, Ref. 1, pages 106-109. When $\sigma_{\text{INPUT}}^2 \approx \sigma_{\text{OUTPUT}}^2$ a solution has been reached]
- vii. stop loop if $\sigma_{\text{INPUT}}^2 = \sigma_{\text{OUTPUT}}^2$; if equality is not reached program reverts to step vi above, if equality is reached latest value of σ^2 becomes σ_f^2 [if equality is reached the program outputs the interpolated value of the length scale L associated with σ_f^2 and the coefficients of the polynomial approximation to the autocorrelation function of the slow component as discussed in Eq. 4.1 of Ref. 1]
- viii. input value of step size of k (DELK)
save image of core if program is to be rerun
- c. Output data files
 - i. data file ITM2 contains values of the autocorrelation function $\phi(\xi)$
 - ii. data file ITM2L contains values of the normalized spectra $\sigma_f^2 L \phi_K(kL)$, Eq. 4.16, Ref. 1.
- d. Examples

Figure C.2 is an example of the execution of program FINAL while the first pages from data files ITM2 and ITM2L are shown in Figs. C.3 and C.4.

```

@LOADER
*FINAL,AK1,AKDAT,ANRP1,GAM,PAR1,PAR2,SIMP2,SIMQ,TRAP3,TRAP6,DGELG$

FINAL 17K CORE, 245 WORDS FREE
LOADER USED 20+5K CORE

EXIT.
^C
$SAVE (CORE FROM) 20 (TO) 777777 (ON) FINAL.SAV;6 [NEW VERSION]
$ST

HAS DATA BEEN COMPUTED ?
N

INPUT DELX,LMAX
2.8525,10

IS RECORD TRANSVERSE OR LONGITUDINAL (T OR L) T

INPUT HIGHEST INDEX OF XI, EH, & M, 122,1202.005,1

RL(NXIH) = 1.186804
SIMPSONS RULE USED IN IS& I6 UP TO L INDEX = 5
PICK A SIGMA SORD AND ITS ASSOCIATED L INDEX
.443,4

FOR SIGMA SORD = 0.4430000 L INT = 61.37044
DLDS = 268.2701
IER = 0
FOR NCOUNT = 0 SIGMA SORD = 0.4287123156550342D+00

STOP LOOP ? (Y OR N) N

PICK A SIGMA SORD AND ITS ASSOCIATED L INDEX
.44,4

FOR SIGMA SORD = 0.4400000 L INT = 60.56512
DLDS = 268.6132
IER = 0
FOR NCOUNT = 1 SIGMA SORD = 0.4327416339782993D+00

STOP LOOP ? (Y OR N) N

PICK A SIGMA SORD AND ITS ASSOCIATED L INDEX
.435,3

FOR SIGMA SORD = 0.4350000 L INT = 59.22062
DLDS = 269.1852
IER = 0
FOR NCOUNT = 2 SIGMA SORD = 0.4396278931497085D+00

STOP LOOP ? (Y OR N) N

PICK A SIGMA SORD AND ITS ASSOCIATED L INDEX
.437,3

FOR SIGMA SORD = 0.4370000 L INT = 59.75876
DLDS = 268.9564
IER = 0
FOR NCOUNT = 3 SIGMA SORD = 0.4368476183747002D+00

STOP LOOP ? (Y OR N) Y

ON LAST PASS NCOUNT = 3 L = 59.75876
AND SIGMA SORD = 0.4368476183747002D+00
A( 0) = 1.436821
A( 1) = -0.2101162E-03
INPUT DELK 4.9559E-05

CPU TIME: 21.79 ELAPSED TIME: 3:1.07
NO EXECUTION ERRORS DETECTED

EXIT.
^C

```

FIG. C.2. TELETYPE PRINTOUT FOR RUNNING PROGRAM FINAL.

DATA FILE CREATED BY PROGRAM FINAL
WITH N = 1 AND LENGTH = 1399.055 AUTOCOR. OF FILE VERT2

OUTPUT FOR PART 9.K
WITH L = 61.10212 AND SIGMA SQRD = 0.4426216249580745D+00
COEFFICIENT A(0) = 1.432016
COEFFICIENT A(1) = -0.2007736E-03

XI	R(XI)	XI CONTD	R(XI) CONTD
0.0000000	1.874637	695.4260	1.292224
8.181482	1.754460	703.6075	1.290595
16.36296	1.689349	711.7889	1.288956
24.54445	1.639438	719.9704	1.287335
32.72593	1.599677	728.1519	1.285703
40.90741	1.566909	736.3334	1.284070
49.08889	1.539465	744.5149	1.282437
57.27037	1.516304	752.6964	1.280803
65.45186	1.496680	760.8778	1.279153
73.63334	1.480015	769.0593	1.277532
81.81482	1.465732	777.2408	1.275895
89.99630	1.453514	785.4223	1.274259
98.17778	1.443010	793.6038	1.272622
106.3593	1.433951	801.7852	1.270985
114.5407	1.426150	809.9667	1.269347
122.7222	1.419331	818.1482	1.267708
130.9037	1.413496	826.3297	1.266111
139.0852	1.408357	834.5112	1.264468
147.2667	1.403872	842.6927	1.262825
155.4482	1.399920	850.8741	1.261183
163.6296	1.396423	859.0556	1.259540
171.8111	1.393309	867.2371	1.257897
179.9926	1.390528	875.4186	1.256255
188.1741	1.388015	883.6001	1.254612
196.3556	1.385737	891.7815	1.252970
204.5371	1.383639	899.9630	1.251327
212.7185	1.381708	908.1445	1.249684
220.9000	1.379908	916.3260	1.248042
229.0815	1.378214	924.5075	1.246399
237.2630	1.376605	932.6890	1.244756
245.4445	1.375066	940.8704	1.243114
253.6259	1.373583	949.0519	1.241471
261.8074	1.372144	957.2334	1.239829
269.9889	1.370738	965.4149	1.238186
278.1704	1.369356	973.5964	1.236543
286.3519	1.367990	981.7778	1.234901
294.5334	1.366636	989.9593	1.233258
302.7148	1.365238	998.1408	1.231615
310.8963	1.363943	1006.322	1.229973

FIG. C.3. OUTPUT DATA FILE ITM2.

DATA FILE CREATED BY PROGRAM FINAL

WITH M = 1 AND LENGTH = 1399.055

AUTOCOR. OF FILE VERT2

OUTPUT FOR PART 9.L

WITH L = 61.10212

AND SIGMA SQRD = 0.4426216249580745D+00

K	L*SIGMA SQRD*PHIK	K CONTD	L*SIGMA SQRD*PHIK CONTD
0.0000000	27.04512	0.2537421E-01	0.9893473
0.4955900E-04	27.05972	0.2542377E-01	0.9861678
0.9911800E-04	27.10322	0.2547333E-01	0.9830046
0.1486770E-03	27.17468	0.2552289E-01	0.9798576
0.1982360E-03	27.27250	0.2557244E-01	0.9767265
0.2477950E-03	27.39492	0.2562200E-01	0.9736113
0.2973540E-03	27.53912	0.2567156E-01	0.9705120
0.3469130E-03	27.70224	0.2572112E-01	0.9674285
0.3964720E-03	27.88102	0.2577068E-01	0.9643604
0.4460310E-03	28.07196	0.2582024E-01	0.9613077
0.4955900E-03	28.27141	0.2586980E-01	0.9582707
0.5451490E-03	28.47563	0.2591936E-01	0.9552487
0.5947080E-03	28.68093	0.2596892E-01	0.9522419
0.6442670E-03	28.88372	0.2601848E-01	0.9492503
0.6938260E-03	29.08057	0.2606803E-01	0.9462735
0.7433850E-03	29.26826	0.2611759E-01	0.9433116
0.7929440E-03	29.44337	0.2616715E-01	0.9403645
0.8425030E-03	29.60477	0.2621671E-01	0.9374319
0.8920620E-03	29.74855	0.2626627E-01	0.9345141
0.9416210E-03	29.87358	0.2631583E-01	0.9315105
0.9911800E-03	29.97794	0.2636539E-01	0.9287214
0.1040739E-02	30.06047	0.2641495E-01	0.9258465
0.1090298E-02	30.12026	0.2646451E-01	0.9229858
0.1139857E-02	30.15670	0.2651407E-01	0.9201392
0.1189416E-02	30.16946	0.2656362E-01	0.9173066
0.1238975E-02	30.15851	0.2661318E-01	0.9144877
0.1288534E-02	30.12403	0.2666274E-01	0.9116827
0.1338093E-02	30.06643	0.2671230E-01	0.9088914
0.1387652E-02	29.98631	0.2676186E-01	0.9061136
0.1437211E-02	29.88441	0.2681142E-01	0.9033495
0.1486770E-02	29.76152	0.2686098E-01	0.9005986
0.1536329E-02	29.61890	0.2691054E-01	0.8978612
0.1585888E-02	29.45732	0.2696010E-01	0.8951369
0.1635447E-02	29.27801	0.2700966E-01	0.8924260
0.1685006E-02	29.08211	0.2705921E-01	0.8897280
0.1734565E-02	28.87081	0.2710877E-01	0.8870430
0.1784124E-02	28.64530	0.2715833E-01	0.8843709
0.1833683E-02	28.40675	0.2720789E-01	0.8817116
0.1883242E-02	28.15633	0.2725745E-01	0.8790650
0.1932801E-02	27.89519	0.2730701E-01	0.8764312
0.1982360E-02	27.62442	0.2735657E-01	0.8738097
0.2031919E-02	27.34508	0.2740613E-01	0.8712008

FIG. C.4. OUTPUT DATA FILE ITM2L.

APPENDIX D

POWER SPECTRAL DENSITY OF THE INSTANTANEOUS VARIANCE $\sigma_f^2(t)$

The estimation procedure for calculating the power spectrum of σ_f^2 is discussed in Sec. 6.2 of Ref. 5. There are two main programs used to compute σ_f^2 -ATURB4.F4 and ITEM3.F4. The first program, ATURB4, computes the two sided power spectrum of high pass filtered atmospheric turbulence data, $w_h(t)$. The unsmoothed power spectral density, $\Phi_\ell(k)$, and its autocorrelation, $R_{w_h}(\xi)$,

of the high pass filtered data are all computed by this one program. From the square of the high pass filtered samples, w_h^2 , the sample spectrum and autocorrelation function, $R_{w_h^2}(\xi)$ are

formed by a slight variation of ATURB4 called ATUR4A.F4.

The second program, ITEM3, computes the autocorrelation function, $R_{\sigma_f^2}(\tau)$, Eq. 6.40, Ref. 5, and the two sided smoothed spectrum of $\sigma_f^2(x)$:

$$\Phi_{\sigma_f^2}(k) = \{E[\sigma_f^2]\} \left\{ \delta(f) + \int_{-M}^M p_0(\xi) \left(\frac{R_{w_h^2}(\xi) - [R_{w_h^2}(0)]^2 - 2[R_{w_h}(\xi)]^2}{[R_{w_h}(0)]^2 + 2[R_{w_h}(\xi)]^2} \right) \times e^{-12\pi k \xi_d \xi} \right\}$$

$$\text{where } p_0(\xi) = \left\{ \frac{1}{\pi} \left| \sin \frac{\pi \xi}{M} \right| + \left(1 - \frac{|\xi|}{M} \right) \cos \frac{\pi \xi}{M} \right\} |\xi| \leq M$$

This equation is similar to Eq. 6.49, Ref. 5 except the Papoulis window, $p_0(\xi)$, has been added to smooth the power spectrum.

Program Outlines and Usage

Program ATURB4: Computes the two sided power spectra and autocorrelation function of high pass filtered turbulence samples. (ATUR4A parallels ATURB4 except it operates on the square of high pass filtered turbulence samples).

a. Subroutines

- i. CFFT and CFFT1 - fast Fourier transform routines for ATURB4 and ATUR4A respectively
- ii. SIMP - integration by Simpson's rule
- iii. HPDES - digital high pass filter (routine C-2 of Appendix C of Ref. 6)

b. Inputs

- i. through v are the same as Sec. B.1.1b i to v
- vi. input the cut-off frequency k_c (FC) for the filter routine; sampling interval (TS) in seconds; and number of filter sections (NS), see Ref. 6
- vii. yes (Y) or no (N) to performing integration check as per B.1.1b vi

c. Outputs

- i. data files PHILK and FPSD2 contain positive frequency domain (k) values of the power spectrum of the high pass filtered data and the high pass filtered squared samples respectively
- ii. data files AUTO and AUTF2 contain the autocorrelation functions of the filtered and filtered squared data

d. Example

The teletype printout involved with program ATURB4 is shown in Fig. D.1. ATUR4A would involve an identical interaction (except CFFT1 would be loaded instead of CFFT with the main program). Figures D.2 through D.5 show examples of outputs of ATURB4 - PHILK and AUTO; and ATUR4A - FPSD2 and AUTF2.

Program ITEM3.F4: Computes the spectrum $\Phi_{\sigma_f^2}(k)$ and autocorrelation function $R_{\sigma_f^2}(\xi)$.

a. Subroutines

- i. CFFT1 - fast Fourier transform program

b. Inputs

- i. step size of ξ (DELX)
- ii. program automatically reads data from files AUTO and AUTF2 containing R_{w_h} and $R_{w_h^2}$

```

(LOADER U^C
^C
@LOADER
♦ATURB4,CFFT,SIMP,HPDESS

ATURB4 69K CORE, 762 WORDS FREE
LOADER USED 72+5K CORE

EXIT.
^C
@SAVE (CORE FROM) 20 (TO) 777777 (ON) ATURB4.SAV;2 [NEW VERSION]
@ST

INPUT SPEED OF CRAFT (M/SEC)
197.05

INPUT TOTAL NO. OF POINTS TO BE USED IN 2L M. OF DATA
AND POWER OF TWO OF THAT NO.
32768
15

INPUT NO. OF POINTS OF W(X) TO BE READ
AND SAMPLING RATE OF DATA 15000,.05

INPUT VALUE OF MPTS
AND POWER OF TWO OF THAT NO.
1024
10

INPUT DATA FILE NAME THAT CONTAINS SAMPLES
OF W(X) AND NO. POINTS NXRAY
VERT2
15000

INPUT CUT-OFF FREQ,SAMPLING INTERVAL,NO. OF FILTER SECTIONS5.9115E-01,.0
♦♦5,2
PERFORM INTEGRATION CHECK (Y OR N) Y

INTEGRAL OF PHI OF L (K) = 0.2211E+00

CPU TIME: 6:50.24 ELAPSED TIME: 13:19.00
NO EXECUTION ERRORS DETECTED

EXIT.
^C

```

FIG. D.1. TELETYPE PRINTOUT FOR RUNNING ATURB4.

DATA FILE CREATED BY PROGRAM ATURB4

HIGH PASS FILTERED DATA

CUT-OFF FREQ = 0.5911500 (K = 0.3000000E-02)

POWER SPECTRUM OF PHI OF L(K)

DATA TAKEN FROM FILE VERT2

32768 DATA POINTS WERE USED IN 2L = 322846.7190 METER
1624 DATA POINTS WERE USED IN M = 10088.9600 METER

17768 ZEROS WERE ADDED TO DATA

MEAN VALUE OF W(X) = -0.74214E-05 M/SEC

MEAN SQ. VALUE = 0.22110E+00 (M/SEC)**2

<W OF L(X)**2> = 0.2210

PRINTOUT OF THE VALUES OF THE POWER SPECTRUM

K	PS VALUE	K CONTD	PS VALUE CONTD
0.000000	0.1348E-13	0.025374	0.8273E+00
0.000003	0.1048E-04	0.025377	0.1553E+00
0.000006	0.0635E-05	0.025380	0.9468E-01
0.000009	0.7124E-05	0.025384	0.4748E-01
0.000012	0.9423E-05	0.025387	0.1841E-01
0.000015	0.7117E-05	0.025390	0.1567E+00
0.000019	0.9116E-05	0.025393	0.5845E+00
0.000022	0.7376E-05	0.025396	0.9412E+00
0.000025	0.8759E-05	0.025399	0.8532E+00
0.000028	0.7711E-05	0.025402	0.4245E+00
0.000031	0.8414E-05	0.025405	0.1030E+01
0.000034	0.8044E-05	0.025408	0.1135E+01
0.000037	0.8110E-05	0.025411	0.3692E+00
0.000040	0.8260E-05	0.025415	0.7724E+00
0.000043	0.7895E-05	0.025418	0.6522E+00
0.000046	0.8417E-05	0.025421	0.2323E+00
0.000050	0.7823E-05	0.025424	0.2026E-01
0.000053	0.8465E-05	0.025427	0.4337E+00
0.000056	0.7831E-05	0.025430	0.8168E+00
0.000059	0.8410E-05	0.025433	0.5790E+00
0.000062	0.7951E-05	0.025436	0.1324E+01
0.000065	0.8351E-05	0.025439	0.1206E+01
0.000068	0.8036E-05	0.025442	0.4376E+00
0.000071	0.8143E-05	0.025446	0.1843E+01

FIG. D.2. OUTPUT DATA FILE PHILK.

DATA FILE CREATED BY PROGRAM ATURB4

AUTOCORRELATION OF
HIGH PASS FILTERED NON-HOMOGENEOUS SAMPLE
DATA TAKEN FROM FILE VERT2

32768 DATA POINTS WERE USED IN 2L = 322846.7190 METER
1024 DATA POINTS WERE USED IN M = 10088.9600 METER

17768 ZEROS WERE ADDED TO DATA

MEAN VALUE OF W(X) = -0.74214E-05 M/SEC
MEAN SQ. VALUE = 0.22110E+00 (M/SEC)**2

<W OF L(X)**2> = 0.2210

TRUNCATION POINT WAS 13000.00 METERS
WHICH CONTAINS 1319 POINTS

PRINTOUT OF THE VALUES OF THE AUTOCORRELATION

X	RL	RL/R0
0.0000000	0.2210828	1.0000000
9.852500	0.1366631	0.6181535
19.70500	0.7223809E-01	0.3267467
29.55750	0.2718233E-01	0.1229509
39.41000	-0.5175197E-02	-0.2340841E-01
49.26250	-0.2545530E-01	-0.1151392
59.11500	-0.3989603E-01	-0.1804574
68.96750	-0.4644818E-01	-0.2100940
78.82000	-0.4698692E-01	-0.2125308
88.57250	-0.4595287E-01	-0.2078536
98.52500	-0.4678660E-01	-0.2116248
108.3775	-0.4089274E-01	-0.1849657
118.2300	-0.3525967E-01	-0.1594862
128.0825	-0.3052934E-01	-0.1380901
137.9350	-0.2747492E-01	-0.1242743
147.7875	-0.2431214E-01	-0.1099685
157.6400	-0.2179220E-01	-0.9857031E-01
167.4925	-0.1709930E-01	-0.7734341E-01
177.3450	-0.1068496E-01	-0.4833010E-01
187.1975	-0.3274173E-02	-0.1480971E-01
197.0500	0.1146689E-02	0.5186693E-02
206.9025	0.6017922E-02	0.2722021E-01
216.7550	0.9448106E-02	0.4273559E-01
226.6075	0.1239062E-01	0.5604516E-01

FIG. D.3. OUTPUT DATA FILE AUTO.

DATA FILE CREATED BY PROGRAM ATURB4

HIGH PASS FILTERED DATA

CUT-OFF FREQ = 0.5911500 (K = 0.3000000E-02)

FILTERED VERSION SQUARED BEFORE TRANSFORMING

POWER SPECTRUM OF PHI OF L(K)

DATA TAKEN FROM FILE VERT2

32768 DATA POINTS WERE USED IN 2L = 322846.7190 METER

1024 DATA POINTS WERE USED IN M = 10088.9600 METER

17768 ZEROS WERE ADDED TO DATA

MEAN VALUE OF W(X) = -0.74214E-05 M/SEC

MEAN SQ. VALUE = 0.22110E+00 (M/SEC)**2

<W OF L(X)**2> = 0.4627

PRINTOUT OF THE VALUES OF THE POWER SPECTRUM

K	PS VALUE	K CONTD	PS VALUE CONTD
0.000000	0.7224E+04	0.025374	0.2112E+01
0.000003	0.3957E+04	0.025377	0.3822E+01
0.000006	0.6459E+03	0.025380	0.2336E+01
0.000009	0.1359E+04	0.025384	0.2389E+01
0.000012	0.1438E+04	0.025387	0.5410E+01
0.000015	0.5900E+03	0.025390	0.3918E+01
0.000019	0.3731E+03	0.025393	0.7773E+00
0.000022	0.1381E+03	0.025396	0.2330E+01
0.000025	0.7375E+02	0.025399	0.3844E+01
0.000028	0.3612E+03	0.025402	0.2807E+01
0.000031	0.5751E+03	0.025405	0.2275E+01
0.000034	0.3753E+03	0.025408	0.1228E+01
0.000037	0.2580E+02	0.025411	0.1186E+00
0.000040	0.2465E+03	0.025415	0.1594E+01
0.000043	0.4334E+03	0.025418	0.2782E+01
0.000046	0.1599E+03	0.025421	0.1292E+01
0.000050	0.2256E+03	0.025424	0.7048E+00
0.000053	0.3186E+03	0.025427	0.3248E+01
0.000056	0.1056E+03	0.025430	0.4490E+01
0.000059	0.2861E+02	0.025433	0.1749E+01
0.000062	0.2717E+02	0.025436	0.4573E+00
0.000065	0.6028E+02	0.025439	0.3822E+01
0.000068	0.1258E+03	0.025442	0.7261E+01

FIG. D.4. OUTPUT DATA FILE FPSD2.

DATA FILE CREATED BY PROGRAM ATURB4

AUTOCORRELATION OF
HIGH PASS FILTERED AND SQUARED NON-HOMOGENEOUS SAMPLE
DATA TAKEN FROM FILE VERT2

32768 DATA POINTS WERE USED IN 2L = 322846.7190 METER
1024 DATA POINTS WERE USED IN M = 10088.9600 METER

17768 ZEROS WERE ADDED TO DATA

MEAN VALUE OF $W(X)$ = $-0.74214E-05$ M/SEC
MEAN SQ. VALUE = $0.22110E+00$ (M/SEC)**2

<W OF L(X)**2> = 0.4627

TRUNCATION POINT WAS 13000.00 METERS
WHICH CONTAINS 1319 POINTS

PRINTOUT OF THE VALUES OF THE AUTOCORRELATION

X	RL	RL/RL0
0.0000000	0.4629917	1.0000000
9.852500	0.2634734	0.5690672
19.70500	0.1654293	0.3573051
29.55750	0.1367036	0.2952614
39.41000	0.1272111	0.2747590
49.26250	0.1450926	0.3133805
59.11500	0.1577147	0.3406427
63.96750	0.1564650	0.3379434
78.82000	0.1460383	0.3154231
88.57250	0.1341455	0.2897362
98.52500	0.1194356	0.2579648
108.3775	0.1148353	0.2480288
118.2300	0.1207423	0.2607872
128.0825	0.1143026	0.2468782
137.9350	0.1129238	0.2439003
147.7875	0.1129470	0.2439503
157.6400	0.1232595	0.2662241
167.4925	0.1278594	0.2761592
177.3450	0.1114554	0.2407286
187.1975	0.1117146	0.2412886
197.0500	0.1127996	0.2436319
206.9025	0.1065844	0.2302079
216.7550	0.1123870	0.2427407
226.6075	0.1149910	0.2483651

FIG. D.5. OUTPUT DATA FILE AUTF2.

- iii. input σ_f^2 (SIGF) and L determined in Sec. C.1.3
- iv. number of points in Fourier transform (MPTS) and power of two of that number (MPWRM)

c. Outputs

- i. data file RSIGF containing values of $R_{\sigma_f^2}(\xi)$
- ii. data file PHIF containing the smoothed power spectrum values of $\Phi_{\sigma_f^2}(k)$.

d. Example

An example of the input through the teletype is given in Fig. D.6 while Figs. D.7 and D.8 show one page from each of the data files RSIGF and PHIF.

@LOADER
♦ITEM3,CFPT1\$

ITEM3 11K CORE, 353 WORDS FREE
LOADER USED 14+5K CORE

EXIT.
^C
@ST

INPUT DELX
9.8525

RWH(MAOFF) = 0.2880105E-02
RWH2(MAOFF) = 0.4643580E-01
INPUT SIGMA SQRD, L
.4609,65.89

INPUT M AND POWER OF 2
1024,10

CPU TIME: 52.13 ELAPSED TIME: 1:53.21
NO EXECUTION ERRORS DETECTED

EXIT.
^C

FIG. D.6. TELETYPE OUTPUT FOR RUNNING ITEM3.

DATA FILE CREATED BY PROGRAM ITEM3
FOR SIGMA SQUARED = 0.4609000 AND AL = 65.89000

XI	RSIGF
0.0000000	0.6707419
9.852500	0.6490613
19.70500	0.5924702
29.55750	0.5766969
39.41000	0.5522717
49.26250	0.6143047
59.11500	0.6435365
68.96750	0.6248567
78.82000	0.5821151
88.67250	0.5366453
98.52500	0.4764114
108.3775	0.4671271
118.2300	0.4993594
128.0825	0.4785250
137.9350	0.4760772
147.7875	0.4792910
157.6400	0.5254913
167.4925	0.5491249
177.3450	0.4821481
187.1975	0.4853142
197.0500	0.4902162
206.9025	0.4625451
216.7550	0.4866718
226.6075	0.4966468
236.4600	0.5135362
246.3125	0.5233681
256.1650	0.5119246
266.0175	0.4612299
275.8700	0.4421263
285.7225	0.4679804
295.5750	0.4891256
305.4275	0.5169329
315.2800	0.5298064
325.1325	0.5117880
334.9850	0.5550355
344.8375	0.5224204
354.6900	0.4742931
364.5425	0.4484167
374.3950	0.4570057
384.2475	0.4884470
394.1000	0.5384777
403.9525	0.5555076
413.8050	0.5482229
423.6575	0.4959692
433.5100	0.4385263
443.3625	0.4048591

FIG. D.7. OUTPUT DATA FILE RSIGF.

DATA FILE CREATED BY PROGRAM ITEM3

SMOOTHED POWER SPECTRUM PHI OF SIGMA SQRD F(K)

1024 DATA POINTS WERE USED IN M = 10080.9600 FT.
RWH(0) = 0.2210028

PRINTOUT OF VALUES OF THE SMOOTHED POWER SPECTRUM
K SPS VALUE K CONTD SPS VALUE CONTD

0.000000	0.1388E+04	0.025374	0.1005E+01
0.000050	0.1007E+04	0.025424	0.3967E+00
0.000099	0.4127E+03	0.025473	-0.1887E+01
0.000149	0.1884E+03	0.025523	-0.3033E+01
0.000198	0.1740E+03	0.025573	-0.7258E+00
0.000248	0.1394E+03	0.025622	0.3225E+01
0.000297	0.8768E+02	0.025672	0.3959E+01
0.000347	0.5885E+02	0.025721	0.1401E+01
0.000396	0.3664E+02	0.025771	0.1112E+01
0.000446	0.2377E+02	0.025820	0.3751E+01
0.000496	0.4061E+02	0.025870	0.3992E+01
0.000545	0.6361E+02	0.025919	0.1706E+01
0.000595	0.6199E+02	0.025969	0.6458E+00
0.000644	0.5725E+02	0.026019	-0.1602E+00
0.000694	0.7100E+02	0.026068	-0.1048E+01
0.000743	0.7442E+02	0.026118	0.3609E+00
0.000793	0.5749E+02	0.026167	0.1975E+01
0.000843	0.5533E+02	0.026217	0.2240E+01
0.000892	0.7045E+02	0.026266	0.2586E+01
0.000942	0.6358E+02	0.026316	0.1789E+01
0.000991	0.8858E+02	0.026365	0.8118E+00
0.001041	-0.1601E+01	0.026415	0.1144E+01
0.001090	-0.9461E+01	0.026465	0.1232E+01
0.001140	-0.6385E+01	0.026514	0.8538E+00
0.001189	-0.5456E+01	0.026564	-0.5051E+00
0.001239	-0.7733E+01	0.026613	-0.2148E+01
0.001289	-0.9891E+01	0.026663	-0.2239E+01
0.001338	-0.6303E+01	0.026712	-0.1680E+01
0.001388	0.3264E+01	0.026762	-0.1419E+01
0.001437	0.8319E+01	0.026811	-0.1143E+01
0.001487	0.8763E+01	0.026861	-0.1863E+01
0.001536	0.9509E+01	0.026911	-0.4201E+01
0.001586	0.7367E+01	0.026960	-0.4662E+01
0.001635	0.8612E+00	0.027010	-0.1028E+01
0.001685	-0.7582E+01	0.027059	0.4374E+02
0.001735	-0.9394E+01	0.027109	-0.2595E+00
0.001784	0.2060E+01	0.027158	0.1573E+00
0.001834	0.1579E+02	0.027208	0.1628E+01
0.001883	0.1587E+02	0.027258	0.1666E+01

FIG. D.8. OUTPUT DATA FILE PHIF.

APPENDIX E

PROBABILITY DENSITY ESTIMATION OF THE INSTANTANEOUS VARIANCE $\sigma_f^2(t)$ AND THE "SLOW" TURBULENCE COMPONENT $w_s(t)$

The methods used to develop the probability density function of $\sigma_f^2(t)$ and $w_s(t)$ are given in Sec. 6.3 and 6.4 of Ref. 5. Three programs are necessary to carry out the computations - MOMENT.F4, ITEM.F4, and GDIST6.F4. MOMENT calculates the moments of σ_f^2 as per Eq. 6.71, Ref. 5. GDIST computes the probability density function of σ_f^2 as described in Ref. 5 - BBN Report 3476, pages 88-91. ITEM4 calculates the moments of w_s , M_{w_s} , Eq. 6.87, Ref. 5 and the probability density function P_{w_s} , Eq. 6.93, Ref. 5.

Program Outlines and Usage

Program MOMENT.F4: Computes first 8 moments of σ_f^2 ; $M_{\sigma_f^2}$, Eq. 6.71, Ref. 5.

a. Subroutines

- i. BIN - tabulates the number of samples in a particular bin
- ii. BINSQ - counts the number of filtered squared samples in a certain bin
- iii. HPDES - high pass digital filter routine from Ref. 6

b. Inputs

- i. number of data points to be used (NPTS)
- ii. name to be used for output data file [FLE, A5 format]
- iii. bin width (BINW)
- iv. number of bins (NBIN)
- v. cut-off frequency k_c (FC); sampling interval (TS), and number of filter section (NS)
- vi. bin width and number of bins for filtered data (BINW and NBIN)
- vii. bin width and number of bins for filtered squared data (BINW1 and NBIN1)
- viii. input σ_f^2 (computed in Sec. C.1.3)

c. Outputs (at TTY)

- i. moments of $w(t)$; M^n $n = 1$ to 8
- ii. moments of w_h ; $M_{w_h}^n$ $n = 1$ to 8

- iii. moments of w_h^2 ; $M_{w_h}^n$ $n = 1$ to 8
- iv. moments of σ_f^2 ; $M_{\sigma_f^2}^n$ $n = 1$ to 8

d. Example

Figure E.1 contains an example of the teletype inputs and outputs.

Program GDIST6.F4: Computes the probability density distribution $P_{\sigma_f^2}$

a. Subroutine

- i. GAM — computes the gamma function Γ

b. Inputs

- i. moments of σ_f^2 ; $M_{\sigma_f^2}$

c. Outputs (at the TTY)

- i. γ — Eq. 6.75, Ref. 5
- ii. $P_{\sigma_f^2}$ — the probability density function of σ_f^2

d. Example

Figure E.2 contains a *partial* listing of the terminal print-out for program GDIST6.

Program ITEM4.F4: Calculates the probability density function of $w_s(t)$: P_{w_s}

a. Subroutine

- i. FAC1 — factorial routine

b. Inputs

- i. moments of w_h and w ; $M_{w_h}^n$ and M_w^n , $n = 1$ to 8
- ii. σ_f^2

```

LOADER
MOMENT,BIN,BINSD,MPDESS

MOMENT 20K CORE, 344 WORDS FREE
LOADER USED 23.5K CORE

EXIT.
-C
QST

INPUT NO. OF DATA POINTS 15000

INPUT DATA RECORD NAME VERT2

INPUT BIN WIDTH
NO. OF BINS TOTAL
.014
1000

TOTAL NO. OF POINTS USED FOR 1ST PASS =15000.00
MOMENTS OF U(X)
K = 1 ALPHU = 0.4662677E-03
K = 2 ALPHU = 1.811680
K = 3 ALPHU = 6.2193678
K = 4 ALPHU = 10.63853
K = 5 ALPHU = 3.433047
K = 6 ALPHU = 116.1995
K = 7 ALPHU = 142.7648
K = 8 ALPHU = 1994.063
INPUT CUT-OFF FREQ,SAMPLING INTERVAL,NO. OF FILTER SECTIONS
5.9113E-01,.05,2

MOMENTS OF FILTERED U

INPUT BIN WIDTH
NO. OF BINS TOTAL
.014
1000

TOTAL NO. OF POINTS USED FOR 1ST PASS =15000.00
K = 1 ALPHU = -0.2426459E-04
K = 2 ALPHU = 0.2211038
K = 3 ALPHU = 0.8224017E-02
K = 4 ALPHU = 6.627326
K = 5 ALPHU = 0.1194177
K = 6 ALPHU = 2.671960
K = 7 ALPHU = 2.380871
K = 8 ALPHU = 25.761930
COMPUTE MOMENTS FOR U=U

INPUT BIN WIDTH
NO. OF BINS
.03521
700

TOTAL NO. OF POINTS USED FOR 1ST PASS =15000.00
K = 1 ALPHU2 = 0.2261919
K = 2 ALPHU2 = 0.4630958
K = 3 ALPHU2 = 2.674421
K = 4 ALPHU2 = 23.85840
K = 5 ALPHU2 = 318.6409
K = 6 ALPHU2 = 4431.121
K = 7 ALPHU2 = 45817.35
K = 8 ALPHU2 = 1817433.
COMPUTE MOMENTS OF SIGMA SQRD F
INPUT SIG SQRD F .4609

N= 1 ALSIG = 0.4609010
N= 2 ALSIG = 0.6409275
N= 3 ALSIG = 1.309544
N= 4 ALSIG = 4.243320
N= 5 ALSIG = 11.84154
N= 6 ALSIG = 30.31175
N= 7 ALSIG = 71.85787
N= 8 ALSIG = 149.1768

CPU TIME: 1:22.33 ELAPSED TIME: 9:59.21
NO EXECUTION ERRORS DETECTED

EXIT.

```

FIG. E.1. TELETYPE PRINTOUT OF INPUTS AND OUTPUTS OF PROGRAM MOMENT.

F40
 ♦GAM=GAM
 GAM ERRORS DETECTED: 0

9K CORE USED

♦^C
 @LOADER ♦
 ♦GDIST6,GAM\$

GDIST6 3K CORE, 197 WORDS FREE
 LOADER USED 6+5K CORE

EXIT.
 ^C
 @ST

INPUT ALPHF :
 .4609
 .6409275
 1.509566
 4.24552
 11.84454
 30.51175

GAMMA = 0.4957514

X	F1PRM	F2PRM	F3PRM	F4PRM
0.07	2.629587	2.568655	2.490394	2.440631
0.13	1.730845	1.729292	1.753583	1.798062
0.20	1.316895	1.340347	1.403987	1.480339
0.26	1.063073	1.098345	1.177042	1.261538
0.33	0.8864120	0.9266947	1.008244	1.089494
0.39	0.7543590	0.7957138	0.8734558	0.9457618
0.46	0.6510620	0.6911143	0.7614079	0.8221444
0.52	0.5676976	0.6050170	0.6660199	0.7143386
0.59	0.4988822	0.5326439	0.5836338	0.6197396
0.65	0.4411022	0.4708831	0.5118475	0.5365624
0.72	0.3919436	0.4175922	0.4489650	0.4634521
0.79	0.3496791	0.3712275	0.3937150	0.3993006
0.85	0.3130320	0.3306351	0.3450979	0.3431546
0.92	0.2810325	0.2949255	0.3022969	0.2941690
0.98	0.2529274	0.2633948	0.2646258	0.2515816
1.05	0.2281204	0.2354738	0.2314955	0.2146990
1.11	0.2061320	0.2106943	0.2023926	0.1828883
1.18	0.1865708	0.1886646	0.1768647	0.1555721
1.24	0.1691139	0.1690534	0.1545104	0.1322240
1.31	0.1534916	0.1515770	0.1349719	0.1123657
1.37	0.1394771	0.1359909	0.1179288	0.9556370E-01
1.44	0.1268777	0.1220822	0.1030944	0.8142707E-01
1.51	0.1155286	0.1096651	0.9021136E-01	0.6960408E-01

FIG. E.2. TELETYPE PRINTOUT FOR RUNNING GDIST6.

c. Outputs

- i. data file PROB containing moments of w_s , M_{w_s} ,
and the probability density function P_{w_s} .

d. Example

Figure E.3 is an example of the teletype printout created by program ITEM4 and Fig. E.4 contains the first page of data file PROB.

ITEM4 3K CORE, 364 WORDS FREE
LOADER USED 6+5K CORE

EXIT.

^C

@ST

INPUT ALPHWH,ALPHW (8 VALUES)

-.2426659E-04,.4662677E-05

.2211058,1.8116

.8224017E-02,.2195678

.462732,10.63055

.1394177,5.433067

2.67196,116.9995

2.380871,162.7648

25.7693,1994.063

INPUT SIGMA SQRD F .4609

PROB DENSITY FN.

CPU TIME: 1.87 ELAPSED TIME: 1:47.82

NO EXECUTION ERRORS DETECTED

EXIT.

^C

@LIST PROB.DAT;3

@

FIG. E.3. TELETYPE PRINTOUT FOR RUNNING ITEM4.

DATA FILE CREATED BY PROGRAM ITEM4
ATMOSPHERIC TURBULENCE TASK, PHASE II
FOR SIGMA SQD F = 0.4609000

THE VALUES OF ALPHA OF WS ARE:
N = 1ALPHA WS = 0.3969848E-04
N = 2ALPHA WS = 1.350700
N = 3ALPHA WS = 0.1949038
N = 4ALPHA WS = 4.884669
N = 5ALPHA WS = 3.326249
N = 6ALPHA WS = 18.19409
N = 7ALPHA WS = 56.67557
N = 8ALPHA WS = -344.2492

THE PROBABILITY DENSITY FN. IS:

WS	P(WS)
0.0000000	0.3294248
-0.4702189E-01	0.3300615
-0.9404378E-01	0.3302435
-0.1410657	0.3299658
-0.1880876	0.3292255
-0.2351094	0.3280209
-0.2821313	0.3263524
-0.3291532	0.3242220
-0.3761751	0.3216337
-0.4231970	0.3185931
-0.4702189	0.3151080
-0.5172408	0.3111880
-0.5642627	0.3068443
-0.6112845	0.3020904
-0.6583064	0.2969412
-0.7053283	0.2914138
-0.7523502	0.2855266
-0.7993721	0.2792998
-0.8463940	0.2727550
-0.8934159	0.2659153
-0.9404378	0.2588050
-0.9874597	0.2514495
-1.034482	0.2438750
-1.081503	0.2361088
-1.128525	0.2281785
-1.175547	0.2201123
-1.222569	0.2119386
-1.269591	0.2036858
-1.316613	0.1953822
-1.363635	0.1870557
-1.410657	0.1787340
-1.457679	0.1704437
-1.504700	0.1622108
-1.551722	0.1540602

FIG. E.4. OUTPUT DATA FILE PROB.

[illegible]

APPENDIX F: COMPUTER PROGRAM LISTING

Subroutine AK1

```

SUBROUTINE AK(I,XP1,AKZ)
C
COMMON/AA/AK1(100),AK2(100)
C
C
IF (XP1.GT.10..OR.XP1.LT.1) GO TO 10
PI = 3.14159265
IXMID = IFIX(XP1/.1 + .5)
IF (I.EQ.1) AKZ = AK1(IXMID)*PI/2.
IF (I.EQ.2) AKZ = AK2(IXMID)*PI/2.
C
GO TO 30
C
10 IF (XP1.LT.1) TYPE 40,XP1
IF (XP1.GT.10.) TYPE 20,XP1
20 FORMAT(' X GT 10.0, CANNOT COMPUTE, X =',G)
40 FORMAT(' X LT 0.1, CANNOT COMPUTE, X =',G)
30 RETURN
C
END

```


Subroutine AK

```

C <RFISHER>AK.F4;10      1-Nov-77 11:55:59      EDIT BY RFISHER
      SUBROUTINE AK (I,XP1,AKZ)
C
C      COMPUTES MODIFIED BESSEL FNS OF FRACTIONAL ORDER 1/3
C      AND 2/3. TABLE OF K(1/3) AND K(2/3) FROM MATH HANDBOOK
C      VALUES OF AK1 ARE FOR MODIFIED BESSEL FNS: (2/PI)*K(1/3)
C      FOR VALUES OF X FROM .1 TO 5.1 IN .1 STEPS
C      VALUES OF AK2 ARE FOR MODIFIED BESSEL FNS: (2/PI)*K(2/3)
C      FOR VALUES OF X AS ABOVE
C      USES SUBROUTINE PARAB TO FIT 3 DATA PTS TO INTERPOLATE
C      BETWEEN STEP OF .1
C
      COMMON/AA/AK1 (51) ,AK2 (51)
      COMMON/BB/X (3) ,AKI (3) ,XPL
C
C      XPL = XP1
C
C      CHECK IF WITHIN RANGE OF .1 TO 5.1
C
      IF (XPL.GT.5..OR.XPL.LT..1) GO TO 10
      IXP1 = IFIX(10.*XP1)
      PI = 3.14159265
      IXMID = IFIX(XPL/.1 + .5)
C
C      CHECK IF INTERPOLATION NECESSARY
C
      IF ((10.*XP1).EQ.IXP1.AND.I.EQ.1) AKZ = AK1 (IXMID) *PI/2.
      IF ((10.*XP1).EQ.IXP1.AND.I.EQ.2) AKZ = AK2 (IXMID) *PI/2.
      IF ((10.*XP1).EQ.IXP1) GO TO 30
C
      IXLOW = IXMID - 1
      IXHIG = IXMID + 1
C
C      CHECK LOWER BOUND OF ARRAY
C
      IF (IXMID.EQ.1) IXHIG = 3
      IF (IXMID.EQ.1) IXLOW = 1
      IF (IXMID.EQ.1) IXMID = 2
      AKI (1) = AK2 (IXLOW)
      AKI (2) = AK2 (IXMID)
      AKI (3) = AK2 (IXHIG)
      IF (I.EQ.1) AKI(1) = AK1 (IXLOW)
      IF (I.EQ.1) AKI(2) = AK1 (IXMID)
      IF (I.EQ.1) AKI(3) = AK1 (IXHIG)
      X (1) = .1*IXLOW
      X (2) = .1*IXMID
      X (3) = .1*IXHIG
      CALL PARAB (AKZ)
      AKZ = AKZ*PI/2.
      GO TO 30

```

```

C
10  IF (XPL.LT..1) TYPE 40,XPL
    IF (XPL.GT.5.) TYPE 20,XPL
20  FORMAT(' X GT 5.0, CANNOT COMPUTE, X =',G)
40  FORMAT(' X LT 0.1, CANNOT COMPUTE, X =',G)
30  RETURN
C
    END

```

Subroutine AKDAT (for use with AK)

```

C      USED BY SUBROUTINE AK, ATMOSPHERIC TURBULENCE
C      VALUES OF AK1 ARE FOR MODIFIED BESSEL FNS:  $(2/\pi)*K(1/3)$ 
C      FOR VALUES OF X FROM .1 TO 5.1 IN .1 STEPS
C      VALUES OF AK2 ARE FOR MODIFIED BESSEL FNS:  $(2/\pi)*K(2/3)$ 
C      FOR VALUES AS ABOVE
C      BLOCK DATA
C
C      COMMON/AA/AK1(51),AK2(51)
C
C      DATA AK1/1.8461,1.2601,.9607,.7676,.6296,.5253,.4434,
1     .3776,.3238,.27911,.24167,.21001,.18306,.160,.14016,.12302,
2     .10818,.09527,.08402,.07419,.06559,.05805,.05142,.04559,
3     .04045,.03592,.03192,.02838,.025249,.022476,.020018,.017838,
4     .015902,.014183,.012654,.011295,.010085,.009008,.008049,
5     .007194,.006432,.005752,.005145,.004604,.004120,.003688,
6     .003303,.0029578,.0026495,.0023739,.0021273/
C
C      DATA AK2/3.026,1.7837,1.2716,.9681,.7678,.6257,.5187,.4354,
1     .3688,.3148,.27024,.23312,.20191,.17547,.15294,.13364,.11704,
2     .10270,.09027,.07947,.07007,.06185,.05466,.04835,.04282,.03795,
3     .03366,.029877,.02654,.023591,.020982,.018672,.016625,.01481,
4     .0132,.01177,.010499,.009369,.008362,.007468,.006671,.005961,
5     .005329,.004764,.004261,.003812,.003411,.003053,.0027332,
6     .0024474,.002192/
C
C      END

```

Subroutine AKDAT (for use with AK1)

C <RFISHER>AKDAT.F4;2
BLOCK DATA

1-NOV-77 10:00:30

EDIT BY RFISHER

USED BY SUBROUTINE AK, ATMOSPHERIC TURBULENCE
VALUES OF AK1 ARE FOR MODIFIED BESSEL FNS; $(2/\pi)*K(1/3)$
FOR VALUES OF X FROM .1 TO 10.0 IN .1 STEPS
VALUES OF AK2 ARE FOR MODIFIED BESSEL FNS; $(2/\pi)*K(2/3)$
FOR VALUES OF X AS ABOVE

COMMON/AA/AK1(100),AK2(100)

DATA AK1/1.8461,1.2601,.9607,.7676,.6296,.5253,.4434,
1 ,3776,.3230,.2791,.2416,.2100,.1830,.160,.14016,.12302,
2 ,10818,.09327,.08402,.07419,.06559,.05805,.05142,.04559,
3 ,04045,.03592,.03192,.02838,.025249,.022476,.020018,.017838,
4 ,015902,.014183,.012654,.011295,.010085,.009008,.008049,
5 ,007194,.006432,.005752,.005145,.004604,.004120,.003688,
6 ,003303,.0029578,.0026495,.0023739,.0021273,.0019067,.0017093,
7 ,0015325,.0013743,.0012326,.0011057,.000992,.0008901,
8 ,0007988,.0007169,.0006469,.0005778,.0005188,.0004658,
9 ,0004184,.0003758,.0003375,.0003032,.00027245,.00024481,
A ,00022,.00019772,.00017779,.00015974,.0001436,.0001291,
B ,00011608,.00010438,.00009386,.8441E-04,.7592E-04,.6828E-04,
C ,6141E-4,.5525E-4,.4971E-4,.4472E-4,.4042E-4,.3621E-4,
D ,3258E-4,.29322E-4,.26389E-4,.23751E-4,.21377E-4,.19242E-4,
E ,17321E-4,.15593E-4,.14038E-4,.12639E-4,.11379E-4/

DATA AK2/3.026,1.7837,1.2716,.9681,.7678,.6257,.5187,.4354,
1 ,3688,.3148,.27024,.23312,.20191,.17547,.15294,.13364,.11704,
2 ,1027,.09027,.07947,.07007,.06185,.05466,.04835,.04282,.03795,
3 ,03366,.029877,.02654,.023591,.020982,.018672,.016625,.01481,
4 ,0132,.01177,.010499,.009369,.008362,.007468,.006671,.005961,
5 ,005329,.004764,.004261,.003812,.003411,.003053,.0027332,
6 ,0024474,.002192,.0019637,.0017594,.0015767,.0014132,
7 ,0012669,.001136,.0010187,.0009137,.0008196,.0007354,.0006599,
8 ,0005922,.0005315,.0004771,.0004283,.0003846,.0003454,
9 ,0003102,.0002786,.00025026,.00022483,.000202,.00018151,
A ,00016312,.0001466,.00013176,.00011844,.00010647,.00009572,
B ,8607E-4,.7739E-4,.6959E-4,.6259E-4,.5629E-4,.5063E-4,.4554E-4,
C ,4097E-4,.3686E-4,.3316E-4,.29837E-4,.26847E-4,.24159E-4,
D ,21741E-4,.19566E-4,.1761E-4,.15851E-4,.14268E-4,.12843E-4,
E ,11562E-4/

END

Subroutine ANRP1


```

C      INTERPOLATION ROUTINE
C
C      SUBROUTINE ANTRP(JMAX,AKNWN,AKJ,NDX,AL,ALINT)
C
C      DIMENSION AKJ(0/11),AL(0/11)
C
C      IF (NDX.EQ.JMAX) GO TO 10
C
C      ALINT = (AL(NDX+1)*(AKNWN-AKJ(NDX))-AL(NDX)*(AKNWN-AKJ(NDX+1)))
1      /((AKJ(NDX+1)-AKJ(NDX)))
C      GO TO 20
10     TYPE 5,JSIG
5      FORMAT(' JSIG = JMAX, INTERPOLATION OUT OF RANGE')
20     RETURN
C
C      END

```

Program ATUR4A

```

C      PROGRAM ATURB4.  ATMOSPHERIC TURBULENCE TASK
C      ITEM 3, PARTS 1,1.A. PHASE II OF ATMOSPHERIC TURBULENCE
C      MEAN VALUE SUBTRACTED FROM DATA BEFORE COMPUTING SPECTRUM
C      SPECTRUM CALCULATIONS (AS IN PHASE I)
C
C      REQUIRES SUBROUTINES:
C          CFFT: TO PERFORM FFT
C          SIMP: TO INTEGRATE BY SIMPSONS RULE
C          HPDES: HIGH PASS FILTER
C
C      READS NASA DATA STORED IN DATA FILE TO BE NAMED
C      PRODUCES DATA FILES:
C          PHLK: VALUES OF PSD PHI OF L(K)
C          AUT7: AUTOCORRELATION OF PSD
C
C      COMMON/53/W(2/65537)
C      COMMON/CC/A(3),B(3),C(3),GR(2,10),F(4,3)
C
C      INPUT INITIAL PARAMETERS
C
C      TYPE 802
802    FORMAT(/1X,'INPUT SPEED OF CRAFT (M/SEC)',/$)
      ACCEPT 801,V
801    FORMAT(G)
      TYPE 802
802    1  FORMAT(/1X,'INPUT TOTAL NO. OF POINTS TO BE USED IN 2L M.',
      ' OF DATA',/1X,'AND POWER OF TWO OF THAT NO.',/$)
      ACCEPT 803,NPTS
      ACCEPT 803,MPWRN
803    FORMAT(I6)
      TYPE 702
      ACCEPT 825,NPTS,SRATE
825    FORMAT(2G)
702    1  FORMAT(/1X,'INPUT NO. OF POINTS OF W(X) TO BE READ',/1X,
      'AND SAMPLING RATE OF DATA ',/$)
      TYPE 804
804    1  FORMAT(/1X,'INPUT VALUE OF NPTS',
      /,' AND POWER OF TWO OF THAT NO.',/$)
      ACCEPT 803,NPTS
      ACCEPT 803,MPWRM
C
      NZERO = NPTS-NPTS
      NLAST = NPTS-1
      TIME = SRATE*NPTS
      TWOL = V*TIME
      PTS = FLOAT(NPTS)
      DELX = TWOL/PTS
      FTM = NZERO*DELX
      FIDAT = TWOL-FTM
      DELK = 1./TWOL
      FTM1 = NPTS*DELX

```

```

C
C      READ IN THE DATA W(X) AND CONVERT TO M/SEC
C
      TYPE 805
806    FORMAT(/1X,'INPUT DATA FILE NAME THAT CONTAINS SAMPLES',
1      /1X,'OF W(X) AND NO. POINTS NXRAY',/$)
      ACCEPT 807,FLE
      ACCEPT 825,NXRAY
807    FORMAT(A5)

      CALL IFILE(20,FLE)
      NX1 = NXRAY/4
      DO 550 I=0,NX1-1
      READ(20,551) W(I),W(I+NX1),W(I+2*NX1),W(I+3*NX1)
      W(I+NX1) = W(I+NX1)*.3048
      W(I+2*NX1) = W(I+2*NX1)*.3048
      W(I+3*NX1) = W(I+3*NX1)*.3048
550    W(I) = W(I)*.3048
551    FORMAT(4(E15.7))
      END FILE 20

C
C      HIGH PASS FILTER THE DATA
C
      TYPE 9
9      FORMAT(' INPUT CUT-OFF FREQ,SAMPLING INTERVAL,'
1      'NO. OF FILTER SECTIONS',/$)
      ACCEPT 2,FC,TS,NS
2      FORMAT(3G)
      FK = FC/V
      CALL HPDES(FC,TS,NS)
      DO 140 N=1,NS+1
      DO 140 M=1,2
140    F(N,M) = 0.0
      DO 150 M=0,NXRAY-1
      F(1,3) = W(M)
      DO 160 N=1,NS
      TEMP = A(N)*(F(N,3)-2.*F(N,2)+F(N,1))
160    F(N+1,3) = TEMP-B(N)*F(N+1,2)-C(N)*F(N+1,1)
      DO 170 N=1,NS+1
      DO 170 MM=1,2
170    F(N,MM) = F(N,MM+1)
150    W(M) = F(NS+1,3)

C
C      COMPUTE AND SUBTRACT OUT MEAN VALUE
C      COMPUTE THE SAMPLE VARIANCE
C
      WBAR = 0.0
      VAR = 0.0
      DO 610 JJ = 0,NOPTS-1
610    WBAR = WBAR + W(JJ)
      WBAR = WBAR/FLOAT(NOPTS)
      DO 600 I=0,NOPTS-1
      W(I) = W(I) - WBAR

```

```

600      VAR = VAR + W(I)*W(I)
        VAR = VAR/FLDAT(NOPTS)
C
C      SQUARE W(I)
C
        DO 1111 I=0,NXRAY-1
1111     W(I) = W(I)*W(I)
C
C      COMPUTE INTEGRAL OF W(K)**2 USING TRAPIZOIDAL RULE
C
        WSUM = 0.0
        DO 370 I=1,NOPTS-2
370      WSUM = WSUM+W(I)*W(I)
        WSUM = WSUM + .5*W(0)*W(0) + .5*W(NOPTS-1)*W(NOPTS-1)
        WSUM = DELX*WSUM/FTDAT
C
C      ADD ZEROS TO DATA
C
        DO 10 I=NOPTS,NLAST
10       W(I) = 0.0
C
C      MAKE ARRAY COMPLEX
C
        DO 512 J=NPTS-1,0,-1
512      W(2*J) = W(J)
        DO 513 J=1,2*NPTS-1,2
513      W(J) = 0.0
C
C      COMPUTE PHI L(K)
C
        CALL CFFT(MPWRN,NPTS,W,1)
C
        NHALF = NPTS/2
        NMIN1 = NHALF-1
        TMS = TWOL*TWOL/FTDAT
        DO 11 I=0,2*NLAST,2
        W(I+1) = 0.0
        W(I) = W(I)*TMS
11       CONTINUE
        W(2*NPTS-1) = 0.0
C
C      OUTPUT VALUES OF PHI OF L(K) TO DATA FILE (FIRST HALF ONLY)
C
        CALL OFILE (20,'FPSD2')
        WRITE(20,900)
        WRITE(20,950) FC,FK,FL
950      FORMAT(/1X,'HIGH PASS FILTERED DATA',/1X,'CUT-OFF FREQ =' ,G,
1         ' (K =' ,G,') ',/1X,
2         ' FILTERED VERSION SQUARED BEFORE TRANSFORMING',/1X,
3         ' POWER SPECTRUM OF PHI OF L(K) ',/1X,
3         ' DATA TAKEN FROM FILE ',A5)
        WRITE(20,906) NPTS,TWOL,NPTS,FTM1
        WRITE(20,907) MZERO
        WRITE(20,913) WBAR,VAR
        WRITE(20,951) WSUM

```

```

951      FORMAT(/1X,'<W OF L(X)**2> = ',F12.4)
      WRITE(20,952)
952      FORMAT(/1X,'PRINTOUT OF THE VALUES OF THE POWER SPECTRUM',
1        /1X,5X,'K',10X,'PS VALUE',8X,'K CONTD',4X,'PS VALUE CONTD',/)
      MHALF = NPTS/4
      M41 = MHALF-1
      DO 823 I=0,M41
      DEL = DELK*I
      XX = W(2*I)
      K = MHALF+I
      DEL2 = DELK*K
      YY = W(2*K)
823      WRITE(20,903) DEL,XX,DEL2,YY
      DEL3 = DELK*NHALF
      ZZ = W(2*NHALF)
      WRITE(20,908) DEL3,ZZ
      END FILE 20

C
C      PERFORM INTEGRATION CHECK
C
      TYPE 810
810      FORMAT(/1X,'PERFORM INTEGRATION CHECK (Y OR N)      ',5)
      ACCEPT 807,CHK
      IF (CHK.EQ.'N') GO TO 23
      EDR = NPTS*DELK
      CALL SIMP(0.0,EDR,DELK,NPTS,ANS)
      TYPE 811,ANS
811      FORMAT(/1X,'INTEGRAL OF PHI OF L (K) = ',E12.4)
C
C      OBTAIN AUTOCORRELATION FUNCTION
C
23      CALL CFFT(MPWRN,NPTS,W,2)
C
C      TYPE OUT AUTOCORPELATION TO DATA FILE
C
      XAOFF = 13000.
      RFCN = PTS/TWOL
      DR0 = W(2)*RFCN
      MAOFF = 13000./DELX
C
      CALL OFILE (20,'AUTF2')
      WRITE(20,900)
      WRITE(20,1950) FILE
1950      FORMAT(/1X,'AUTOCORRELATION OF',/1X,
1        'HIGH PASS FILTERED AND SQUARED NON-HOMOGENEOUS SAMPLE',/1X,
2        'DATA TAKEN FROM FILE ',A5)
      WRITE(20,906) NPTS,TWOL,MPTS,FTM1
      WRITE(20,907) MZERO
      WRITE(20,910) WBAR,VAR
      WRITE(20,951) WSUM
      WRITE(20,381) XAOFF,MAOFF

```

```

381      FORMAT(/1X,' TRUCATION POINT WAS',G,' METERS',/1X,
1      'WHICH CONTAINS ',I7,' POINTS')
      WRITE(20,1952)
1952      FORMAT(/1X,'PRINTOUT OF THE VALUES OF THE AUTOCORRELATION',
1      /1X,8X,'X',12X,'      RL      ',9X,'RL/R0')
      DO 1823 I=0,NAOFF
      DEL = DELX*I
      XX = W(2*I)*RFFCON
      YY = XX/DR0
1823      WRITE(20,915) DEL,XX,YY
915      FORMAT(1X,3(2X,G))
      END FILE 20

C

900      FORMAT(/1X,'DATA FILE CREATED BY PROGRAM ATURB4')
903      FORMAT(1X,F10.6,3X,E12.4,3X,F10.6,3X,E12.4)
905      FORMAT(/1X,'RL(0) = ',F12.4)
906      FORMAT(/1X,I6,' DATA POINTS WERE USED IN 2L = ',F15.4,
1      ' METER',/1X,I5,' DATA POINTS WERE USED IN M = ',F16.4,
1      ' METER')
907      FORMAT(/1X,I6,' ZEROS WERE ADDED TO DATA')
908      FORMAT(29X,F10.4,3X,E12.4)
910      FORMAT(/1X,'MEAN VALUE OF W(X) = ',E15.5,' M/SEC',/1X,
1      'MEAN SQ. VALUE = ',E15.5,' (M/SEC)**2')

C
C
4999      END

```

Program ATURB2


```

C <RFISHER>ATURB2.F4;5      1-NOV-77 10:14:27      EDIT BY RFISHER
C PROGRAM ATP2.  ATMOSPHERIC TURBULENCE TASK
C ITEM 1, PARTS 1,1.A. PHASE II OF ATMOSPHERIC TURBULENCE
C MEAN VALUE SUBTRACTED FROM DATA BEFORE COMPUTING SPECTRUM
C SPECTRUM CALCULATIONS (AS IN PHASE I)
C
C REQUIRES SUBROUTINES:
C
C       CFPT: TO PERFORM FFT
C       SIMP: TO INTEGRATE BY SIMPSIONS RULE
C
C READS NASA DATA STORED IN DATA FILE TO BE NAMED
C PRODUCES DATA FILES:
C       PHILK: VALUES OF PSD PHI OF L(K)
C       AUTO: AUTOCORRELATION OF PSD
C       DSPS: SMOOTHED POWER SPECTRUM
C
C COMMON/SG/W(0/65537)
C
C INPUT INITIAL PARAMETERS
C
C TYPE 800
800  FORMAT(/1X,'INPUT SPEED OF CRAFT (M/SEC)',/S)
      ACCEPT 801,V
801  FORMAT(G)
      TYPE 802
802  FORMAT(/1X,'INPUT TOTAL NO. OF POINTS TO BE USED IN 2L M.',
1     ' OF DATA',/1X,'AND POWER OF TWO OF THAT NO.',/S)
      ACCEPT 803,NPTS
      ACCEPT 803,MPWRN
803  FORMAT(I6)
      TYPE 702
      ACCEPT 825,NOPTS,SRATE
825  FORMAT(2G)
702  FORMAT(/1X,'INPUT NO. OF POINTS OF W(X) TO BE READ',/1X,
1     'AND SAMPLING RATE OF DATA ',/S)
      TYPE 804
804  FORMAT(/1X,'INPUT VALUE OF MPTS',
1     '/, ' AND POWER OF TWO OF THAT NO.',/S)
      ACCEPT 803,MPTS
      ACCEPT 803,MPWRM
C
      MZERO = NPTS-NOPTS
      NLAST = NPTS-1
      TIME = SRATE*NPTS
      TWOL = V*TIME
      PTS = FLOAT(NPTS)
      DELX = TWOL/PTS
      FTM = MZERO*DELX
      FTDAT = TWOL-FTM
      DELK = 1./TWOL
      FTM1 = MPTS*DELX

```

```

C
C      READ IN THE DATA W(X) AND CONVERT TO M/SEC
C
      TYPE 806
806    1  FORMAT(/1X,'INPUT DATA FILE NAME THAT CONTAINS SAMPLES',
      /1X,'OF W(X) AND NO. POINTS NXRAY',/8)
      ACCEPT 807,FLE
      ACCEPT 825,NXRAY
807    FORMAT(A5)

      CALL IFILE(20,FLE)
      NX1 = NXRAY/4
      DO 550 I=0,NX1-1
      READ(20,551)W(I),W(I+5700),W(I+11400),W(I+17100)
      W(I+5700) = W(I+5700)*.3048
      W(I+11400) = W(I+11400)*.3048
      W(I+17100) = W(I+17100)*.3048
550    W(I) = W(I)*.3048
551    FORMAT(4(E15.7))
      END FILE 20

C
C      COMPUTE AND SUBTRACT OUT MEAN VALUE
C      COMPUTE THE SAMPLE VARIANCE
C
      WBAR = 0.0
      VAR = 0.0
      DO 610 JJ = 0,NOPTS-1
610    WBAR = WBAR + W(JJ)
      WBAR = WBAR/FLOAT(NOPTS)
      DO 600 I=0,NOPTS-1
      W(I) = W(I) - WBAR
600    VAR = VAR + W(I)*W(I)
      VAR = VAR/FLOAT(NOPTS)

C
C      COMPUTE INTEGRAL OF W(K)**2 USING TRAPIZOIDAL RULE
C
      WSUM = 0.0
      DO 370 I=1,NOPTS-2
370    WSUM = WSUM+W(I)*W(I)
      WSUM = WSUM + .5*W(0)*W(0)+.5*W(NOPTS-1)*W(NOPTS-1)
      WSUM = DELX*WSUM/FTDAT

C
C      ADD ZEROS TO DATA
C
      DO 10 I=NOPTS,NLAST
10    W(I) = 0.0

C
C      MAKE ARRAY COMPLEX
C
      DO 512 J=NPTS-1,0,-1

```

```

512      W(2*J) = W(J)
      DO 513 J=1,2*NPTS-1,2
513      W(J) = 0.0
C
C      COMPUTE PHI L(K)
C
      CALL CFFT(MPWRN,NPTS,W,1)

      NHALF = NPTS/2
      NMIN1 = NHALF-1
      TMS = TWOL*TWOL/FTDAT
      DO 11 I=0,2*NLAST,2
      W(I+1) = 0.0
      W(I) = W(I)*TMS
11      CONTINUE
      W(2*NPTS-1) = 0.0
C
C      OUTPUT VALUES OF PHI OF L(K) TO DATA FILE (FIRST HALF ONLY)
C
      CALL OFILE (20,'PHILK')
      WRITE(20,900)
      WRITE(20,950)FLE
950      1  FORMAT(/1X,'POWER SPECTRUM OF PHI OF L(K)',/1X,
      'DATA TAKEN FROM FILE ',A5)
      WRITE(20,906)NPTS,TWOL,MPTS,FTM1
      WRITE(20,907)MZERO
      WRITE(20,910)WBAR,VAR
      WRITE(20,951)WSUM
951      FORMAT(/1X,'<W OF L(X)**2> = ',F12.4)
      WRITE(20,952)
952      1  FORMAT(/1X,'PRINTOUT OF THE VALUES OF THE POWER SPECTRUM',
      /1X,5X,'K',10X,'PS VALUE',8X,'K CONTD',4X,'PS VALUE CONTD',/)
      MHALF = NPTS/4
      M41 = MHALF-1
      DO 823 I=0,M41
      DEL = DELK*I
      XX = W(2*I)
      K = MHALF+I
      DEL2 = DELK*K
      YY = W(2*K)
823      WRITE(20,903)DEL,XX,DEL2,YY
      DEL3 = DELK*MHALF
      ZZ = W(2*MHALF)
      WRITE(20,908)DEL3,ZZ
      END FILE 20
C
C      PERFORM INTEGRATION CHECK
C
      TYPE 810

```

```

810  FORMAT(/1X,'PERFORM INTEGRATION CHECK (Y OR N) ',8)
      ACCEPT 807,CHK
      IF (CHK.EQ.'N') GO TO 23
      EDR = NPTS*DELK
      CALL SIMP(0.0,EDR,DELK,NPTS,ANS)
      TYPE 811,ANS
811  FORMAT(/1X,'INTEGRAL OF PHI OF L (K) =' ,E12.4)
C
C      OBTAIN AUTOCORRELATION FUNCTION
      TYPE 4444,W(0),W(1),W(2),W(3)
      TYPE 4444,W(4),W(5),W(6),W(7)
      TYPE 4444,W(8),W(10)
      TYPE 4444,W(NPTS)
      TYPE 4444,W(NPTS+2),W(NPTS+2)
      TYPE 4444,W(2*NLAST-2),W(2*NLAST-4),W(2*NLAST-6)
4444  FORMAT(4G)
C
23  CALL CFFT(MPWRN,NPTS,W,2)
C
C      TYPE OUT AUTOCORRELATION TO DATA FILE
C
      XAOFF = 13000.
      RFCON = PTS/TWOL
      DR0 = W(0)*RFCON
      MAOFF = 13000./DELX
C
      CALL OFILE (20,'AUTO')
      WRITE(20,900)
      WRITE(20,1950)FLE
1950  FORMAT(/1X,'AUTOCORRELATION OF STATIONARY SAMPLE',/1X,
1  'DATA TAKEN FROM FILE ',A5)
      WRITE(20,906)NPTS,TWOL,MPTS,FTM1
      WRITE(20,907)MZERO
      WRITE(20,910)WBAR,VAR
      WRITE(20,951)WSUM
      WRITE(20,381)XAOFF,MAOFF
381  FORMAT(/1X,'TRUCATION POINT WAS',G,' METERS',/1X,
1  'WHICH CONTAINS ',I7,' POINTS')
      WRITE(20,1952)
1952  FORMAT(/1X,'PRINTOUT OF THE VALUES OF THE AUTOCORRELATION',
1  /1X,8X,'X',12X,' RL ',9X,'RL/R0')
      DO 1823 I=0,MAOFF
      DEL = DELX*I
      XX = W(2*I)*RFCON
      YY = XX/DR0
1823  WRITE(20,915)DEL,XX,YY
915  FORMAT(1X,3(2X,G))
      END FILE 20
C
C      COMPUTE WINDOW FN. AND RL(Z)/(1=Z/(2L-M))

```

```

C      DR0 = W(0)/DELX
      PI = 3.141592654
      FTM1 = MPTS*DELX
      DO 13 J=0,MPTS
      ARG = DELX*J
      CC = PI*ARG/FTM1
      DD = 1.-ARG/FTM1
      EE = ABS(SIN(CC))
      WINDO = EE/PI+DD*COS(CC)
      DVSOR = 1.-ARG/FTDAT
13     W(2*J) = W(2*J)*WINDO/(DELX*DVSOR)
      M2PWR = MPWRM+1
      M2PTS = 2*MPTS
      FT2M = 2.*FTM1
      MMIN1 = MPTS-1
      DO 14 JK=1,MMIN1
      KK = M2PTS-JK
14     W(2*KK) = W(2*JK)
      DO 521 K=1,2*M2PTS-1,2
521    W(K) = 0.0
C
C      COMPUTE SMOOTHED POWER SPECTRUM
C
      CALL CFFT(M2PWR,M2PTS,W,2)
C
C      PRINT DATA FILE 'DSPS' , FIRST HALF OF DATA POINTS
C
      CALL OFILE(20,'DSPS')
      WRITE(20,900)
      WRITE(20,901)FLE
      WRITE(20,906)NPTS,TWOL,MPTS,FTM1
      WRITE(20,910)WBAR,VAR
      WRITE(20,951)WSUM
      WRITE(20,907)MZERO
      WRITE(20,905)DR0
      WRITE(20,902)
      MHALF = MPTS/2
      M41 = MHALF-1
      DELK = 1./FT2M
      DO 56 I=0,M41
      DEL = DELK*I
      XX = W(2*I)*FT2M
      K = MHALF+I
      DEL2 = DELK*K
      YY = W(2*K)*FT2M
56     WRITE(20,903)DEL,XX,DEL2.YY
      DEL3 = DELK*MPTS
      ZZ = W(2*MPTS)*FT2M
      WRITE(20,908)DEL3,ZZ
      END FILE 20

```

```

900      FORMAT(//1X,'DATA FILE CREATED BY PROGRAM ATURB2')
901      FORMAT(//1X,'SMOOTHED POWER SPECTRUM PHI OF P(K)',
1        /1X,'DATA TAKEN FROM DATA FILE ',A5)
902      FORMAT(//1X,'PRINTOUT OF VALUES OF THE SMOOTHED',
1        ' POWER SPECTRUM',/5X,'K',10X,'SPS VALUE',8X,'K CONTD',
2        4X,'SPS VALUE CONTD',/)
903      FORMAT(1X,F10.6,3X,E12.4,3X,F10.6,3X,E12.4)
905      FORMAT(//1X,'RL(0) = ',E12.4)
906      FORMAT(//1X,I6,' DATA POINTS WERE USED IN 2L = ',F15.4,
1        ' METER',/1X,I5,' DATA POINTS WERE USED IN M = ',F16.4,
1        ' METER')
907      FORMAT(//1X,I6,' ZEROS WERE ADDED TO DATA')
908      FORMAT(29X,F10.4,3X,E12.4)
910      FORMAT(//1X,'MEAN VALUE OF W(X) = ',E15.5,' M/SEC',/1X,
1        'MEAN SQ. VALUE = ',E15.5,' (M/SEC)**2')
C
C
4999      END

```

Program ATURB3

```

C      PROGRAM ATURB2.  ATMOSPHERIC TURBULENCE TASK
C      ITEM 2, PARTS 1,1,A.  PHASE II OF ATMOSPHERIC TURBULENCE
C      MEAN VALUE SUBTRACTED FROM DATA BEFORE COMPUTING SPECTRUM
C      SPECTRUM CALCULATIONS (AS IN PHASE I)
C
C      REQUIRES SUBROUTINES:
C
C          CFFT: TO PERFORM FFT
C          SIMP: TO INTEGRATE BY SIMPSIONS RULE
C
C      READS NASA DATA STORED IN DATA FILE TO BE NAMED
C      PRODUCES DATA FILES:
C          PHILK: VALUES OF PSD PHI OF L(K)
C          AUTO: AUTOCORRELATION OF PSD
C          DSPS: SMOOTHED POWER SPECTRUM
C
C      COMMON/SG/W(0/65537)
C
C      INPUT INITIAL PARAMETERS
C
C      TYPE 800
800    FORMAT(/1X,'INPUT SPEED OF CRAFT (M/SEC)',/8)
      ACCEPT 801,V
801    FORMAT(G)
      TYPE 802
802    FORMAT(/1X,'INPUT TOTAL NO. OF POINTS TO BE USED IN 2L M.',
1      ' OF DATA',/1X,'AND POWER OF TWO OF THAT NO.',/8)
      ACCEPT 803,NPTS
      ACCEPT 803,MPWRN
803    FORMAT(I6)
      TYPE 702
      ACCEPT 825,NOPTS,SRATE
825    FORMAT(2G)
702    FORMAT(/1X,'INPUT NO. OF POINTS OF W(X) TO BE READ',/1X,
1      'AND SAMPLING RATE OF DATA ',/8)
      TYPE 804
804    FORMAT(/1X,'INPUT VALUE OF MPTS',
1      '/, AND POWER OF TWO OF THAT NO.',/8)
      ACCEPT 803,MPTS
      ACCEPT 803,MPWRM
C
      MZERO = NPTS-NOPTS
      NLAST = NPTS-1
      TIME = SRATE*NPTS
      TWOL = V*TIME
      PTS = FLOAT(NPTS)
      DELX = TWOL/PTS
      FTM = MZERO*DELX
      FTDAT = TWOL-FTM
      DELK = 1./TWOL
      FTM1 = MPTS*DELX

```



```

C
C      READ IN THE DATA W(X) AND CONVERT TO M/SEC
C
      TYPE 806
806    1  FORMAT(/1X,'INPUT DATA FILE NAME THAT CONTAINS SAMPLES',
      1  /1X,'OF W(X) AND NO. POINTS NXRAY',/S)
      ACCEPT 807,FLE
      ACCEPT 825,NXRAY
807    FORMAT(A5)

      CALL IFILE(20,FLE)
      NX1 = NXRAY/4
      DO 550 I=0,NX1-1
      READ(20,551)W(I),W(I+NX1),W(I+2*NX1),W(I+3*NX1)
      W(I+NX1) = W(I+NX1)*.3048
      W(I+2*NX1) = W(I+2*NX1)*.3048
      W(I+3*NX1) = W(I+3*NX1)*.3048
550    W(I) = W(I)*.3048
551    FORMAT(4(E15.7))
      END FILE 20

C
C      COMPUTE AND SUBTRACT OUT MEAN VALUE
C      COMPUTE THE SAMPLE VARIANCE
C
      WBAR = 0.0
      VAR = 0.0
      DO 610 JJ = 0,NOPTS-1
610    WBAR = WBAR + W(JJ)
      WBAR = WBAR/FLOAT(NOPTS)
      DO 600 I=0,NOPTS-1
      W(I) = W(I) - WBAR
600    VAR = VAR + W(I)*W(I)
      VAR = VAR/FLOAT(NOPTS)

C
C      COMPUTE INTEGRAL OF W(K)**2 USING TRAPIZOIDAL RULE
C
      WSUM = 0.0
      DO 370 I=1,NOPTS-2
370    WSUM = WSUM+W(I)*W(I)
      WSUM = WSUM + .5*W(0)*W(0)+.5*W(NOPTS-1)*W(NOPTS-1)
      WSUM = DELX*WSUM/FTDAT

C
C      ADD ZEROS TO DATA
C
      DO 10 I=NOPTS,NLAST
10    W(I) = 0.0

C
C      MAKE ARRAY COMPLEX
C
      DO 512 J=NPTS-1,0,-1

```

```

512      W(2*J) = W(J)
      DO 513 J=1,2*NPTS-1,2
513      W(J) = 0.0
C
C      COMPUTE PHI L(K)
C
      CALL CFFT(MPWRN,NPTS,W,1)

      NHALF = NPTS/2
      NMIN1 = NHALF-1
      TMS = TWOL*TWOL/FTDAT
      DO 11 I=0,2*NLAST,2
      W(I+1) = 0.0
      W(I) = W(1)*TMS
11      CONTINUE
      W(2*NPTS-1) = 0.0
C
C      OUTPUT VALUES OF PHI OF L(K) TO DATA FILE (FIRST HALF ONLY)
C
      CALL OFILE (20,'PHILK')
      WRITE(20,900)
      WRITE(20,950)FLE
950      FORMAT(/1X,'POWER SPECTRUM OF PHI OF L(K)',/1X.
1      'DATA TAKEN FROM FILE ',A5)
      WRITE(20,906)NPTS,TWOL,MPTS,FTM1
      WRITE(20,907)MZERO
      WRITE(20,910)WBAR,VAR
      WRITE(20,951)WSUM
951      FORMAT(/1X,'<W OF L(X)**2> = ',F12.4)
      WRITE(20,952)
952      FORMAT(/1X,'PRINTOUT OF THE VALUES OF THE POWER SPECTRUM',
1      /1X,5X,'K',10X,'PS VALUE',8X,'K CONTD',4X,'PS VALUE CONTD',/)
      MHALF = NPTS/4
      M41 = MHALF-1
      DO 823 I=0,M41
      DEL = DELK*I
      XX = W(2*I)
      K = MHALF+I
      DEL2 = DELK*K
      YY = W(2*K)
823      WRITE(20,903)DEL,XX,DEL2,YY
      DEL3 = DELK*NHALF
      ZZ = W(2*NHALF)
      WRITE(20,908)DEL3,ZZ
      END FILE 20
C
C      ABOVE WAS DESCRIPTIVE DATA FILE NOW MAKE DATA FILE THAT
C      IS READ BY OTHER PROGRAMS
C
      CALL OFILE(21,'PSD')
      DO 1010 JK=1,32
      K = 500*(JK-1)

```

```

1010 WRITE(21,9606)(W(2*I),I=0+K,499+K)
9606 FORMAT(500G)
WRITE(21,9606)(W(2*I),I=16000,16384)
END FILE 21

C
C
C
PERFORM INTEGRATION CHECK

TYPE 810
810 FORMAT(/1X,'PERFORM INTEGRATION CHECK (Y OR N) ',8)
ACCEPT 807,CHK
IF (CHK.EQ.'N') GO TO 23
EDR = NPTS*DELK
CALL SIMP(0.0,EDR,DELK,NPTS,ANS)
TYPE 811,ANS
811 FORMAT(/1X,'INTEGRAL OF PHI OF L (K) =' ,E12.4)

C
C
C
OBTAIN AUTOCORRELATION FUNCTION
TYPE 4444,W(0),W(1),W(2),W(3)
TYPE 4444,W(4),W(5),W(6),W(7)
TYPE 4444,W(8),W(10)
TYPE 4444,W(NPTS)
TYPE 4444,W(NPTS-2),W(NPTS+2)
TYPE 4444,W(2*NLAST-2),W(2*NLAST-4),W(2*NLAST-6)
4444 FORMAT(4G)

C
23 CALL CFFT(MPWRN,NPTS,W,2)

C
C
C
TYPE OUT AUTOCORRELATION TO DATA FILE

XAOFF = 13000.
RFCON = PTS/TWOL
DR0 = W(0)*RFCON
MAOFF = 13000./DELX

C
CALL OFILE (20,'AUTO')
WRITE(20,900)
WRITE(20,1950)FLE
1950 FORMAT(/1X,'AUTOCORRELATION OF STATIONARY SAMPLE',/1X,
1 'DATA TAKEN FROM FILE ',A5)
WRITE(20,906)NPTS,TWOL,MPTS,FTM1
WRITE(20,907)MZERO
WRITE(20,910)WBAR,VAR
WRITE(20,951)WSUM
WRITE(20,381)XAOFF,MAOFF
381 FORMAT(/1X,'TRUNCATION POINT WAS',G,' METERS',/1X,
1 'WHICH CONTAINS ',I7,' POINTS')
WRITE(20,1952)
1952 FORMAT(/1X,'PRINTOUT OF THE VALUES OF THE AUTOCORRELATION',
1 /1X,8X,'X',12X,' RL ',9X,'RL/R0')
DO 1823 I=0,MAOFF
DEL = DELX*I
XX = W(2*I)*RFCON
YY = XX/DR0

```

```

1823 WRITE(20,915)DEL,XX,YY
915  FORMAT(1X,3(2X,G))
    END FILE 20

C
C   ABOVE WAS DES CRIPTIVE DATA FILE NOW MAKE DATA FILE
C   TO BE READ BY FOLLOWING PROGRAMS
C
C*****DONT FORGET THESE NUMBERS MUST BE MULTIPLIED BY RFCON *****
C
    IEND = MAOFF/500
    CALL OFILE(21,'RLNH')
    DO 1011 JK = 1,IEND
    K = 500*(JK-1)
1011  WRITE(21,9606)(W(2*I),I=0+K,499+K)
    KEND = 500*(IEND-1)
    WRITE(21,9606)(W(2*I),I=KEND,MAOFF)

C
C   COMPUTE WINDOW FN. AND RL(Z)/((1-Z)/(2L-M))
C
    DR0 = W(0)/DELX
    PI = 3.141592654
    FTM1 = MPTS*DELX
    DO 13 J=0,MPTS
    ARG = DELX*J
    CC = PI*ARG/FTM1
    DD = 1.-ARG/FTM1
    EE = ABS(SIN(CC))
    WINDO = EE/PI+DD*COS(CC)
    DVSOR = 1.-ARG/FTDAT
13    W(2*J) = W(2*J)*WINDO/(DELX*DVSOR)
    M2PWR = MPWRM+1
    M2PTS = 2*MPTS
    FT2M = 2.*FTM1
    MMIN1 = MPTS-1
    DO 14 JK=1,MMIN1
    KK = M2PTS-JK
14    W(2*KK) = W(2*JK)
    DO 521 K=1,2*M2PTS-1,2
521  W(K) = 0.0

C
C   COMPUTE SMOOTHED POWER SPECTRUM
C
    CALL CFFT(M2PWR,M2PTS,W,2)

C
C   PRINT DATA FILE 'DSPS' , FIRST HALF OF DATA POINTS

```

6

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CALLS OF FILE(20, 'DSPS')
WRITE(20, 907)
WRITE(20, 908) FILE
WRITE(20, 909) HPTS, THOL, NPTS, PTM1
WRITE(20, 910) AEAR, VAR
WRITE(20, 911) WSON
WRITE(20, 907) M/EHO
WRITE(20, 905) DR0
WRITE(22, 907)
MHALF = MPTS/2
N41 = MHALF+1
DELK = 1./PT2M
DO 50 I=0, N41
DEL2 = DELK*I
XX = W(2*I)*PT2M
K = MHALF-I
DEL2 = DELK*K
YY = W(2*K)*PT2M
WRITE(20, 901) DEL, XX, DEL2, YY
DEL3 = DELK*MPTS
ZZ = W(2*MPTS)*PT2M
WRITE(20, 908) DEL3, Z1
END FILE 21

```

56

```

500  FORMAT(//1X, 'DATA FILE CREATED BY PROGRAM ATUR13')
901  FORMAT(//1X, 'SMOOTHED POWER SPECTRUM PHJ OF PCK')
1    /1X, 'DATA TAKEN FROM DATA FILE 1, A5'
902  FORMAT(//1X, 'PRINTOUT OF VALUES OF THE SMOOTHED')
1    ' POWER SPECTRUM', /1X, 'X', 'Y', 'Z', 'SPS VALUE', 'K CONTR',
2    /1X, 'SPS VALUE CONID', /)
903  FORMAT(1X, F10.6, 3X, E12.4, 1X, F10.6, 3X, E12.4)
905  FORMAT(//1X, 'R1(0) = ', E12.4)
906  FORMAT(//1X, 16, ' DATA POINTS WERE USED IN 21 = ', F15.4,
1    ' METER', /1X, 16, ' DATA POINTS WERE USED IN M = ', F16.4,
1    ' METER')
907  FORMAT(//1X, 16, ' ZEROS WERE ADDED TO DATA')
908  FORMAT(25X, F12.4, 3X, E12.4)
910  FORMAT(//1X, 'MEAN VALUE OF W(X) = ', E15.5, ' M/SEC', /1X,
1    'MEAN SQ. VALUE = ', E15.5, ' (M/SEC)*2')

```

6

12

4999

END

Program ATURB4

```

C      PROGRAM ATURB4.  ATMOSPHERIC TURBULENCE TASK
C      ITEM 3, PARTS 1,1,A. PHASE II OF ATMOSPHERIC TURBULENCE
C      MEAN VALUE SUBTRACTED FROM DATA BEFORE COMPUTING SPECTRUM
C      SPECTRUM CALCULATIONS (AS IN PHASE I)
C
C      REQUIRES SUBROUTINES:
C          CFFT; TO PERFORM FFT
C          SIMP; TO INTEGRATE BY SIMPSIONS RULE
C          HPDES; HIGH PASS FILTER
C
C      READS NASA DATA STORED IN DATA FILE TO BE NAMED
C      PRODUCES DATA FILES:
C          PHILK; VALUES OF PSD PHI OF L(K)
C          AUTO; AUTOCORRELATION OF PSD
C
C      COMMON/SG/W(0/65537)
C      COMMON/CC/A(3),B(3),C(3),GR(2,10),F(4,3)
C
C      INPUT INITIAL PARAMETERS
C
C      TYPE 800
800    FORMAT(/1X,'INPUT SPEED OF CRAFT (M/SEC)',/s)
      ACCEPT 801,V
801    FORMAT(G)
      TYPE 802
802    1  FORMAT(/1X,'INPUT TOTAL NO. OF POINTS TO BE USED IN 2L M.',
      ' OF DATA',/1X,'AND POWER OF TWO OF THAT NO.',/s)
      ACCEPT 803,NPTS
      ACCEPT 803,MPWRN
803    FORMAT(I6)
      TYPE 702
      ACCEPT 825,NOPTS,SRATE
825    FORMAT(2G)
702    1  FORMAT(/1X,'INPUT NO. OF POINTS OF W(X) TO BE READ',/1X,
      ' AND SAMPLING RATE OF DATA ',s)
      TYPE 804
804    1  FORMAT(/1X,'INPUT VALUE OF MPTS',
      '/, AND POWER OF TWO OF THAT NO.',/s)
      ACCEPT 803,MPTS
      ACCEPT 803,MPWRM
C
      MZERO = NPTS-NOPTS
      NLAST = NPTS-1
      TIME = SRATE*NPTS
      TWOL = V*TIME
      PTS = FLOAT(NPTS)
      DELX = TWOL/PTS
      FTM = MZERO*DELX
      FTDAT = TWOL-FTM
      DELK = 1./TWOL
      FTM1 = MPTS*DELX

```

```

C
C      READ IN THE DATA W(X) AND CONVERT TO M/SEC
C
      TYPE 806
806      FORMAT(/1X,'INPUT DATA FILE NAME THAT CONTAINS SAMPLES',
1        /1X,'OF W(X) AND NO. POINTS NXRAY',/S)
      ACCEPT 807,FLE
      ACCEPT 825,NXRAY
807      FORMAT(A5)

      CALL IFILE(20,FLE)
      NX1 = NXRAY/4
      DO 550 I=0,NX1-1
      READ(20,551)W(I),W(I+NX1),W(I+2*NX1),W(I+3*NX1)
      W(I+NX1) = W(I+NX1)*.3048
      W(I+2*NX1) = W(I+2*NX1)*.3048
      W(I+3*NX1) = W(I+3*NX1)*.3048
550      W(I) = W(I)*.3048
551      FORMAT(4(E15.7))
      END FILE 20

C
C      HIGH PASS FILTER THE DATA
C
      TYPE 9
9      FORMAT(' INPUT CUT-OFF FREQ,SAMPLING INTERVAL,
1        'NO. OF FILTER SECTIONS',S)
      ACCEPT 2,FC,TS,NS
2      FORMAT(3G)
      FK = FC/V
      CALL HPDES(FC,TS,NS)
      DO 140 N=1,NS+1
      DO 140 M=1,2
140      F(N,M) = 0.0
      DO 150 M=0,NXRAY-1
      F(1,3) = W(M)
      DO 160 N=1,NS
      TEMP = A(N)*(F(N,3)+2.*F(N,2)+F(N,1))
160      F(N+1,3) = TEMP-B(N)*F(N+1,2)+C(N)*F(N+1,1)
      DO 170 N=1,NS+1
      DO 170 MM=1,2
170      F(N,MM) = F(N,MM+1)
150      W(M) = F(NS+1,3)

C
C      COMPUTE AND SUBTRACT OUT MEAN VALUE
C      COMPUTE THE SAMPLE VARIANCE
C
      WBAR = 0.0
      VAR = 0.0
      DO 610 JJ = 0,NOPTS-1
610      WBAR = WBAR + W(JJ)
      WBAR = WBAR/FLOAT(NOPTS)
      DO 600 I=0,NOPTS-1
      W(I) = W(I) - WBAR

```



```

600    VAR = VAR + W(I)*W(I)
      VAR = VAR/FLOAT(NOPTS)
C
C    COMPUTE INTEGRAL OF W(K)**2 USING TRAPIZOIDAL RULE
C
      WSUM = 0.0
      DO 370 I=1,NOPTS-2
370    WSUM = WSUM+W(I)*W(I)
      WSUM = WSUM + .5*W(0)*W(0)+.5*W(NOPTS-1)*W(NOPTS-1)
      WSUM = DELX*WSUM/FTDAT
C
C    ADD ZEROS TO DATA
C
      DO 10 I=NOPTS,NLAST
10    W(I) = 0.0
C
C    MAKE ARRAY COMPLEX
C
      DO 512 J=NPTS-1,0,-1
512    W(2*J) = W(J)
      DO 513 J=1,2*NPTS-1,2
513    W(J) = 0.0
C
C    COMPUTE PHI L(K)
C
      CALL CFFT(MPWRN,NPTS,W,1)

      NHALF = NPTS/2
      NMIN1 = NHALF-1
      TMS = TWOL*TWOL/FTDAT
      DO 11 I=0,2*NLAST,2
      W(I+1) = 0.0
      W(I) = W(I)*TMS
11    CONTINUE
      W(2*NPTS-1) = 0.0
C
C    OUTPUT VALUES OF PHI OF L(K) TO DATA FILE (FIRST HALF ONLY)
C
      CALL OFILE (20,'PHILK')
      WRITE(20,900)
      WRITE(20,950)FC,FK,FLE
950    FORMAT(/1X,'HIGH PASS FILTERED DATA',/1X,'CUT-OFF FREQ =',G,
1    ' (K =',G,')',/1X,
2    'POWER SPECTRUM OF PHI OF L(K)',/1X,
3    'DATA TAKEN FROM FILE ',A5)
      WRITE(20,906)NPTS,TWOL,MPTS,FTM1
      WRITE(20,907)MZERO
      WRITE(20,910)WBAR,VAR
      WRITE(20,951)WSUM
951    FORMAT(/1X,'<W OF L(X)**2> = ',F12.4)
      WRITE(20,952)

```

```

952      FORMAT(/1X,'PRINTOUT OF THE VALUES OF THE POWER SPECTRUM',
1        /1X,5X,'K',10X,'PS VALUE',8X,'K CONTD',4X,'PS VALUE CONTD',/)
      MHALF = NPTS/4
      M41 = MHALF-1
      DO 823 I=0,M41
      DEL = DELK*I
      XX = W(2*I)
      K = MHALF+I
      DEL2 = DELK*K
      YY = W(2*K)
823      WRITE(20,903)DEL,XX,DEL2,YY
      DEL3 = DELK*NHALF
      ZZ = W(2*NHALF)
      WRITE(20,908)DEL3,ZZ
      END FILE 20

C
C      PERFORM INTEGRATION CHECK
C
      TYPE 810
810      FORMAT(/1X,'PERFORM INTEGRATION CHECK (Y OR N)      ',S)
      ACCEPT 807,CHK
      IF (CHK.EQ.'N') GO TO 23
      EDR = NPTS*DELK
      CALL SIMP(0,0,EDR,DELK,NPTS,ANS)
      TYPE 811,ANS
811      FORMAT(/1X,'INTEGRAL OF PHI OF L (K) =',E12.4)

C
C      OBTAIN AUTOCORRELATION FUNCTION
C
23      CALL CFFT(MPWRN,NPTS,W,2)
C
C      TYPE OUT AUTOCORRELATION TO DATA FILE
C
      XAOFF = 13000,
      RFCON = PTS/TWOL
      DR0 = W(0)*RFCON
      MAOFF = 13000./DELX

C
      CALL OFILE (20,'AUTO')
      WRITE(20,900)
      WRITE(20,1950)FLE
1950      FORMAT(/1X,'AUTOCORRELATION OF',/1X,
1        'HIGH PASS FILTERED NON-HOMOGENEOUS SAMPLE',/1X,
2        'DATA TAKEN FROM FILE ',A5)
      WRITE(20,906)NPTS,TWOL,MPTS,FTM1
      WRITE(20,907)MZERO
      WRITE(20,910)WBAR,VAR
      WRITE(20,951)WSUM
      WRITE(20,381)XAOFF,MAOFF
381      FORMAT(/1X,'TRUCATION POINT WAS',G,' METERS',/1X,
1        'WHICH CONTAINS ',I7,' POINTS')
      WRITE(20,1952)

```

```

1952  FORMAT(//1X,'PRINTOUT OF THE VALUES OF THE AUTOCORRELATION',
1      /1X,8X,'X',12X,' RL ',9X,'RL/R0')
      DO 1823 I=0,MAOFF
      DEL = DELX*I
      XX = W(2*I)*RFCON
      YY = XX/DR0
1823  WRITE(20,915)DEL,XX,YY
915   FORMAT(1X,3(2X,G))
      END FILE 20
C
900   FORMAT(//1X,'DATA FILE CREATED BY PROGRAM ATURB4')
903   FORMAT(1X,F10.6,3X,E12.4,3X,F10.6,3X,E12.4)
905   FORMAT(//1X,'RL(0) = ',E12.4)
906   FORMAT(//1X,I6,' DATA POINTS WERE USED IN 2L = ',F15.4,
1      ' METER',/1X,I5,' DATA POINTS WERE USED IN M = ',F16.4,
1      ' METER')
907   FORMAT(//1X,I6,' ZEROS WERE ADDED TO DATA')
908   FORMAT(29X,F10.4,3X,E12.4)
910   FORMAT(//1X,'MEAN VALUE OF W(X) = ',E15.5,' M/SEC',/1X,
1      'MEAN SQ. VALUE = ',E15.5,' (M/SEC)**2')
C
C
4999  END

```

Subroutine BIN

```

C      SUBROUTINE BIN(S)
C      PROGRAM BIN.  ATMOSPHERIC TURBULENCE TASK
C      ITEM 4

COMMON/AA/BIN1(501),BIN2(501),TOTAL,NBIN,BINW,NPTS
DIMENSION S(0/15000)

TYPE 798
FORMAT(/1X,' INPUT BIN WIDTH',/1X,' NO. OF BINS TOTAL',/S)
ACCEPT 744,BINW
ACCEPT 744,NBIN
744    FORMAT(G)

NBIN2 = NBIN/2
DO 10 J=1,NBIN2
BIN1(J) = 0.0
10    BIN2(J) = 0.0
TOTAL = 0.0
TBIN = 0.0

DO 11 J= 0,NPTS-1
IF (S(J).LT.0.0) GO TO 50
IF (S(J).EQ.0.0) BIN1(1) = BIN1(1)+1.
IF (S(J).EQ.0.0) GO TO 11
RSLT1 = S(J)/BINW
IR = IFIX(RSLT1) + 1
BIN1(IR) = BIN1(IR) + 1.
GO TO 11
50    RSLT2 = ABS(S(J))/BINW
IR = IFIX(RSLT2) + 1
BIN2(IR) = BIN2(IR) + 1.
11    CONTINUE

TOTAL = 0.0
DO 13 J=1,NBIN2
13    TOTAL = TOTAL + BIN1(J) + BIN2(J)
TYPE 802,TOTAL
802    FORMAT(/1X,'TOTAL NO. OF POINTS USED FOR 1ST PASS =',F8.2)

C
C      RETURN
C
C      END

```

Subroutine BINSQ

```

SUBROUTINE BINSQ(S)
C PROGRAM BINSQ.  ATMOSPHERIC TURBULENCE TASK
C ITEM 4
C
COMMON/BB/BIN3(700),TOT1,NBIN1,BINW1,NPTS1
DIMENSION S(0/15000)

TYPE 798
798 FORMAT(' INPUT BIN WIDTH ',/1X,' NO. OF BINS',/4)
ACCEPT 744,BINW1
ACCEPT 744,NBIN1
744 FORMAT(G)

DO 10 J=1,NBIN1
10 BIN3(J) = 0.0
TOT1 = 0.0
TBIN = 0.0

DO 11 J= 0,NPTS1-1
IF (S(J).EQ.0.0) BIN3(1) = BIN3(1)+1.
IF (S(J).EQ.0.0) GO TO 11
RSLT = S(J)/BINW1
IR = IFIX(RSLT) + 1
BIN3(IR) = BIN3(IR) + 1.0
11 CONTINUE
TOT1 = 0.0
DO 13 J=1,NBIN1
13 TOT1 = TOT1 + BIN3(J)
TYPE 802,TOT1
802 FORMAT(/1X,'TOTAL NO. OF POINTS USED FOR 1ST PASS =',F8.2)

C
RETURN
C
END

```

Subroutine CFFT1


```

C      SUBROUTINE CFFT.  CALCULATES FFT OF ANY DATA ARRAY
C      NUMBER OF DATA POINTS IS POWER OF 2(M)
C

```

```

      SUBROUTINE CFFT(MPOWR,NPTS,S,NWAY)

```

```

      DIMENSION B(2)
      COMPLEX S(0/5000),U,T,W1
      EQUIVALENCE (W1,B)
      DO 301 I=NPTS,1,-1
      NRAY = I-1
301    S(I) = S(NRAY)
D454   FORMAT(1X,4(E12,4,3X))
D      TYPE 454,S(1),S(2)
D      TYPE 454,S(3),S(4)
D      TYPE 454,S(5),S(6)
D      TYPE 454,S(NPTS/2+1)
D      TYPE 454,S(NPTS/2),S(NPTS/2+2)
D      TYPE 454,S(NPTS),S(NPTS-1),S(NPTS-2)
D      TYPE 454,S(NPTS/2-1),S(NPTS/2-2)
D      TYPE 454,S(NPTS/2+3),S(NPTS/2+4)
C
C

```

```

      NV2 =NPTS/2
      NM1 = NPTS-1
      J = 1
      DS = 1./FLOAT(NPTS)
      DO 7 I=1,NM1
      IF (I.GE.J) GO TO 5
      T = S(J)
      S(J) = S(I)
      S(I) = T
5      K = NV2
6      IF (K.GE.J) GO TO 7
      J= J-K
      K=K/2
      GO TO 6
7      J = J+K
      PI = 3.141592654
      DO 30 L=1,MPOWR
      LE =2**L
      LE1 = LE/2
      FLE1= FLOAT(LE1)
      U = (1.,0.,0.)
      PL=PI/FLE1
      PL =-PL
      B(1) = COS(PL)
      B(2) = SIN(PL)
      DO 20 J=1,LE1
      DO 11 I=J,NPTS,LE
      IP =I+LE1
      T=S(IP)*U
      S(IP)=S(I)*T

```

```

11      S(I) = S(I)+T
20      U=U*W1
30      CONTINUE
        NOP = NPTS-1
        IF(NWAY.EQ.1) GO TO 200
        DO 40 I=0,NOP
            IDX = I+1
            XX2 = DS*REAL(S(IDX))
40      S(I) = CMPLX(XX2,0.0)
        GO TO 210
200     DO 300 I=0,NOP
            IDX1 = I+1
            XYZ = (CABS(S(IDX1)))*DS
            XX3 = XYZ*XYZ
300     S(I) = CMPLX(XX3,0.0)
210     CONTINUE

D      TYPE 454,S(0),S(1)
D      TYPE 454,S(2),S(3)
D      TYPE 454,S(4),S(5)
D      TYPE 454,S(6),S(7)
D      TYPE 454,S(NPTS-4),S(NPTS-3)
D      TYPE 454,S(NPTS-2),S(NPTS-1)
        RETURN
        END

```

Subroutine CFFT

```

C      SUBROUTINE CFFT.  CALCULATES FFT OF ANY DATA ARRAY
C      NUMBER OF DATA POINTS IS POWER OF 2(M)
C

```

```

C      SUBROUTINE CFFT(MPOWR,NPTS,S,NWAY)

      DIMENSION B(2)
      COMPLEX S(0/32768),U,T,W1
      EQUIVALENCE (W1,B)
      DO 301 I=NPTS,1,-1
      NRAY = I-1
301      S(I) = S(NRAY)
D454      FORMAT(1X,4(E12.4,3X))
D      TYPE 454,S(1),S(2)
D      TYPE 454,S(3),S(4)
D      TYPE 454,S(5),S(6)
D      TYPE 454,S(NPTS/2+1)
D      TYPE 454,S(NPTS/2),S(NPTS/2+2)
D      TYPE 454,S(NPTS),S(NPTS-1),S(NPTS-2)
D      TYPE 454,S(NPTS/2-1),S(NPTS/2-2)
D      TYPE 454,S(NPTS/2+3),S(NPTS/2+4)
C

```

```

      NV2 =NPTS/2
      NM1 = NPTS-1
      J = 1
      DS = 1./FLOAT(NPTS)
      DO 7 I=1,NM1
      IF (I.GE.J) GO TO 5
      T = S(J)
      S(J) = S(I)
      S(I) = T
5      K = NV2
6      IF (K.GE.J) GO TO 7
      J= J-K
      K=K/2
      GO TO 6
7      J = J+K
      PI = 3.141592654
      DO 30 L=1,MPOWR
      LE =2**L
      LE1 = LE/2
      FLE1= FLOAT(LE1)
      U = (1.,0.,0)
      PL=PI/FLE1
      PL =-PL
      B(1) = COS(PL)
      B(2) = SIN(PL)
      DO 20 J=1,LE1
      DO 11 I=J,NPTS,LE
      IP =I+LE1
      T=S(IP)*U
      S(IP)=S(I)*T

```

```

11      S(I) =S(I)+T
20      U=U*W1
30      CONTINUE
        NOP = NPTS-1
        IF(NWAY.EQ.1) GO TO 200
        DO 40 I=0,NOP
            IDX = I+1
            XX2 = DS*REAL(S(IDX))
40      S(I) = CMPLX(XX2,0.0)
        GO TO 210
200     DO 300 I=0,NOP
            IDX1 = I+1
            XYZ = (CABS(S(IDX1)))*DS
            XX3 = XYZ*XYZ
300     S(I) = CMPLX(XX3,0.0)
210     CONTINUE

D      TYPE 454,S(0),S(1)
D      TYPE 454,S(2),S(3)
D      TYPE 454,S(4),S(5)
D      TYPE 454,S(6),S(7)
D      TYPE 454,S(NPTS-4),S(NPTS-3)
D      TYPE 454,S(NPTS-2),S(NPTS-1)
        RETURN
        END

```

Subroutine DGELG

.....

SUBROUTINE DGELG

PURPOSE

TO SOLVE A GENERAL SYSTEM OF SIMULTANEOUS LINEAR EQUATIONS.

USAGE

CALL DGELG(R,A,M,N,EPS,IER)

DESCRIPTION OF PARAMETERS

- R - DOUBLE PRECISION M BY N RIGHT HAND SIDE MATRIX
 (DESTROYED). ON RETURN R CONTAINS THE SOLUTIONS
 OF THE EQUATIONS.
- A - DOUBLE PRECISION M BY M COEFFICIENT MATRIX
 (DESTROYED).
- M - THE NUMBER OF EQUATIONS IN THE SYSTEM.
- N - THE NUMBER OF RIGHT HAND SIDE VECTORS.
- EPS - SINGLE PRECISION INPUT CONSTANT WHICH IS USED AS
 RELATIVE TOLERANCE FOR TEST ON LOSS OF
 SIGNIFICANCE.
- IER - RESULTING ERROR PARAMETER CODED AS FOLLOWS
 IER=0 - NO ERROR,
 IER=-1 - NO RESULT BECAUSE OF M LESS THAN 1 OR
 PIVOT ELEMENT AT ANY ELIMINATION STEP
 EQUAL TO 0,
 IER=K - WARNING DUE TO POSSIBLE LOSS OF SIGNIFI-
 CANCE INDICATED AT ELIMINATION STEP K+1,
 WHERE PIVOT ELEMENT WAS LESS THAN OR
 EQUAL TO THE INTERNAL TOLERANCE EPS TIMES
 ABSOLUTELY GREATEST ELEMENT OF MATRIX A.

REMARKS

INPUT MATRICES R AND A ARE ASSUMED TO BE STORED COLUMNWISE
IN M*N RESP. M*M SUCCESSIVE STORAGE LOCATIONS. ON RETURN
SOLUTION MATRIX R IS STORED COLUMNWISE TOO.
THE PROCEDURE GIVES RESULTS IF THE NUMBER OF EQUATIONS M IS
GREATER THAN 0 AND PIVOT ELEMENTS AT ALL ELIMINATION STEPS
ARE DIFFERENT FROM 0. HOWEVER WARNING IER=K - IF GIVEN -
INDICATES POSSIBLE LOSS OF SIGNIFICANCE. IN CASE OF A WELL
SCALED MATRIX A AND APPROPRIATE TOLERANCE EPS, IER=K MAY BE
INTERPRETED THAT MATRIX A HAS THE RANK K. NO WARNING IS
GIVEN IN CASE M=1.

SUBROUTINES AND FUNCTION SUBPROGRAMS REQUIRED

NONE

METHOD

SOLUTION IS DONE BY MEANS OF GAUSS-ELIMINATION WITH
COMPLETE PIVOTING.

```

.....
SUBROUTINE DGELG(R,A,M,N,EPS,IER)

```

```

DIMENSION A(1),R(1)
DOUBLE PRECISION R,A,PIV,TB,TOL,PIVI
IF(M)23,23,1

```

```

SEARCH FOR GREATEST ELEMENT IN MATRIX A

```

```

1 IER=0
  PIV=0,D0
  MM=M*M
  NM=N*M
  DO 3 L=1,MM
    TB=DABS(A(L))
    IF(TB=PIV)3,3,2
2 PIV=TB
  I=L
3 CONTINUE
  TOL=EPS*PIV
  A(I) IS PIVOT ELEMENT. PIV CONTAINS THE ABSOLUTE VALUE OF A(I).

```

```

START ELIMINATION LOOP
LST=1
DO 17 K=1,M

```

```

TEST ON SINGULARITY

```

```

IF(PIV)23,23,4
4 IF(IER)7,5,7
5 IF(PIV-TOL)6,6,7
6 IER=K+1
7 PIVI=1,D0/A(I)
  J=(I-1)/M
  I=I-J*M-K
  J=J+1=K
  I+K IS ROW-INDEX, J+K COLUMN-INDEX OF PIVOT ELEMENT

```

```

PIVOT ROW REDUCTION AND ROW INTERCHANGE IN RIGHT HAND SIDE R

```

```

DO 8 L=K,NM,M
  LL=L+I
  TB=PIVI*R(LL)
  R(LL)=R(L)
8 R(L)=TB

```

```

IS ELIMINATION TERMINATED
IF(K=M)9,18,18

```



```

C      COLUMN INTERCHANGE IN MATRIX A
9      LEND=LST+M-K
      IF(J)12,12,10
10     II=J-M
      DO 11 L=LST,LEND
      TB=A(L)
      LL=L+II
      A(L)=A(LL)
11     A(LL)=TB

C
C      ROW INTERCHANGE AND PIVOT ROW REDUCTION IN MATRIX A
12     DO 13 L=LST,MM,M
      LL=L+I
      TB=PIVI*A(LL)
      A(LL)=A(L)
13     A(L)=TB

C
C      SAVE COLUMN INTERCHANGE INFORMATION
      A(LST)=J

C
C      ELEMENT REDUCTION AND NEXT PIVOT SEARCH
      PIV=0,D0
      LST=LST+1
      J=0
      DO 16 II=LST,LEND
      PIV= A(II)
      IST=II+M
      J=J+1
      DO 15 L=IST,MM,M
      LL=L-J
      A(L)=A(L)+PIVI*A(LL)
      TB=DABS(A(L))
      IF(TB=PIV)15,15,14
14     PIV=TB
      I=L
15     CONTINUE
      DO 16 L=K,NM,M
      LL=L+J
16     R(LL)=R(LL)+PIVI*R(L)
17     LST=LST+M
      END OF ELIMINATION LOOP

C
C
C      BACK SUBSTITUTION AND BACK INTERCHANGE
18     IF(M=1)23,22,19
19     IST=MM+M
      LST=M+1
      DO 21 I=2,M
      II=LST-I
      IST=IST-LST
      L=IST-M

```

```

      L=A(L)+.5D0
      DO 21 J=II,NM,M
      TB=R(J)
      LL=J
      DO 20 K=IST,MM,M
      LL=LL+1
20    TB=TB-A(K)*R(LL)
      K=J+L
      R(J)=R(K)
21    R(K)=TB
22    RETURN

C
C
C      ERROR RETURN
23    IER=-1
      RETURN
      END

```

Subroutine FAC1

```

C      SUBROUTINE FAC1: COMPUTES K!/((K-M)!M!)
C
C      SUBROUTINE FAC1(K,M,AKM)
C
      IF (K.EQ.0.OR.K.EQ.1) GO TO 35
      FK =FLCAT(K)
      FKC = FK
      DO 30 I=1,K-1
      AI =FLOAT(I)
30      FK = FK*(FKC-AI)
35      IF (K.EQ.1.OR.K.EQ.2) FK =1.
      IF (M.EQ.1.OR.M.EQ.2) GO TO 45
      FM =FLCAT(M)
      FMC = FM
      DO 40 I=1,M-1
      AI = FLOAT(I)
40      FM = FM*(FMC-AI)
45      IF (M.EQ.1.OR.M.EQ.2) FM = 1.
      FMK = FLCAT(K)-FLCAT(M)
      IF (FMK.EQ.1..OR.FMK.EQ.0.0) GO TO 55
      KM = K-M-1
      FKMC = FMK
      DO 50 I=1,KM
      AI = FLCAT(I)
50      FMK = FMK*(FKMC-AI)
55      IF (FMK.EQ.1.0.CR.FMK.EQ.0.0) FMK=1.0
      AKM = FK/(FMK*FM)
      RETURN
      END

```

Program FINAL

PROGRAM FINAL. ATMOSPHERIC TURBULENCE TASK
 ITEM 2(REVISION 2&3), PARTS 6,7,8, AND 9. PHASE II OF ATMOS-
 PHERIC TURBULENCE
 MEAN VALUE SUBTRACTED FROM DATA BEFORE COMPUTING SPECTRUM
 COMPONENT WITH LOW FREQUENCY CONTAMINATION

REQUIRES SUBROUTINES GAM,PAR1 & INTERPOLATION
 ROUTINE: ANRP1; AK1 REQUIRES PAR,AKDAT,SIMQ,DGELG;
 USES SUBROUTINES TRAP(3&6) AND SIMP2 FOR INTEGRATIONS
 ALSO USES SUBROUTINE SET AND FUNCTION DECT INCLUDED IN PROGRAM

LMAX MAX = 10; NCOUNT MAX = 10; NXIH MAX = 290; M MAX = 4

DATA FILES READ BY PROGRAM ARE:

AUTO1: AUTOCORRELATION -FROM ATURB2

DATA FILE CREATED BY THIS PROGRAM ARE:

ITM2: VALUES OF CAP R (XI), PART 9,K, ITEM 2

ITM2L: VALUES OF SIGMA SQRD*L*PHI OF K(KL), PART 9,L

DOUBLE PRECISION AMTRX,XC,DIJ,A,SUM,AM,S1,S2,RI,CJ,ALG2
 DOUBLE PRECISION S3,S4

COMMON/BB/X(3),AKI(3),XPL
 COMMON/TR5/DELX1,NXIH1,AI5I(0/290)
 COMMON/TR2/DELX,NXIH,AI2I(0/290)

DIMENSION DIJ(7,7),RI(7),CJ(7)
 DIMENSION AL(0/11),SIG2(0/11),PH1(0/100),PH2(0/100)
 DIMENSION PHI(0/11,0/290),PHI1(0/11,0/290),AI(0/6)
 DIMENSION AI3(0/4,0/11),AI4(0/4,0/11),AI5(0/11),AI6(0/11)
 DIMENSION AI7(0/9),AI8(0/4),RL(0/290)
 DIMENSION SIG2J(0/11),XC(6),A(6,6),AMTRX(36)
 DIMENSION AI1(0/11),AI2(0/11),AI4T(0/4),AI3T(0/4)
 DIMENSION NDO(0/11),NUPR(0/11),ELEFT(0/11)
 DIMENSION EWAY(0/11),AI5X(0/11),AI6X(0/11)

TYPE 450

FORMAT(' HAS DATA BEEN COMPUTED ?',/8)

ACCEPT 807,AND11

IF (AND11.EQ.'Y') GO TO 451

TYPE 1001

FORMAT(/1X,'INPUT DELX,LMAX',/8)

ACCEPT 1750,DELX,LMAX

FORMAT(2G)

TYPE 1751

FORMAT(' IS RECORD TRANSVERSE OR LONGITUDIONAL (T OR L) ',/8)

ACCEPT 807,WHICH

007

FORMAT(A5)
DELX1= DELX

C
C
C
C

SIGMA SQRD & L COMPUTED FROM PROGRAM PART5

AL(0) = 20.
AL(1) = 30.
AL(2) = 40.
AL(3) = 50.
AL(4) = 60.
AL(5) = 70.
AL(6) = 80.
AL(7) = 90.
AL(8) = 100.
AL(9) = 110.
AL(10) = 120.
AL(11) = 130.
SIG2(0) = .339271
SIG2(1) = .3400205
SIG2(2) = .3666069
SIG2(3) = .4009919
SIG2(4) = .4378971
SIG2(5) = .4753911
SIG2(6) = .5126825
SIG2(7) = .5494336
SIG2(8) = .5855068
SIG2(9) = .6208566
SIG2(10) = .6554832
SIG2(11) = .73219

C
C
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C

PART 7. REQUIRES SUBROUTINES GAM, TRAP(3&6),AK
COMPUTES VON-KARMEN AUTOCORRELATION

TYPE 1511

1511 FORMAT(' INPUT HIGHEST INDEX OF XI, EH, & M, ')
ACCEPT 1752,NXIH,EH,MM

1752 FORMAT(3G)
NXIH1=NXIH
PI = 3.14159265
C116 = 11./6.
C56 = 5./6.
C43 = 4./3.
CALL GAM(C116,G1)
CALL GAM(C43,G2)
GAM13 = 2.67893853
BETA = 2.*G1*(PI**.5)/(5.*G2)
C23 = 2./3.
C13 = 1./3.
CONST = -BETA*(2.**C23)/GAM13
CONS1 = BETA/(GAM13*2.**C13)

```

CONS2 = ((2.**C23)/GAM13)
PH2(0) = 0.0
PH1(0) = 1.0
DO 1070 I=1,100
  E1 = .1*I
  CALL AK(1,E1,AK1)
  CALL AK(2,E1,AK2)
  PH1(I) = CONS2*(E1**C13)*AK1-CONS2*(E1**C13)*AK2*E1/2.
  IF (WHICH.EQ.'L') PH1(I) = CONS2*
1      (E1**C13)*AK1
  PH2(I)=CONS1*((E1**C43)*AK1 -
1      8.*(E1**C13)*AK2/3.)
  IF (WHICH.EQ.'L') PH2(I) = CONST*(E1**C13)*AK2
1070 CONTINUE
D TYPE 6100,PH1(97),PH1(98),PH1(99),PH1(100)
D6100 FORMAT(' PH1: ',4G)
D TYPE 6101,PH2(97),PH2(98),PH2(99),PH2(100)
D6101 FORMAT(' PH2 : ',4G)
C
C DO INTEGRALS 1 TO 4
C PH1 AND PH2 = 0 IF MORE THAN 100 PTS USED (NUPR IS ARRAY INDEX)
C IF LESS THAN 100 PTS USED IN INTEGRAL, SIMPSON'S RULE
C IS USED UP TO LAST "EVEN" PT IN INTEGRAL. THE END OF THE
C INTEGRAL IS DONE AS PER HILDEBRAND (P.111) FOR THE REMAINING
C PART OF THE INTEGRAL (USING TERMS UP TO THE FOURTH POWER)
C
EBAR = .1/BETA
DO 6068 L = 0,LMAX+1
  EHMAX = EH/AL(L)
  NUPR(L) = IFIX(EHMAX/EBAR)
  NDO(L) = 0
  IF (NUPR(L).GT.100) NUPR(L) = 100
  IF (NUPR(L).EQ.100) GO TO 444
  NDO(L) = 1
  AODD = FLOAT(NUPR(L))/2.
  NODD = NUPR(L)/2
  REM = AODD = FLOAT(NODD)
  IF (REM.GE'.5) NUPR(L) = NUPR(L) +1
  ELEFT(L) = EHMAX = FLOAT(NUPR(L))*EBAR
D TYPE 447,AODD,NODD,REM,ELEFT(L)
D447 FORMAT(' AODD,NODD,REM,ELEFT(L) :',G,2X,I3,2X,2G)
444 CONTINUE
D TYPE 6003,NUPR(L),L
D6003 FORMAT(' NUPR(L) = ',I4,' L INDEX = ',I3)
DO 3002 I=0,NUPR(L)
  E = EBAR*I
  AI2I(I) = PH1(I)*PH2(I)*E
  AI5I(I) = PH1(I)*PH1(I)
  DO 3003 J=0,MM
    J1 = J+1
    PHI(J,I) = (E**J1)*PH2(I)
    PHI1(J,I) = (E**J)*PH1(I)

```



```

3003 CONTINUE
3002 CONTINUE
      CALL SIMP(EBAR,NUPR(L),AI2I,A1)
      CALL SIMP(EBAR,NUPR(L),AI5I,A2)
      AI1(L) = A1
      AI2(L) = A2
      IF (NDO(L).EQ.0) GO TO 445
D      TYPE 449,L,AI1(L),AI2(L)
D449  FORMAT(' FOR L =',I3,' AI1 =',G,' AI2 =',G)
      AI1(L) = AI1(L) + DECT(ELEFT(L),AI2I,NUPR(L))
D      TYPE 448,AI1(L)
D448  FORMAT(' AI1 =',G)
      AI2(L) = AI2(L) + DECT(ELEFT(L),AI5I,NUPR(L))
445   CONTINUE
D      TYPE 6609,L,AI1(L),AI2(L)
D6609 FORMAT(1X,'L,AI1,AI2 :',1X,I4,2G)
      DO 3004 JJ=0,MM
      DO 3104 I=0,NUPR(L)
      AI2I(I) = PHI(JJ,I)
3104  AI5I(I) = PHI(JJ,I)
      CALL SIMP(EBAR,NUPR(L),AI2I,A1)
      AI3(JJ,L) = A1
      CALL SIMP(EBAR,NUPR(L),AI5I,A2)
      AI4(JJ,L) = A2
      IF (NDO(L).EQ.0) GO TO 446
D      TYPE 453,JJ,L,AI3(JJ,L),AI4(JJ,L)
D453  FORMAT('JJ,L,AI3,AI4:',I3,1X,I2,1X,2G)
      AI3(JJ,L) = AI3(JJ,L) + DECT(ELEFT(L),AI2I,NUPR(L))
      AI4(JJ,L) = AI4(JJ,L) + DECT(ELEFT(L),AI5I,NUPR(L))
446   CONTINUE
D      TYPE 6608,JJ,L,AI3(JJ,L),AI4(JJ,L)
D6608 FORMAT(1X,'J,L,AI3,AI4 :',2X,2I4,2G)
3004  CONTINUE
6068  CONTINUE
C
C      READ IN AUTOCORRELATION FN.
C
      CALL IFILE (20,'AUTO')
      READ(20,900)
      READ(20,1950)FLE
1950  FORMAT(/1X,'AUTOCORRELATION OF STATIONARY SAMPLE',/1X,
1     'DATA TAKEN FROM FILE ',A5)
      READ(20,906)NPTS,TWOL,MPTS,FTM1
      READ(20,907)MZERO
      READ(20,910)WBAR,VAR
      READ(20,951)WSUM
951   FORMAT(/1X,'<W OF L(X)**2> = ',F12.4)
      READ(20,381)XAOFF,MAOFF
381   FORMAT(/1X,' TRUCATION POINT WAS',G,' METERS',/1X,
1     'WHICH CONTAINS ',I7,' POINTS')
      READ(20,1952)

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1952 1 FORMAT(/1X,'PRINTOUT OF THE VALUES OF THE AUTOCORRELATION',
/1X,8X,'X',12X,' RL ',9X,'RL/R0')
DO 1823 I=0,NXIH
1823 READ(20,1915)DEL,RL(I),YY
1915 FORMAT(1X,3(2X,G))
END FILE 20
TYPE 4324,RL(NXIH)
4324 FORMAT(' RL(NXIH) =',G)
C
C DETERMINE WHETHER SIMPSON'S RULE OR TRAP. RULE IS TO BE
C USED IN INTEGRAL I5 & I6
C
DO 7023 L=0,LMAX+1
EB = DELX/AL(L)
EWAY(L) = 'TRP'
7023 IF (EB.GT.EBAR) EWAY(L) = 'SMP'
LCHNG = -2
TP = 1
DO 7024 L=0,LMAX+1
IF (EWAY(L).EQ.'SMP') TP = 0
7024 IF (EWAY(L).EQ.'SMP') LCHNG = L
IF (TP.EQ.1) TYPE 7006
7006 FORMAT(' SIMPSONS RULE NOT USED FOR I5 & I6')
IF (TP.EQ.0) TYPE 7007,LCHNG
7007 FORMAT(' SIMPSONS RULE USED IN I5& I6 UP TO L INDEX =',I3)
C
C DO INTEGRAL I5 AND I6 WHEN E/L > EBAR
C
IF (TP.EQ.1) GO TO 7008
DO 7050 L=0,LMAX+1
IF (EWAY(L).EQ.'TRP') GO TO 7050
ALBAR = AL(L)*EBAR
DO 7002 I=0,NUPR(L)
ALDT = I*ALBAR
NPLCE = IFIX(ALDT/DELX+.5)
IXMID = NPLCE
IXLOW = IXMID - 1
IXHIG = IXMID + 1
IF (IXMID.EQ.0) IXLOW = 0
IF (IXMID.EQ.0) IXHIG = 2
IF (IXMID.EQ.0) IXMID = 1
IF (IXMID.GE.NXIH) IXLOW = NXIH-2
IF (IXMID.GE.NXIH) IXHIG = NXIH
IF (IXMID.GE.NXIH) IXMID = NXIH-1
AKI(1) = RL(IXLOW)
AKI(2) = RL(IXMID)
AKI(3) = RL(IXHIG)
X(1) = IXLOW*DELX
X(2) = IXMID*DELX
X(3) = IXHIG*DELX
XPL = ALDT
CALL PARAB(AK2)
RLI = AK2
E=EBAR*I
AI2I(I) = E*PH2(I)*RLI

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7002  AI5I(I) = PHI(I)*RLI
      CALL SIMP(EBAR,NUPR(L),AI2I,A3)
      CALL SIMP(EBAR,NUPR(L),AI5I,A4)
      AI5(L) = A3/RL(0)
      AI6(L) = A4/RL(0)
D
D7449  TYPE 7449,L,AI5(L),AI6(L)
      FORMAT(' L,AI5,AI6:',3G)
      IF (NDO(L).EQ.0) GO TO 7050
      AI5(L) = AI5(L) + (DECT(ELEFT(L),AI2I,NUPR(L)))/RL(0)
      AI6(L) = AI6(L) + (DECT(ELEFT(L),AI5I,NUPR(L)))/RL(0)
D
D7450  TYPE 7450,AI5(L),AI6(L)
      FORMAT(' AI5,AI6:',2G)
7050  CONTINUE

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C
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C
C

INTERPOLATE PHI AND PH2 WHEN E/L < EBAR

```

7008  DO 1050 L=0,LMAX+1
      IF (L.LE.LCHNG) GO TO 1050
      PHI(L,0) = 1.0
      PHI1(L,0) = 0.0
      EB = DELX/AL(L)
D
D1071  TYPE 1071,EB,EBAR
      FORMAT(' EB =',G,' EBAR =',G)
      XMAX2 = DELX*NXIH
      XMAX1 = 100.*EBAR
      DO 2071 JX = 1,NXIH
        IXMID = IFIX(EB*JX/EBAR+.5)
        IXLOW = IXMID-1
        IXHIG = IXMID+1
        IF (IXMID.GE.100) IXHIG = 100
        IF (IXMID.GE.100) IXLOW = 98
        IF (IXMID.GE.100) IXMID = 99
        IF (IXMID.EQ.0) IXHIG = 2
        IF (IXMID.EQ.0) IXLOW = 0
        IF (IXMID.EQ.0) IXMID = 1
        AKI(1) = PHI(IXLOW)
        AKI(2) = PHI(IXMID)
        AKI(3) = PHI(IXHIG)
        X(1) = IXLOW*EBAR
        X(2) = IXMID*EBAR
        X(3) = IXHIG*EBAR
        XPL = EB*JX
        IF (XPL.GT.XMAX1) GO TO 2072
        CALL PARAB(AKZ)
        PHI(L,JX) = AKZ
        AKI(1) = PH2(IXLOW)
        AKI(2) = PH2(IXMID)
        AKI(3) = PH2(IXHIG)
        X(1) = IXLOW*EBAR
        X(2) = IXMID*EBAR
        X(3) = IXHIG*EBAR
        XPL = EB*JX
        CALL PARAB(AKZ)
        PHI1(L,JX) = AKZ

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2072 IF (XPL.GT.XMAX1) PHI(L,JX) = 0.0
      IF (XPL.GT.XMAX1) PHI1(L,JX) = 0.0
2071 CONTINUE
1050 CONTINUE
C
C      INTEGRALS 5 AND 6 WHEN E/L < EBAR
C
      DO 3010 L=0,LMAX+1
      IF (L.LE.LCHNG) GO TO 3010
      DO 3011 I=0,NXIH
      AI5I(I) = DELX*I*PHI1(L,I)*RL(I)
3011 AI2I(I) = PHI(L,I)*RL(I)
      CALL TRAP2(AIXY)
      AI6(L) = AIXY/(RL(0)*AL(L))
      CALL TRAP5(AIXY)
      AI5(L) = AIXY/(RL(0)*AL(L)*AL(L))
3010 CONTINUE
D      DO 3019 L=0,LMAX+1
D      IF (EWAY(L).EQ.'SMP') GO TO 3019
D      TYPE 6007,L,AI5(L)
D      TYPE 6008,L,AI6(L)
D3019 CONTINUE
D6007 FORMAT(' FOR L= ',I3,' AI5 = ',G)
D6008 FORMAT(' FOR L= ',I3,' AI6 = ',G)
C
C      INTEGRAL 18(J)
C
      DO 1080 J2 = 0,MM
      DO 1081 I=0,NXIH
1081 AI2I(I) = ((DELX*I)**J2)*RL(I)
      CALL TRAP2(AIXY)
1080 AI8(J2) = AIXY
C
C      INTEGRAL 17(J)
C
      MMTWO = 2*MM
      DO 1084 J2 = 0,MMTWO
      IEX = J2+1
1084 AI7(J2) = (EH**IEX)/(FLOAT(J2) + 1.)
D      TYPE 5219,AI7(MMTWO)
D5219 FORMAT(' FINAL VALUE OF I7 IS: ',/1X,G)
C
C      START ITERATION
C
      NCOUNT = -1
19      NCOUNT = NCOUNT + 1
      TYPE 4
4      FORMAT(' PICK A SIGMA SQRD AND ITS ASSOCIATED L INDEX',/8)
      ACCEPT 1753,SIG2J(NCOUNT),LSIG
1753 FORMAT(2G)
C

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C      DETERMINE CLOSEST MIDPT FOR INTERPOLATION ROUTINES
C
DIY1 = ABS(SIG2J(NCOUNT)-SIG2(LSIG))
DIY2 = ABS(SIG2J(NCOUNT)-SIG2(LSIG+1))
LCTR = LSIG
IF (DIY2.LT.DIY1) LCTR = LSIG + 1
D      TYPE 7028,LSIG,LCTR
D7028  FORMAT(' LSIG =',I3,' LCTR =',I3)
C
C      CALL INTERPOLATION ROUTINES
C
CALL SET(AL,SIG2,SIG2J(NCOUNT),LCTR,BK1)
ALINT = BK1
TYPE 6056,SIG2J(NCOUNT),ALINT
6056   FORMAT(' FOR SIGMA SORD =',G,' L INT =',G)
C
IF (LCTR.LT.(LCHNG+2)) GO TO 7012
DO 7010 L=0,LMAX+1
AI5X(L) = AI5(L)*AL(L)*AL(L)
7010   AI6X(L) = AI6(L)*AL(L)
CALL SET(AI5X,AL,ALINT,LCTR,BK1)
AI5T = BK1/(ALINT*ALINT)
CALL SET(AI6X,AL,ALINT,LCTR,BK1)
AI6I = BK1/ALINT
GO TO 7011
7012   CALL SET(AI5,AL,ALINT,LCTR,BK1)
AI5T = BK1
CALL SET(AI6,AL,ALINT,LCTR,BK1)
AI6I = BK1
7011   CALL SET(AI1,AL,ALINT,LCTR,BK1)
AI1T = BK1
CALL ANTRP(LMAX,ALINT,AL,LSIG,AI2,AI2T)
C
D      TYPE 9084,AI5T,AI6I,AI1T,AI2T
D9084  FORMAT(' FOR INT AI5I =',G,' AI6I =',G,' AI1T =',G,' AI2T =',G)
DO 6069 IJK =0,MM
AKI(1) = AI3(IJK,LCTR-1)
AKI(2) = AI3(IJK,LCTR)
AKI(3) = AI3(IJK,LCTR+1)
X(1) = AL(LCTR-1)
X(2) = AL(LCTR)
X(3) = AL(LCTR+1)
XPL = ALINT
CALL PARAB(AKZ)
AI3T(IJK) = AKZ
AKI(1) = AI4(IJK,LCTR-1)
AKI(2) = AI4(IJK,LCTR)
AKI(3) = AI4(IJK,LCTR+1)
X(1) = AL(LCTR-1)
X(2) = AL(LCTR)
X(3) = AL(LCTR+1)
XPL = ALINT
CALL PARAB(AKZ)
AI4T(IJK) = AKZ

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D      TYPE 6010,IJK,AI3T(IJK),AI4T(IJK)
D6010  FORMAT(' J =',I3,'AI3T =',G,'AI4T =',G)
6069   CONTINUE
C
C      COMPUTE DEL(L)/DEL(SIGMA SQRD)
C
      AKI(1) = SIG2(LCTR=1)
      AKI(2) = SIG2(LCTR)
      AKI(3) = SIG2(LCTR+1)
      X(1) = AL(LCTR=1)
      X(2) = AL(LCTR)
      X(3) = AL(LCTR+1)
      XPL = SIG2J(NCOUNT)
      CALL PAR1(DLDS)
      TYPE 7004,DLDS
7004   FORMAT(' DLDS = ',G)
C
C      COMPUTE COEFFICIENTS XC = R.H.S., ARRAY A = L.H.S.
C
      XC(1) = RL(0)*(-SIG2J(NCOUNT)*DLDS*AI5T+ALINT*AI6T)
      A(1,1) = -SIG2J(NCOUNT)*DLDS*AI1T+ALINT*AI2T
      DO 3013 K=2,MM+2
      K2 = K-2
      L=K
      L2 = K2
      XC(K) = AI0(K2)/ALINT**K2
      A(K,1) = ALINT*AI4T(K2)
      A(1,L) = -SIG2J(NCOUNT)*DLDS*(ALINT**L2)*AI3T(L2)
      + (ALINT**(L-1))*AI4T(L2)
1      DO 3014 LA = 2,MM+2
      LAK = K+LA-4
3014   A(K,LA) = (ALINT**(LA-LAK-2))*AI7(LAK)
3013   CONTINUE
D      TYPE 7003,XC(1),A(1,1),A(1,2)
D      TYPE 7005,XC(2),A(2,1),A(2,2)
D7003  FORMAT(' XC(1),A(1,1),A(1,2):',3G)
D7005  FORMAT(' XC(2),A(2,1),A(2,2):',3G)
C
C      DO SIM. LINEAR EQN USING DOUBLE PRECISION PROGRAM DGELC
C
      MM2 = MM+2
      ALG2 = DLOG10(2.D0)
      MM3 = MM+3
      DO 10 I=1,MM2
      DO 10 J=1,MM2
10     DIJ(I,J) = DLOG10(DABS(A(I,J)))/ALG2
      DO 20 I=1,MM2
20     DIJ(I,MM3) = DLOG10(DABS(XC(I)))/ALG2
C      IF MM =4 USE DOUBLE PRECISION CONSTANTS BELOW

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C      IF NOT USE FOLLOWING LOGIC
      DMM23 = 42.D0
      DMM2 = 6.D0
      DMM3 = 7.D0
      IF (MM.NE.0) GO TO 6683
      DMM23 = 6.D0
      DMM2 = 2.D0
      DMM3 = 3.D0
      GO TO 6681
6683   IF (MM.NE.1) GO TO 6680
      DMM23 = 12.D0
      DMM2 = 3.D0
      DMM3 = 4.D0
      GO TO 6681
6680   IF (MM.NE.2) GO TO 6682
      DMM23 = 20.D0
      DMM2 = 4.D0
      DMM3 = 5.D0
      GO TO 6681
6682   IF (MM.NE.3) GO TO 6681
      DMM23 = 30.D0
      DMM2 = 5.D0
      DMM3 = 6.D0
6681   SUM = 0.0
      DO 30 J=1,MM3
      DO 30 I=1,MM2
30      SUM = SUM +DIJ(I,J)
      AM = -SUM/DMM23
      DO 40 I=1,MM2
      S1 = 0.0
      DO 50 J=1,MM3
50      S1 = S1 + (DIJ(I,J)+AM)
40      RI(I) = -S1/DMM3
      DO 60 J=1,MM3
      S2 = 0.0
      DO 70 I=1,MM2
70      S2 = S2 + (DIJ(I,J)+AM)
60      CJ(J) = -S2/DMM2
      DO 80 J=1,MM2
      IX = (J-1)*MM2
      DO 80 I=1,MM2
80      AMTRX(I+IX) = A(I,J)*2.D0** (RI(I)+CJ(J)+AM)
      DO 90 I=1,MM2
90      XC(I) = XC(I)*2.D0** (RI(I)+CJ(MM3)+AM)
      S3 = 0.0
      DO 110 I=1,MM2
110     S3 = S3 + RI(I)
      S4 = 0.0
      DO 120 J=1,MM3
120     S4 = S4 + CJ(J)
      CALL DGELG(XC,AMTRX,MM2,1,1,E=10,IER)
      TYPE 3016,IER

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3016  FORMAT(' IER = ',G)
      DO 100 I=1,MM2
100    XC(I) = XC(I)*2.**((CJ(I)-CJ(MM3)))
      SIG2J(NCOUNT+1) = XC(1)
      TYPE 1620,NCOUNT,XC(1)
1620  FORMAT(' FOR NCOUNT = ',I2,' SIGMA SQRD = ',G,/1X)
C
C
28    TYPE 1627
1627  FORMAT(' STOP LOOP ? (Y OR N) ',I)
      ACCEPT 807,STOP
      IF (STOP.EQ.'N') GO TO 19
C
C
      COMPUTE FINAL A'S
      DO 3017 L=2,MM2
3017  AI(L=2) = XC(L)
C
      TYPE OUT FINAL RESULTS
C
31    TYPE 1628,NCOUNT,ALINT,XC(1)
1628  FORMAT(' ON LAST PASS NCOUNT = ',I3,' L = ',G,/1X,
1      ' AND SIGMA SQRD = ',G)
      DO 32 I2 = 0,MM
32    TYPE 1629,I2,AI(I2)
1629  FORMAT(' A(',I2,') = ',G)
C
      COMPUTE R(XI), PART 9.K
C
      CALL OFILE(20,'ITM2')
      WRITE(20,915),MM,EH,FLE
      WRITE(20,1632)ALINT,XC(1)
1632  FORMAT(/1X,'OUTPUT FOR PART 9.K',/1X,'WITH L = ',G,
1      ' AND SIGMA SQRD = ',G)
      DO 7016 I=0,MM
7016  WRITE(20,7015),I,AI(I)
7015  FORMAT(' COEFFICIENT A(',I1,') = ',G)
      WRITE(20,1631)
1631  FORMAT(/3X/,' XI',12X,'R(XI)',12X,'XI CONTD',8X,'R(XI) CONTD')
      ARG = ALINT*.1/BETA
      NX2 = IFIX(EH/ARG)
      NX12 = (NX2/2) - 1
      DO 41 I=0,NX12
      E1 = ARG*I
      E2 = ARG*(I+NX12+1)
      SUM1 = 0.
      SUM2 = 0.
      DO 42 I0 = 0,MM
      SUM2 = SUM2+AI(I0)*E2**I0
42    SUM1 = SUM1+AI(I0)*E1**I0
      IF (I.GT.100) R = SUM1
      IF (I.GT.100) GO TO 43
      R = XC(1)*PH1(I)+SUM1

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43      IF ((I+NX12+1).GT.100) R2 = SUM2
      IF ((I+NX12+1).GT.100) GO TO 41
      R2 = XC(1)*PHI(I+NX12)+SUM2
41      WRITE(20,1630)E1,R,E2,R2
1630    FORMAT(4G)
      END FILE 20

C
C
C      DO PART 9,L, COMPUTE SIGMA SQRD*L*PHIK

      CALL OFILE(21,'ITM2L')
      WRITE(21,915),MM,EH,FLE
      WRITE(21,1633)ALINT,XC(1)
1633    FORMAT(/' OUTPUT FOR PART 9,L',/1X,' WITH L =',G,' AND SIGMA',
1      ' SQRD =',G,'/1X,' K',8X,' L*SIGMA SQRD*PHIK',5X,' K CONTD',4X,
2      'L*SIGMA SQRD*PHIK CONTD')
      SIG2L = XC(1)*ALINT
      TYPE 7013
7013    FORMAT(' INPUT DELK ',2X.6)
      ACCEPT 7014,DELK2
7014    FORMAT(G)
      DO 44 I = 0,511
      EXC1 = DELK2*I
      EXC2 = DELK2*(I+512)
      AKLR = ALINT*EXC1
      AKLR1 = ALINT*EXC2
      ALKI2 = AKLR*AKLR
      ALKI3 = AKLR1*AKLR1
      PHIK = (2./(1.+70.78*ALKI2)**C56)*SIG2L
      PHIK1 = (2./(1.+70.78*ALKI3)**C56)*SIG2L
      IF (WHICH.EQ.'T') PHIK = ((1.+188.75*ALKI2)/(1.+70.78*ALKI2)
1      **C116)*SIG2L
      IF (WHICH.EQ.'T') PHIK1 = ((1.+188.75*ALKI3)/(1.+70.78*ALKI3)
1      **C116)*SIG2L
44      WRITE(21,1630)EXC1,PHIK,EXC2,PHIK1
      END FILE 21

C
915    FORMAT(/1X,'DATA FILE CREATED BY PROGRAM FINAL',/1X,
1      'WITH M =',I3,' AND LENGTH =',G,' AUTOCOR. OF FILE ',A5)

900    FORMAT(/1X,'DATA FILE CREATED BY PROGRAM ATURB3')
903    FORMAT(1X,F10.6,3X,E12.4,3X,F10.6,3X,E12.4)
906    FORMAT(/1X,I6,' DATA POINTS WERE USED IN 2L = ',F15.4,
1      ' METER',/1X,I5,' DATA POINTS WERE USED IN M = ',F16.4,
1      ' METER')
907    FORMAT(/1X,I6,' ZEROS WERE ADDED TO DATA')
910    FORMAT(/1X,'MEAN VALUE OF W(X) = ',E15.5,' M/SEC',/1X,
1      'MEAN SQ. VALUE = ',E15.5,' (M/SEC)**2')

C
9999   END
C

```

```

C      FUNCTION DECT(ULEFT,AIXZ,NAU1)
      DIMENSION AIXZ(0/290)
      DECT = ULEFT*(2.6402778*AIXZ(NAU1)-3.852778*AIXZ(NAU1-1)
1          +3.6333333*AIXZ(NAU1-2)-1.7694444*AIXZ(NAU1-3)
2          +.34861111*AIXZ(NAU1-4))
      RETURN
C
C      END
C
C      SUBROUTINE SET(A1,B1,C1,LSIG,BK1)
C
C      DIMENSION A1(0/11),B1(0/11)
      COMMON/BB/X(3),AKI(3),XPL
C
      AKI(1) = A1(LSIG-1)
      AKI(2) = A1(LSIG)
      AKI(3) = A1(LSIG+1)
      X(1) = B1(LSIG-1)
      X(2) = B1(LSIG)
      X(3) = B1(LSIG+1)
      XPL = C1
      CALL PARAB(AKZ)
      BK1 = AKZ
C
C      RETURN
C
C      END

```

Subroutine GAM

```

C <RFISHER>GAM.F4;1      28-Sep-77 13:47:22      EDIT BY RFISHER
      SUBROUTINE GAM(GAMMA,GG)
C      PROGRAM TO COMPUTE GAMMA FUNCTIONS
C      PROGRAM DERIVED FROM "HANDBOOK OF MATH. FUNCTIONS, NAT.
C      BUREAU OF STANDARDS, APPLIED MATH SERIES 55", 1964
C      PROGRAM USES EQN 6.1.15 (P 256) AND APPROX. 6.1.36 (P.257)
C
C      ANS RETURNED AS GG, INPUT IS GAMMA
C
      IF (GAMMA.GE.2.) GO TO 1
      X = GAMMA -1.
      RES =1.
      GO TO 10
1      N = IFIX(GAMMA)-1
      RES = 1.
      DO 20 I=1,N
20     RES = RES*(GAMMA-FLOAT(I))
      X = GAMMA - FLOAT(N+1)
10     GG = 1. - .577191652*X +.988205891*X**2 - .897056937*X**3
1      + .918206857*X**4 - .756704278*X**5 + .482199394*X**6
2      - .193527818*X**7 + .035868343*X**8
C
      GG = GG*RES
C
      RETURN
C
      END

```

Program GDIST6

```

C      PROGRAM TO COMPUTE DIST. SEVERAL WAYS. NOTES FROM BILL
C      MARK OF SEPT 25, 1977. *APPROXIMATIONS USING DERIVATIVE
C      FORMS OF GENERALIZED GRAM-CHARLIS
C      UPDATED VERSION PAGES 6,7
C      USED WITH ITEM4, AT. TURBULENCE Part 5
C      REQUIRES SUBROUTINE GAM
C

```

```

      PI = 3.141592654
      PI2 = PI/2.
      TPI = (2.*PI)**-.5

```

```

C
C
7001  TYPE 7001
      FORMAT(' INPUT ALPHF :',/s)
      ACCEPT 702,ALPH1F
      ACCEPT 702,ALPH2F
      ACCEPT 702,ALPH3F
      ACCEPT 702,ALPH4F
      ACCEPT 702,ALPH5F
      ACCEPT 702,ALPH6F
702   FORMAT(G)
C

```

```

      ALPB = ALPH2F = ALPH1F*ALPH1F
      ALPB2 = SQRT(ALPB)
      GAMMA = ALPH1F*ALPH1F/ALPB
1010  TYPE 1010,GAMMA
      FOMAT(' GAMMA =',G)
      ALAM = GAMMA/ALPH1F
      GAM2 = GAMMA/2.
      GAM12 = (GAMMA+1.)/2.
      TGAM = 2.**(GAMMA-.5)
      DELX = .1*ALPB2

```

```

C
C
      CALL GAM(GAM2,GG1)
      CALL GAM(GAM12,GG2)
C

```

```

      BE1 = (GAMMA-1.)/2.
      BETA = GAMMA - 1.
      B2 = BETA*BETA
      B3 = BETA*B2
      B4 = BETA*B3
      AL3 = ALAM*ALAM*ALAM
      AL4 = AL3*ALAM
      G12 = GAMMA*(GAMMA+1.)*(GAMMA+2.)/6.
      G123 = G12*(GAMMA+3.)/4.
      G4 = G123*(GAMMA+4.)/5.
      G5 = G4*(GAMMA+5.)/6.
      CO1 = (ALAM/GAMMA)
      CO2 = (CO1*ALAM/(GAMMA+1.))

```

```

CO3 = (CO2*ALAM/(GAMMA+2.))
CO4 = (CO3*ALAM/(GAMMA+3.))
CO5 = (CO4*ALAM/(GAMMA+4.))
CO6 = (CO5*ALAM/(GAMMA+5.))
BB3 = G12*(1.-3.*CO1*ALPH1F+3.*CO2*ALPH2F-CO3*ALPH3F)
BB4 = G123*(1.-4.*CO1*ALPH1F+6.*CO2*ALPH2F-4.*CO3*ALPH3F+
1      CO4*ALPH4F)
BB5 = G4*(1.-5.*CO1*ALPH1F+10.*CO2*ALPH2F-10.*CO3*ALPH3F+5.*
1      CO4*ALPH4F-CO5*ALPH5F)
BB6 = G5*(1.-6.*CO1*ALPH1F+15.*CO2*ALPH2F-20.*CO3*ALPH3F+15.*
1      CO4*ALPH4F-6.*CO5*ALPH5F+CO6*ALPH6F)

```

C

```

F0N0 = 0.
F0N1 = 0.
F0N2 = 0.
F1N0 = 0.
F1N1 = 0.
F1N2 = 0.
F1N3 = 0.
F2N0 = 0.
F2N1 = 0.
F2N2 = 0.
F2N3 = 0.
F2N4 = 0.
F3N0 = 0.
F3N1 = 0.
F3N2 = 0.
F3N3 = 0.
F3N4 = 0.
F3N5 = 0.
F4N0 = 0.
F4N1 = 0.
F4N2 = 0.
F4N3 = 0.
F4N4 = 0.
F4N5 = 0.
F4N6 = 0.

```

C

600

C

```

TYPE 600
FORMAT(/, '      X', 11X, 'F1PRM', 9X, 'F2PRM', 9X, 'F3PRM', 9X, 'F4PRM')

```

C

```

DO 20 I=1,100
X = DELX*I
ALX = ALAM*X
PHI0 = (ALAM*EXP(-ALX))*((ALX**BE1)/GG1)*((ALX**BE1)/GG2))/
1      (TPI*TGAM)

P3 = 1. - 3.*CO1*X + 3.*CO2*X*X - CO3*X*X*X
P4 = 1. - 4.*CO1*X + 6.*CO2*X*X - 4.*CO3*X*X*X + CO4*X**4
P5 = 1.-5.*CO1*X+10.*CO2*X*X-10.*CO3*X**3+5.*CO4*X**4-CO5*X**5
P6 = 1.-6.*CO1*X+15.*CO2*X*X-20.*CO3*X**3+15.*CO4*X**4-
1      6.*CO5*X**5+CO6*X**6
F1PRM = PHI0 + BB3*PHI0*P3
F2PRM = F1PRM + BB4*PHI0*P4
F3PRM = F2PRM + BB5*PHI0*P5
F4PRM = F3PRM + BB6*PHI0*P6

```

C

```

F0N0 = F0N0 + PHI0
F0N1 = F0N1 + PHI0*X
F0N2 = F0N2 + PHI0*X*X
F1N0 = F1N0 + F1PRM
F1N1 = F1N1 + F1PRM*X
F1N2 = F1N2 + F1PRM*X*X
F1N3 = F1N3 + F1PRM*X**3
F2N0 = F2N0 + F2PRM
F2N1 = F2N1 + F2PRM*X
F2N2 = F2N2 + F2PRM*X*X
F2N3 = F2N3 + F2PRM*X**3
F2N4 = F2N4 + F2PRM*X**4
F3N0 = F3N0 + F3PRM
F3N1 = F3N1 + F3PRM*X
F3N2 = F3N2 + F3PRM*X*X
F3N3 = F3N3 + F3PRM*X**3
F3N4 = F3N4 + F3PRM*X**4
F3N5 = F3N5 + F3PRM*X**5
F4N0 = F4N0 + F4PRM
F4N1 = F4N1 + F4PRM*X
F4N2 = F4N2 + F4PRM*X*X
F4N3 = F4N3 + F4PRM*X**3
F4N4 = F4N4 + F4PRM*X**4
F4N5 = F4N5 + F4PRM*X**5
F4N6 = F4N6 + F4PRM*X**6

```

TYPE 700,X,F1PRM,F2PRM,F3PRM,F4PRM

20

CONTINUE

700

FORMAT(F6,2,2X,4G)

C

TYPE 7004,F0N0,F0N1,F0N2

TYPE 7004,F1N0,F1N1,F1N2,F1N3

TYPE 7004,F2N0,F2N1,F2N2,F2N3,F2N4

7004

FORMAT(5G)

TYPE 7004,F3N0,F3N1,F3N2,F3N3,F3N4

TYPE 7004,F3N5

TYPE 7004,F4N0,F4N1,F4N2,F4N3,F4N4

TYPE 7004,F4N5,F4N6

C

C

END

Subroutine HPDES

HPDES
 HIPASS BUTTERWORTH FILTER DESIGN SUBROUTINE
 INPUTS ARE CUTOFF (3-DB) FREQUENCY FC IN HERTZ,
 SAMPLING INTERVAL T IN SECONDS, AND
 NUMBERS OF FILTER SECTIONS.
 OUTPUTS ARE NS SETS OF FILTERED COEFFICIENTS, I.E.,
 A(K) THRU C(K) FOR K=1 THRU NS, AND
 10 PAIRS OF FREQUENCY AND POWER GAIN, I.E.,
 GR(1,K) AND GR(2,K) FOR K=1 THRU 10
 NOTE THAT A(K),B(K),C(K) AND GR(2,10) MUST BE DIMENSIONED
 IN CALLING PROG.

THE DIGITAL FILTER HAS NS SECTIONS IN CASCADE. THE KTH
 SECTION HAS THE TRANSFER FUNCTION

$$H(Z) = \frac{A(K) * Z(Z^{*2} - 2 * Z + 1)}{Z^{*2} + B(K) * Z + C(K)}$$

THUS IF F(M) AND G(M) ARE THE INPUT AND OUTPUT OF THE
 KTH SECTION AT TIME M*T, THEN

$$G(M) = A(K) * (F(M) - 2 * F(M-1) + F(M-2)) - B(K) * G(M-1) - C(K) * G(M-2)$$

SUBROUTINE HPDES(FC,T,NS)
 COMMON/CC/A(3),B(3),C(3),GR(2,10),F(4,3)
 PI = 3.1415926536
 WCP = SIN(FC*PI*T)/COS(FC*PI*T)
 DO 120 K=1,NS
 CS = COS(FLOAT(2*(K+NS)-1)*PI/FLOAT(4*NS))
 A(K) = 1./((1.+WCP*WCP-2.*WCP*CS)
 B(K) = 2.*((WCP*WCP-1.)*A(K)
 120 C(K) = (1.+WCP*WCP+2.*WCP*CS)*A(K)
 DO 130 K=1,10
 GR(2,K) = .01+.98*FLOAT(K-1)/9.
 X = ATAN(WCP*(1./GR(2,K)-1.))*(-1./FLOAT(4*NS))
 130 GR(1,K) = X/(PI*T)
 RETURN
 END

Program ITEM3

PROGRAM ITEM3. ATMOSPHERIC TURBULENCE TASK
ITEM 3, PARTS 5 AND 6. PHASE II OF ATMOSPHERIC TURB.

PROGRAM READS DATA FILE :
AUTOF= AUTOCORRELATION FN FROM ATURB4
AUTF2= AUTOCORRELATION PRODUCES WITH W**2= FROM
MODIFIED VERSION OF ATURB4
PROGRAM PRODUCES DATA FILES:
RSIGF= VALUES OF R (SIGMA SQRD F), PART 5 ITEM 3
PHIF= VALUES OF PHI OF F, PART 6, ITEM 3

REQUIRES SUBROUTINE CFFT1

RUN PROGRAM ATURB4 TO DETERMINE PHI L, R L, AND PHI L ,
R L FOR SQUARED W(T) VALUES (SEE MARK'S NOTES ITEM3)
RUN ITEM2 TO DETERMINE SIGMA SQRD AND L

DIMENSION RWH(0/1300),RWH2(0/1300)
DIMENSION AUT(0/5000)

INPUT INITIAL PARAMETERS

TYPE 1 FC (filter cut-off frequency)
FORMAT(/1X,'INPUT DELX',/8)
ACCEPT 1750,DELX ,
FORMAT(3G)

READ IN AUTOCORRELATION FN.

CALL IFILE (20,'AUTOF')

READ(20,900)

READ(20,1950)FLE

FORMAT(/1X,'AUTOCORRELATION OF',/1X,
1 'HIGH PASS FILTERED NON-HOMOGENEOUS SAMPLE',/1X,

2 'DATA TAKEN FROM FILE ',A5)

READ(20,906)NPTS,TWOL,MPTS,FTM1

READ(20,907)MZERO

READ(20,910)WBAR,VAR

READ(20,951)WSUM

FORMAT(/1X,'<W OF L(X)**2> = ',F12.4)

READ(20,381)XAOFF,MAOFF

FORMAT(/1X,' TRUCATION POINT WAS',G,' METERS',/1X,

1 'WHICH CONTAINS ',I7,' POINTS')

READ(20,1952)

FORMAT(/1X,'PRINTOUT OF THE VALUES OF THE AUTOCORRELATION',

1 /1X,8X,'X',12X,' RL ',9X,'RL/R0')

DO 1823 I=0,MAOFF

1823 READ(20,1915)DEL,RWH(I),YY

1915 FORMAT(1X,3(2X,G))

END FILE 20

TYPE 4324,RWH(MAOFF)

```

4324      FORMAT( ' RWH(MAOFF) =',G)
C
C      READ IN AUTOCORRELATION FN. OF SQUARED W(T)
C
      CALL IFILE (20,'AUTF2')
      READ(20,900)
1965      FORMAT(/1X,'AUTOCORRELATION OF',/1X,
1      'HIGH PASS FILTERED AND SQUARED NON-HOMOGENEOUS SAMPLE',/1X,
2      'DATA TAKEN FROM FILE ',A5)
      READ(20,1965)FLE
      READ(20,906)NPTS,TWOL,MPTS,FTM1
      READ(20,907)MZERO
      READ(20,910)WBAR,VAR
      READ(20,951)WSUM
      READ(20,381)XAOFF,MAOFF
      READ(20,1952)
      DO 1824 I=0,MAOFF
1824      READ(20,1915)DEL,RWH2(I),YY
      END FILE 20
      TYPE 4325,RWH2(MAOFF)
4325      FORMAT( ' RWH2(MAOFF) =',G)
C
C      READ IN SIGMA SQRD, L (AFTER RUNNING ITEM2)
C
      TYPE 111
111      FORMAT(' INPUT SIGMA SQRD, L',/S)
      ACCEPT 2,SIGF,AL
2      FORMAT(2G)
C
C      DO PART 5, R (SIGMA SQRD F)
C
      CALL OFILE(20,'RSIGF')
      WRITE (20,911)
      WRITE(20,1953)SIGF,AL
1953      FORMAT(' FOR SIGMA SQUARED =',G,' AND AL =',G)
      WRITE(20,1954)
1954      FORMAT(//' XI',7X,'RSIGF')
      SIGF2 = SIGF*SIGF
      DO 500 I=0,MAOFF
      XI = DELX*I
      RSIGF = SIGF2*RWH2(I)/(RWH(0)**2+2.*RWH(I)*RWH(I))
      IF (RSIGF,LT.,.0001) GO TO 501
500      WRITE(20,2)XI,RSIGF
501      END FILE 20
C
C      DO PART 6, PHI(SIGF)
C
      TYPE 3
3      FORMAT(' INPUT M AND POWER OF 2 ',/S)
      ACCEPT 1750,MPTS,MPWRM

```

C

```

PI = 3.141592654
RWH02 = RWH(0)*RWH(0)
FTM1 = MPTS*DELX
DO 13 J=0,MPTS
  ARG = DELX*J
  CC = PI*ARG/FTM1
  DD = 1.-ARG/FTM1
  EE = ABS(SIN(CC))
  WINDO = EE/PI+DD*COS(CC)
  RTP = RWH02+2.*RWH(J)*RWH(J)
13  AUT(2*J) = ((RWH2(J)-RTP)/RTP)*WINDO
  M2PWR = MPWRM+1
  M2PTS = 2*MPTS
  FT2M = 2.*FTM1
  MMIN1 = MPTS-1
  DO 14 JK=1,MMIN1
    KK = M2PTS+JK
14  AUT(2*KK) = AUT(2*JK)
  DO 521 K=1,2*M2PTS-1,2
521  AUT(K) = 0.0
C
C  COMPUTE SMOOTHED POWER SPECTRUM
C
C  CALL CFFT(M2PWR,M2PTS,AUT,2)
C
C  PRINT DATA FILE 'PHIF'
C
C  CALL OFILE(20,'PHIF')
  WRITE(20,911)
  WRITE(20,901) FC
901  FORMAT(/1X,'SMOOTHED POWER SPECTRUM PHI OF SIGMA SQRD F(K)')
  WRITE(20,1906)MPTS,FTM1
1906  FORMAT(/1X,I5,' DATA POINTS WERE USED IN M = ',F16.4,' FT. ')
  WRITE(20,1905)RWH(0)
1905  FORMAT(' RWH(0) = ',G)
  ADD TO FORMAT STATEMENT 901
  WRITE(20,902) [(X,' WITH FILTER CUT-OFF=',
  G,'M**-1')]
  MHALF = MPTS/2
  M41 = MHALF-1
  DELK = 1./FT2M
  DO 56 I=0,M41
    DEL = DELK*I
    XX = AUT(2*I)*FT2M*SIGF*SIGF
    K = MHALF+I
    DEL2 = DELK*K
    YY = AUT(2*K)*FT2M*SIGF*SIGF
56  WRITE(20,903)DEL,XX,DEL2,YY
  DEL3 = DELK*MPTS
  ZZ = AUT(2*MPTS)*FT2M*SIGF*SIGF
  WRITE(20,908)DEL3,ZZ
  END FILE 20

```

```

C
902  FORMAT(//1X,'PRINTOUT OF VALUES OF THE SMOOTHED',
      1  ' POWER SPECTRUM',/5X,'K',10X,'SPS VALUE',8X,'K CONTD',
      2  4X,'SPS VALUE CONTD',/)
908  FORMAT(29X,F10.4,3X,E12.4)
900  FORMAT(//1X,'DATA FILE CREATED BY PROGRAM ATURB4')
903  FORMAT(1X,F10.6,3X,E12.4,3X,F10.6,3X,E12.4)
906  FORMAT(//1X,I6,' DATA POINTS WERE USED IN 2L = ',F15.4,
      1  ' METER',/1X,I5,' DATA POINTS WERE USED IN M = ',F16.4,
      1  ' METER')
907  FORMAT(//1X,I6,' ZEROS WERE ADDED TO DATA')
910  FORMAT(//1X,'MEAN VALUE OF W(X) = ',E15.5,' M/SEC',/1X,
      1  'MEAN SQ. VALUE = ',E15.5,' (M/SEC)**2')
911  FORMAT(' DATA FILE CREATED BY PROGRAM ITEM3')
C
END

```

Program ITEM4


```

C      PROGRAM ITEM4 ,PART 6   ATMOS. TURBULENCE PHASE II
C
      DIMENSION ALWF(8),ALWS(0/8),ALPHWH(8),PROB(0/100)
      DIMENSION ALPHW(8)
C
      TYPE 1
1      FORMAT(' INPUT ALPHWH,ALPHW (8 VALUES) ',/8)
      ACCEPT 2,ALPHWH(1),ALPHW(1)
      ACCEPT 2,ALPHWH(2),ALPHW(2)
      ACCEPT 2,ALPHWH(3),ALPHW(3)
      ACCEPT 2,ALPHWH(4),ALPHW(4)
      ACCEPT 2,ALPHWH(5),ALPHW(5)
      ACCEPT 2,ALPHWH(6),ALPHW(6)
      ACCEPT 2,ALPHWH(7),ALPHW(7)
      ACCEPT 2,ALPHWH(8),ALPHW(8)
C
2      FORMAT(2G)
      TYPE 3
3      FORMAT(' INPUT SIGMA SQRD F ',8)
      ACCEPT 2,SIG
C
      AK = (SIG/ALPHWH(2))**.5
      DO 100 K = 1,8
100     ALWF(K) = (AK**K)*ALPHWH(K)
      ALWS(0) = 1.0
      DO 110 N = 1,8
      SUM = 0.0
      DO 120 K=0,N-1
      CALL FAC1(N,K,ANK)
120     SUM = SUM + ANK*ALWF(N-K)*ALWS(K)
110     ALWS(N) = ALPHW(N) * SUM
C
C      COMPUTE PROB DENSITY FN
C
      TYPE 8
8      FORMAT(' PROB DENSITY FN',8)
      DELWS = .1*(SQRT(ALPHWH(2)-ALPHWH(1)*ALPHWH(1)))
      SIGWS = ALWS(2)**.5
      TWOSG = 2.*SIGWS*SIGWS
      C1 = 1./((2.*3.14159265)**.5*SIGWS)
      SIG3 = SIGWS**3
      SIG4 = SIGWS**4
      SIG2 = SIGWS**2
C
      DO 130 I=0,100
      WS = DELWS*I
      WS2 = WS*WS
      WS3 = WS2*WS
      WS4 = WS3*WS
      EX = -(WS2/TWOSG)

```

```

130  PROB(I) = C1*EXP(EX)*(((1.+((ALWS(3)*(WS3/SIG3=
1      3.*WS/SIGWS)/(6.*SIG3))+(((ALWS(4)/SIG4)-3.)*
2      (WS4/SIG4-6.*WS2/SIG2+3.))/24.))
      GAM1 = ALWS(3)/(ALWS(2)**(3./2.))
      GAM2 = (ALWS(4)/(ALWS(2)**2))-3.

C
C
C  PRINT OUT VALUES OF PROBABILITYE

      CALL OFILE(20,'PROB')
      WRITE(20,900)
900  FORMAT(' DATA FILE CREATED BY PROGRAM ITEM4',/1X,
1      ' ATMOSPHERIC TURBULENCE TASK, PHASE II')
      WRITE(20,901)SIG
901  FORMAT(' FOR SIGMA SQRD F =',G,/1X,'THE VALUES OF ALPHA ',
1      'OF WS ARE:')
      DO 170 N=1,8
170  WRITE(20,902)N,ALWS(N)
902  FORMAT(' N =',I3,'ALPHA WS =',G)
      WRITE(20,903)
903  FORMAT(/1X,'THE PROBABILITY DENSITY FN. IS:')
1      /7X,'WS',13X,'P(WS)')
      DO 180 I=0,100
      WS = DELWS*I
180  WRITE(20,904)WS,PROB(I)
904  FORMAT(1X,G,2X,G)
      WRITE(20,905)GAM1,GAM2
905  FORMAT(/1X,'COEFFICIENT OF SKEWNESS =',G,/1X,
1      'COEFFICIENT OF EXCESS =',G)

```

Program MOMENT

```

C      PROGRAM MOMENT   ATMOSPHERIC TURBULENCE PHASE II
C      ITEM 4   Parts 1-4
C
C      CALL SUBROUTINES BIN(5Q),HPDES
C
C      COMMON/AA/BIN1(501),BIN2(501),TOTAL,NBIN,BINW,NPTS
C      COMMON/BB/BIN3(700),TOT1,NBIN1,BINW1,NPTS1
C      COMMON/CC/A(3),B(3),C(3),GR(2,10),F(4,3)
C
C      DIMENSION D(8),ALSIG(8)
C      DIMENSION ALWS(0/8)
C      DIMENSION W(0/15000),ALPHW(8),ALPHWH(8),ALPWH2(8)
C
C      READ IN W(X) DATA
C
C      TYPE 1
1      FORMAT(' INPUT NO. OF DATA POINTS',8)
C      ACCEPT 2,NPTS
2      FORMAT(3G)
C      TYPE 3
3      FORMAT(' INPUT DATA RECORD NAME ',8)
C      ACCEPT 4,FLE
4      FORMAT(A5)
C
C      READ IN FILE AND CONVERT TO METERS
C
C      CALL IFILE(20,FLE)
C      NX1 = NPTS/4
C      DO 550 I=0,NX1-1
C      READ(20,551)W(I),W(I+NX1),W(I+2*NX1),W(I+3*NX1)
C      W(I+NX1) = W(I+NX1)*.3048
C      W(I+2*NX1) = W(I+2*NX1)*.3048
C      W(I+3*NX1) = W(I+3*NX1)*.3048
550    W(I) = W(I)*.3048
551    FORMAT(4(E15.7))
C      END FILE 20
C
C      SUBTRACT OUT MEAN
C
C      WBAR = 0.0
C      DO 610 JJ = 0,NPTS-1
610    WBAR = WBAR+W(JJ)
C      WBAR = WBAR/FLOAT(NPTS)
C      DO 611 I=0,NPTS-1
611    W(I) = W(I)-WBAR
C
C      COMPUTE BINS
C
C      CALL BIN(W)
C
C      COMPUTE MOMENTS OF W(T)

```

```

BINHF = BINW/2.
NBIN2 = NBIN/2

```

C

```

TYPE 900
FORMAT(' MOMENTS OF W(X) ')
DO 10 K=1,8
SUM = 0.0
DO 20 J=1,NBIN2
WMID1 = (BINW*J-BINHF)**K
WMID2 = (BINW*(-J)+BINHF)**K
20 SUM = SUM + WMID1*BIN1(J) + WMID2*BIN2(J)
ALPHW(K) = SUM/TOTAL
10 TYPE 901,K,ALPHW(K)
901 FORMAT(' K =',I3,' ALPHW = ',G)

```

C

C

C

C

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C

C

C

```

TYPE 9
FORMAT(' INPUT CUT-OFF FREQ,SAMPLING INTERVAL, '
1 'NO. OF FILTER SECTIONS',/S)
ACCEPT 2,FC,TS,NS
CALL HPDES(FC,TS,NS)
DO 140 N=1,NS+1
DO 140 M=1,2
140 F(N,M) = 0.0
DO 150 M=0,NPTS-1
F(1,3) = W(M)
DO 160 N=1,NS
TEMP = A(N)*(F(N,3)-2.*F(N,2)+F(N,1))
160 F(N+1,3) = TEMP+B(N)*F(N+1,2)-C(N)*F(N+1,1)
DO 170 N=1,NS+1
DO 170 MM=1,2
170 F(N,MM) = F(N,MM+1)
150 W(M) = F(NS+1,3)

```

MOMENTS OF WH

```

TYPE 5
FORMAT(' MOMENTS OF FILTERED W')

```

```

CALL BIN(W)
BINHF = BINW/2.
NBIN2 = NBIN/2

```

C

```

DO 30 K=1,8
SUM = 0.0
DO 40 J = 1,NBIN2
WMID1 = (BINW*J-BINHF)**K
WMID2 = (BINW*(-J)+BINHF)**K

```

```

40      SUM = SUM + WMID1*BIN1(J) + WMID2*BIN2(J)
      ALPHWH(K) = SUM/TOTAL
30      TYPE 902,K,ALPHWH(K)
902     FORMAT(' K =',I3,' ALPHWH =',G)
C
C      COMPUTE WH**2, PART 3
C
C      DO 50 I=0,NPTS-1
50      W(I) = W(I)*W(I)
C
C      COMPUTE BINS
C
C      TYPE 6
6      FORMAT(' COMPUTE MOMENTS FOR W*W')
C
C      NPTS1 = NPTS
      CALL BINSQ(W)
C
C      BINHF = BINW1/2.
C
C      DO 60 K=1,8
      SUM = 0.0
      DO 70 J=1,NBIN1
70      WMID = (BINW1*J-BINHF)**K
      SUM = SUM + WMID*BIN3(J)
      ALPHW2(K) = SUM/TOT1
60      TYPE 903,K,ALPHW2(K)
903     FORMAT(' K =',I3,' ALPHW2 =',G)
C
C      PART4, MOMENTS OF SIGMA SQRD F
C
C      D(1) = 1.
      D(2) = 3.
      D(3) = 15.
      D(4) = 105.
      D(5) = 945.
      D(6) = 10395.
      D(7) = 135135.
      D(8) = 2027025.
      TYPE 905
905     FORMAT(' COMPUTE MOMENTS OF SIGMA SQRD F')
C
C      TYPE 7
7      FORMAT(' INPUT SIG SQRD F ',8)
      ACCEPT 2,SIG
C
C      DO 80 N=1,8
      ALSIG(N) = ((SIG/ALPHW2(1))**N)*ALPHW2(N)/D(N)
80      TYPE 904,N,ALSIG(N)
904     FORMAT(' N =',I3,' ALSIG =',G)
C
      END

```

Subroutine PAR1

```

SUBROUTINE PAR1(AKZ)
C
C CALL SIMQ, FORMS EQN Y**2+DX+EY+F = 0.0
C
COMMON/BB/X(3),AKI(3),XPL
DIMENSION Y(3),A(9),B(3)
EQUIVALENCE(Y,AKI)
C
YMID = AKI(2)
DO 20 J=1,3
B(J) = -Y(J)**2
20 A(J) = X(J)
C
DO 25 K=4,6
25 A(K) = Y(K-3)
A(7) = 1.
A(8) = 1.
A(9) = 1.
C
KK = 0
C
CALL SIMQ(A,B,3,KK)
C
IF (KK.EQ.1) TYPE 518
518 FORMAT(' K = 1 IN SIMQ; SINGULAR SOLUTION = BAD')
C
D = B(1)
E = B(2)
F = B(3)
AKZ = (-2,*XPL-E)/D
C
RETURN
C
END

```


Subroutine PAR2

```

SUBROUTINE PARAB(AKZ)
C
C CALLS SIMQ, FORMS EQN  $Y''+2+DX+EY+F = 0.0$ 
C
COMMON/BB/X(3),AKI(3),XPL
DIMENSION Y(3),A(9),B(3)
EQUIVALENCE(Y,X)
C
IF (AKI(2).EQ.AKI(3)) AKZ = AKI(2)
IF (AKI(2).EQ.AKI(3)) GO TO 40
DO 20 J=1,3
B(J) = -Y(J)**2
20 A(J) = AKI(J)
C
DO 25 K=4,6
25 A(K) = Y(K-3)
A(7) = 1.
A(8) = 1.
A(9) = 1.
C
KK = 0
C
CALL SIMQ(A,B,3,KK)
C
IF (KK.EQ.1) TYPE 518,X(1),AKI(1),XPL
518 FORMAT(' K = 1 IN SIMQ; SINGULAR SOLUTION = BAD',/1X,
1 'X(1)=',G,'AKI(1)=',G,'XPL =',G)
C
D = B(1)
E = B(2)
F = B(3)
AKZ = (-XPL*XPL-E*XPL-F)/D
C
C
C
40 RETURN
END

```

Subroutine PARAB

```

C <RFISHER>PARAB.F4;6      1-NOV-77 10:06:33      EDIT BY RFISHER
      SUBROUTINE PARAB(AKZ)
C
C      CALL SIMQ, FORMS EQN Y**2+DX+EY+F = 0,0
C
      COMMON/BB/X(3),AKI(3),XPL
      DIMENSION Y(3),A(9),B(3)
      EQUIVALENCE(Y,AKI)
C
      YMID = AKI(2)
      DO 20 J=1,3
      B(J) = -Y(J)**2
20      A(J) = X(J)
C
      DO 25 K=4,6
25      A(K) = Y(K-3)
      A(7) = 1.
      A(8) = 1.
      A(9) = 1.
C
      KK = 0
C
      CALL SIMQ(A,B,3,KK)
C
      IF (KK.EQ.1) TYPE 518
518      FORMAT(' K = 1 IN SIMQ; SINGULAR SOLUTION = BAD')
C
      D = B(1)
      E = B(2)
      F = B(3)
      CONST = F + D*XPL
      AKZ1 = (-E + (E**2-4.*CONST)**.5)/2.
      AKZ2 = (-E - (E**2-4.*CONST)**.5)/2.
      DIF1 = ABS(AKZ1-YMID)
      DIF2 = ABS(AKZ2-YMID)
      AKZ = AKZ2
      IF (DIF2.GT.DIF1) AKZ = AKZ1
C
      RETURN
C
      END

```

Program PART2

```

C      PROGRAM PART2.  ATMOSPHERIC TURBULENCE TASK
C      ITEM 1, PARTS 2,3,4,5,6,7, AND 8. PHASE II OF ATMOS-
C      PHERIC TURBULENCE
C      MEAN VALUE SUBTRACTED FROM DATA BEFORE COMPUTING SPECTRUM
C      MAXIMUM LIKELIHOOD ESTIMATION OF INTEGRAL SCALE OF VERTICAL
C      COMPONENT OF STATIONARY SAMPLE
C
C      READS IN PSD FROM FILE PHILK (PRODUCED BY ATURB2.F4)
C      PRODUCES DATA FILES;
C          LG: VALUES OF LG(K;L) PART4, ITEM1
C          PHIXI: VALUES OF PHI OF K(XI)= VON-KARMEN AUTO-
C              CORRELATION FN., PART 7, ITEM 1
C          PHIK: VALUES OF PHI OF K(K)= PART 8, ITEM 1
C
C      REQUIRES SUBROUTINES GAM AND AK
C      SUBR. AK REQUIRES PARAB, AKDAT, SIMQ
C
C      DIMENSION PHIKT(6500),DVSR(4),E(20),ALV(20),ALKI2(6500)
C      DIMENSION W(0/6500)
C
C      INPUT INITIAL PARAMETERS
C
C      TYPE 1
C      1  FORMAT(/1X,'INPUT NO. OF POINTS TO BE READ, DELK,DELX',/8)
C      1750 ACCEPT 1750,NPTS,DELK,DELX
C      1750 FORMAT(3G)
C
C      INPUT VALUES OF PHI OF L(K)
C
C      CALL IFILE (20,'PHILK')
C      READ(20,900)
C      READ(20,950)FLE
C      950  FORMAT(/1X,'POWER SPECTRUM OF PHI OF L(K)',/1X,
C      1  'DATA TAKEN FROM FILE ',A5)
C      READ(20,906)NX2J,TWOL,MPTS,FTM1
C      READ(20,907)MZERO
C      READ(20,910)WBAR,VAR
C      READ(20,951)WSUM
C      951  FORMAT(/1X,'<W OF L(X)**2> = ',F12.4)
C      READ(20,952)
C      952  FORMAT(/1X,'PRINTOUT OF THE VALUES OF THE POWER SPECTRUM',
C      1  /1X,5X,'K',10X,'PS VALUE',8X,'K CONTD',4X,'PS VALUE CONTD',/)
C      DO 823 I=0,NPTS
C      READ(20,903)DEL,W(I),DEL2,YY
C      823  CONTINUE
C      END FILE 20
C
C      START PART2; COMPUTE INTEGRAL SCALE OF L USING EQN FOR E(L)
C
C      1040 TYPE 1500.
C      1500 FORMAT(' INPUT N,L, & STEP SIZE OF L ',/8)
C      ACCEPT 1750,N,ALV(1),STPL

```

```

C      AN = FLOAT(N)
      C116 = 11./6.
      DVSR(1) = 1.
      DVSR(2) = 1.
      DVSR(3) = 10.
      DVSR(4) = 10.

C
      NDX = 0

C
1030    NDX = NDX + 1
      E(NDX) = 0.0

C
      SUM1 = 0.0
      DO 1020 JJ=1,N
      ALKI2(JJ) = (ALV(NDX)*DELK*JJ)**2
      PHIKT(JJ) = (1.+188.75*ALKI2(JJ))/(1.+70.78*ALKI2(JJ))*C116
1020    SUM1 = SUM1 + W(JJ)/PHIKT(JJ)
      DO 1010 I=1,N
      GKIL = 117.97*ALKI2(I)*(1.-188.75*ALKI2(I))/
1010    (ALV(NDX)*(1.+70.78*ALKI2(I))*(1.+188.75*ALKI2(I)))
      E(NDX) = E(NDX) + GKIL*((SUM1/AN)-(W(I)/PHIKT(I)))
      E(NDX) = E(NDX)/AN
      TYPE 1525,NDX,E(NDX)
1525    FORMAT(' ON ',I3,' PASS E =',G)

C
C      PICK NEW L FOR SECOND AND 4TH PASS
C
      IF (E(NDX).LT.0.0) ALV(NDX+1) = ALV(NDX) + STPL/DVSR(NDX)
      IF (E(NDX).GT.0.0) ALV(NDX+1) = ALV(NDX) - STPL/DVSR(NDX)

C
C      USE LINEAR INTERPOLATION FOR NDX = 2 OR 4
C
      IF (NDX.EQ.1) GO TO 1030
      IF (NDX.EQ.3) GO TO 1030
      TYPE 1503
1503    FORMAT(' ARE THERE POS. AND NEG. E'S (Y OR N) ',8)
      ACCEPT 807,ANS
807     FORMAT(A5)
      IF (ANS.EQ.'N') GO TO 1040

C
C      DO LINEAR INTERPOLATION
C
      ALV(NDX+1) = (ALV(NDX)*(-E(NDX-1))+ALV(NDX-1)*E(NDX))/
1      (E(NDX)-E(NDX-1))
      TYPE 1502,NDX,ALV(NDX+1)
1502    FORMAT(' FOR ',I2,' PASS INTERPOLATED L =',G)
      IF (NDX.EQ.2) GO TO 1030

C
C      FINAL L HAS BEEN SELECTED AFTER 4 PASSES
C
      ALF = ALV(NDX+1)

```

C
C
C
PART 4 , COMPUTE LG(KI,L)

CALL OFILE(20,'LG')

WRITE(20,1507)ALF

1507 FORMAT(' OUTPUT OF ITEM 1, PART 4, PHASE II ATMO. TURB.',

1 /1X,'CREATED BY PROGRAM PART2',/1X,

2 ' FOR L =',G,' METERS',/1X,' K',7X,'LG',7X,'K CONTD',5X,

3 'LG CONTD')

NH2 = N/2

DO 1050 I1 = 1,NH2

AKZ = DELK*I1

AKIL = ALF*AKZ

AK1 = DELK*(NH2+I1)

AKIL3 = ALF*AK1

AKIL2 = AKIL*AKIL

AKIL4 = AKIL3*AKIL3

ALG = (117.97*AKIL2*(1.+188.75*AKIL2))/

1 ((1.+70.78*AKIL2)*(1.+188.75*AKIL2))

ALG2 = (117.97*AKIL4*(1.+188.75*AKIL4))/

1 ((1.+70.78*AKIL4)*(1.+188.75*AKIL4))

1050 WRITE(20,1505)AKZ,ALG,AK1,ALG2

1505 FORMAT(4G)

END FILE 20

C
C
C
TYPE 1506

1506 FORMAT(' CONTINUE OR ABORT? (C OR A) ',S)

ACCEPT 807,GOON

IF (GOON.EQ.'A') GO TO 9999

C
C
C
DO PART 5

SIG2 = 0.0

DO 1060 I=1,N

1060 SIG2 = SIG2 + W(I)/(ALF*PHIKT(I))

SIG2 = SIG2/AN

TYPE 1508,SIG2

1508 FORMAT(' SIG2 =',G)

C
C
C
CONTINUE ?

TYPE 1506

ACCEPT 807,GOON

IF (GOON.EQ.'A') GO TO 9999

C
C
C
PART 7, REQUIRES SUBROUTINES GAM AND AK

CALL OFILE(21,'PHIXI')

WRITE (20,915)

WRITE(20,1512)SIG2


```

1512  FORMAT(/1X,'OUTPUT FOR PART 7 OF ITEM I, PHASE II',/1X,
      1  'CALCULATIONS OF VON-KARMEN AUTOCORRELATION FN',/1X,
      2  'SIGMA SQRD =',G)
      WRITE(20,1511)
1511  FORMAT('  XI',5X,'PHI(XI)',5X,'XI CONTD',5X,'PHI(XI) CONTD')
      PI = 3.14159265
      C43 = 4./3.
      CALL GAM(C116,G1)
      CALL GAM(C43,G2)
      GAM13 = 2.67893853
      BETA = 2.*G1*(PI**.5)/(5.*G2)
      ARG = BETA/ALF
      C13 = 1./3.
      C23 = 2./3.
      C13 = 1./3.
      CONST = SIG2*(2.**C23)/GAM13
      DO 1070 I=1,NH2
      E1 = DELX*I
      E2 = DELX*(I+NH2)
      ETA = E1*ARG
      ETA2 = E2*ARG
      IF (ETA.GT'.5'.OR.ETA.LT'.1')TYPE 1515,ETA
      CALL AK(1,ETA,AK1)
      CALL AK(2,ETA,AK2)
      CALL AK(1,ETA2,AK3)
      CALL AK(2,ETA2,AK4)
      PHIK = CONST*(ETA**C13)*(AK1-AK2*ETA/2.)
      PHIK2 = CONST*(ETA2**C13)*(AK3-AK4*ETA2/2.)
1070  WRITE(20,1505) E1,PHIK,E2,PHIK2
1515  FORMAT('  ETA =',G,' WHICH IS OUT OF RANGE FOR AK SUBR.')
      END FILE 20

C
C
C
      PART 8

      CALL OFILE(20,'PHIK')
      WRITE(20,915)
      WRITE(20,1595)ALF,SIG2
1595  FORMAT(/1X,'OUTPUT FOT PART 8 OF ITEM I, PHASE II',/1X,
      1  'CALCULATIONS OF PHI K WITH L =',G,'AND SIG SQRD =',G)
      WRITE(20,1514)
1514  FORMAT('  K',5X,'PHIK',7X,'K CONTD',5X,'PHIK CONTD')
      CONS1 = SIG2*ALF
      STEP = DELK*ALF
      NPTS2 = NPTS/2
      DO 1080 I=1,NPTS2
      AK0 = DELK*I
      AK1 = DELK*(I+NPTS2)
      AK2 = (STEP*I)*(STEP*I)
      AK4 = (STEP*(I+NPTS2))**2
      PHIK = CONS1*(1.+188.75*AK2)/((1.+70.78*AK2)**C116)
      PHIK2 = CONS1*(1.+188.75*AK4)/((1.+70.78*AK4)**C116)

```

```

1080 WRITE(20,1505)AK0,PHIK,AK1,PHIK2
      END FILE 20
C
C
900  FORMAT(//1X,'DATA FILE CREATED BY PROGRAM ATURB2')
915  FORMAT(//1X,'DATA FILE CREATED BY PROGRAM PART2')
903  FORMAT(1X,F10.6,3X,E12.4,3X,F10.6,3X,E12.4)
906  FORMAT(//1X,I6,' DATA POINTS WERE USED IN 2L = ',F15.4,
1    ' METER',/1X,I5,' DATA POINTS WERE USED IN M = ',F16.4,
1    ' METER')
907  FORMAT(//1X,I6,' ZEROS WERE ADDED TO DATA')
908  FORMAT(29X,F10.4,3X,E12.4)
910  FORMAT(//1X,'MEAN VALUE OF W(X) = ',E15.5,' M/SEC',/1X,
1    'MEAN SQ. VALUE = ',E15.5,' (M/SEC)**2')
C
C
9999  END

```

Program PART5

```

C PROGRAM ITEM2, ATMOSPHERIC TURBULENCE TASK
C ITEM 2, PART 5, PHASE II OF ATMOS-
C PHERIC TURBULENCE
C MEAN VALUE SUBTRACTED FROM DATA BEFORE COMPUTING SPECTRUM
C COMPONENT WITH LOW FREQUENCY CONTAMINATION

```

```

C DATA FILES READ BY PROGRAM ARE:
C PHILK: INPUT PSD FROM ATURB2.F4
C

```

```

C DIMENSION W(0/13000),AL(15),SIG2(15)
C

```

```

C INPUT INITIAL PARAMETERS
C

```

```

C TYPE 1001

```

```

1001 FORMAT(/1X,'INPUT NO. OF POINTS TO BE READ, DELK,DELX',/S)
ACCEPT 1750,M41,DELK,DELX

```

```

1750 FORMAT(3G)

```

```

C TYPE 1751

```

```

1751 FORMAT(' IS RECORD TRANSVERSE OR LONGITUDIONAL (T OR L) ',S)
ACCEPT 807,WHICH

```

```

807 FORMAT(A5)
C

```

```

C INPUT VALUES OF PHI OF L(K)
C

```

```

CALL IFILE (20,'PHILK')

```

```

READ(20,900)

```

```

READ(20,950)FLE

```

```

950 FORMAT(/1X,'POWER SPECTRUM OF PHI OF L(K)',/1X,

```

```

1 'DATA TAKEN FROM FILE ',A5)

```

```

READ(20,906)NPTS,TWOL,MPTS,FTM1

```

```

READ(20,907)MZERO

```

```

READ(20,910)WBAR,VAR

```

```

READ(20,951)WSUM

```

```

951 FORMAT(/1X,'<W OF L(X)**2> = ',F12.4)

```

```

READ(20,952)

```

```

952 FORMAT(/1X,'PRINTOUT OF THE VALUES OF THE POWER SPECTRUM',

```

```

1 /1X,5X,'K',10X,'PS VALUE',8X,'K CONTD',4X,'PS VALUE CONTD',/)

```

```

MHALF = NPTS/4

```

```

M51 = MHALF-1

```

```

DO 823 I=0,M51

```

```

K= MHALF+I

```

```

IF (K,GE,M41) K=M41+1

```

```

823 READ(20,903)DEL,W(I),DEL2,W(K)

```

```

END FILE 20
C

```

```

C TYPE 1500

```

```

1500 FORMAT(' INPUT KL,KU,N ',/S)

```

```

ACCEPT 1752,KL,KU,N

```

```

1752     FORMAT(3G)
C
1010     TYPE 1503
1503     FORMAT(' INPUT J,L ',*)
        ACCEPT 1752,J,AL(J)
        JMAX = J
C
C
C
        DO PART 5
        AN = FLOAT(N)
        C116 = 11./6.
        C56 = 5./6.
        SIG2(J) = 0.0
        DO 1060 I=KL,KU
        ALKI2 = (AL(J)*DELK*I)**2
        PHIK = 2./(1.+70.78*ALKI2)**C56
        IF (WHICH.EQ.'T')PHIK=(1.+180.75*ALKI2)/
1          (1.+70.78*ALKI2)**C116
1060     SIG2(J) = SIG2(J) + W(I)/(AL(J)*PHIK)
        SIG2(J) =SIG2(J)/AN
        TYPE 1508,J,SIG2(J)
1508     FORMAT(' FOR J =',I3,' SIG2 =',G)
C
        TYPE 1502
1502     FORMAT(' PICK ANOTHER J,L (Y OR N) ',*)
        ACCEPT 807,AJL
        IF (AJL.EQ.'Y') GO TO 1010
900     FORMAT('//1X,'DATA FILE CREATED BY PROGRAM ATURB3')
903     FORMAT(1X,F10.6,3X,E12.4,3X,F10.6,3X,E12.4)
906     FORMAT('//1X,I6,' DATA POINTS WERE USED IN 2L = ',F15.4,
1       ' METER',/1X,I5,' DATA POINTS WERE USED IN M = ',F16.4,
1       ' METER')
907     FORMAT('//1X,I6,' ZEROS WERE ADDED TO DATA')
910     FORMAT('//1X,'MEAN VALUE OF W(X) = ',E15.5,' M/SEC',/1X,
1       'MEAN SQ. VALUE = ',E15.5,' (M/SEC)**2')
        END

```

Subroutine SIMP2

```

C      SUBROUTINE SIMP, USES SIMPSON'S RULE TO DO INTEGRALS.
C      H = SPACING BETWEEN VALUES
C      NPTS = NO. OF POINTS TO BE COVERED, W=ARRAY CONT-
C      AINING VALUES TO BE INTEGRATED, ANS = INTEGRAL OF W(I)
C
C      SUBROUTINE SIMP(H,NPTS,W,ANS)
C
C      DIMENSION W(0/500)
C
C      ARG = H/3.
C      ANS1 = 0.0
C      DO 10 I=1,NPTS-3,2
10     ANS1 = ANS1 + 4.*W(I)
C      ANS2 = 0.0
C      DO 11 I=2,NPTS-2,2
11     ANS2 = ANS2 + 2.*W(I)
C      ANS = (ANS2+ANS1+W(0)+W(NPTS-1))*ARG
C
C      RETURN
C
C      END

```

Subroutine SIMP


```

C      SUBROUTINE SIMP. USES SIMPSON'S RULE TO DO INTEGRALS.
C      A = STARTING VALUE, B = END VALUE, H = SPACING BETWEEN
C      VALUES, NPTS = NO. OF POINTS TO BE COVERED, W=ARRAY CONT-
C      AINING VALUES TO BE INTEGRATED, ANS = INTEGRAL OF W(I)
C
C      SUBROUTINE SIMP(A,B,H,NPTS,ANS)
C
C      COMMON/SG/W(0/65537)
C
C      ARG = H/3.
C      ANS1 = 0.0
C      DO 10 I=1,NPTS-3,2
10      ANS1 = ANS1 + 4.*W(2*I)
C      ANS2 = 0.0
C      DO 11 I=2,NPTS-2,2
11      ANS2 = ANS2 + 2.*W(2*I)
C      ANS = (ANS2+ANS1+W(0)+W(2*NPTS-2))*ARG
C
C      RETURN
C
C      END

```

Subroutine SIMQ

.....

SUBROUTINE SIMQ

PURPOSE

OBTAIN SOLUTION OF A SET OF SIMULTANEOUS LINEAR EQUATIONS,
AX=B

USAGE

CALL SIMQ(A,B,N,KS)

DESCRIPTION OF PARAMETERS

- A - MATRIX OF COEFFICIENTS STORED COLUMNWISE. THESE ARE DESTROYED IN THE COMPUTATION. THE SIZE OF MATRIX A IS N BY N.
- B - VECTOR OF ORIGINAL CONSTANTS (LENGTH N). THESE ARE REPLACED BY FINAL SOLUTION VALUES, VECTOR X.
- N - NUMBER OF EQUATIONS AND VARIABLES. N MUST BE .GT. ONE.
- KS - OUTPUT DIGIT
 - 0 FOR A NORMAL SOLUTION
 - 1 FOR A SINGULAR SET OF EQUATIONS

REMARKS

MATRIX A MUST BE GENERAL.
IF MATRIX IS SINGULAR , SOLUTION VALUES ARE MEANINGLESS.
AN ALTERNATIVE SOLUTION MAY BE OBTAINED BY USING MATRIX
INVERSION (MINV) AND MATRIX PRODUCT (GMPRD).

SUBROUTINES AND FUNCTION SUBPROGRAMS REQUIRED

NONE

METHOD

METHOD OF SOLUTION IS BY ELIMINATION USING LARGEST PIVOTAL
DIVISOR. EACH STAGE OF ELIMINATION CONSISTS OF INTERCHANGING
ROWS WHEN NECESSARY TO AVOID DIVISION BY ZERO OR SMALL
ELEMENTS.
THE FORWARD SOLUTION TO OBTAIN VARIABLE N IS DONE IN
N STAGES. THE BACK SOLUTION FOR THE OTHER VARIABLES IS
CALCULATED BY SUCCESSIVE SUBSTITUTIONS. FINAL SOLUTION
VALUES ARE DEVELOPED IN VECTOR B, WITH VARIABLE 1 IN B(1),
VARIABLE 2 IN B(2),....., VARIABLE N IN B(N).
IF NO PIVOT CAN BE FOUND EXCEEDING A TOLERANCE OF 0.0,
THE MATRIX IS CONSIDERED SINGULAR AND KS IS SET TO 1. THIS
TOLERANCE CAN BE MODIFIED BY REPLACING THE FIRST STATEMENT.

.....

SUBROUTINE SIMQ(A,B,N,KS)
DIMENSION A(1),B(1)

```

C      FORWARD SOLUTION
C
TOL=0.0
KS=0
JJ=N
DO 65 J=1,N
  JY=J+1
  JJ=JJ+N+1
  BIGA=0
  IT=JJ-J
  DO 30 I=J,N
C
C      SEARCH FOR MAXIMUM COEFFICIENT IN COLUMN
C
  IJ=IT+I
  IF(ABS(BIGA)-ABS(A(IJ))) 20,30,30
20  BIGA=A(IJ)
  IMAX=I
30  CONTINUE
C
C      TEST FOR PIVOT LESS THAN TOLERANCE (SINGULAR MATRIX)
C
  IF(ABS(BIGA)-TOL) 35,35,40
35  KS=1
  RETURN
C
C      INTERCHANGE ROWS IF NECESSARY
C
40  I1=J+N*(J-2)
  IT=IMAX-J
  DO 50 K=J,N
    I1=I1+N
    I2=I1+IT
    SAVE=A(I1)
    A(I1)=A(I2)
    A(I2)=SAVE
C
C      DIVIDE EQUATION BY LEADING COEFFICIENT
C
50  A(I1)=A(I1)/BIGA
  SAVE=B(IMAX)
  B(IMAX)=B(J)
  B(J)=SAVE/BIGA
C
C      ELIMINATE NEXT VARIABLE
C
  IF(J=N) 55,70,55
55  IQS=N*(J-1)
  DO 65 IX=JY,N
    IXJ=IQS+IX
    IT=J-IX
    DO 60 JX=JY,N
      IXJX=N*(JX-1)+IX
      JJX=IXJX+IT
60  A(IXJX)=A(IXJX)-(A(IXJ)*A(JJX))
65  B(IX)=B(IX)-(B(J)*A(IXJ))

```

C
C
C

BACK SOLUTION

```
70 NY=N-1
   IT=N*N
   DO 80 J=1,NY
     IA=IT-J
     IB=N-J
     IC=N
     DO 80 K=1,J
       B(IB)=B(IB)-A(IA)*B(IC)
     IA=IA-N
80  IC=IC-1
   RETURN
END
```

Subroutine TRAP3

```

C <RFISHER>TRAP2.F4;1      1-NOV-77 10:09:04      EDIT BY RFISHER
      SUBROUTINE TRAP2(AI)
C
C      INTEGRALS BY TRAP. RULE
C
      COMMON/TR2/DETX,NXIH,AI2I(0/290)
C
      AI = 0.0
      DO 10 I=1,NXIH-1
10    AI = AI + AI2I(I)
      AI = .5*AI2I(0) + .5*AI2I(NXIH) + AI
      AI = AI*DETX
C
      RETURN
C
      END

```

Subroutine TRAP6


```

C <RFISHER>TRAP5.F4;1      1-NOV-77 10:09:04      EDIT BY RFISHER
      SUBROUTINE TRAP5(AI)
C
C      INTEGRALS BY TRAP. RULE
C
      COMMON/TR5/DELX1,NXIH1,AI5I(0/290)
C
      AI = 0.0
      DO 10 I=1,NXIH1-1
10     AI = AI + AI5I(I)
      AI = .5*AI5I(0) + .5*AI5I(NXIH1) + AI
      AI = AI*DELX1
C
      RETURN
C
      END

```

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